

Development Plan of Phase-Ⅲ TOBA

2026/4/27

東京大学理学系研究科物理学専攻
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杉岡達哉

- Introduction of TOBA
- Summary of my master's work
- Minor update plan for next suspension design
- Toward Phase-III, VIS & Cryogenic

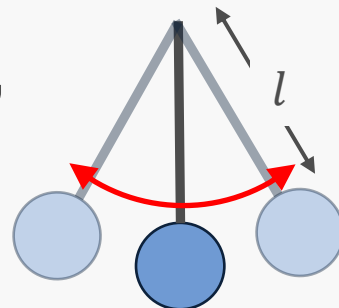
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Low frequency GW

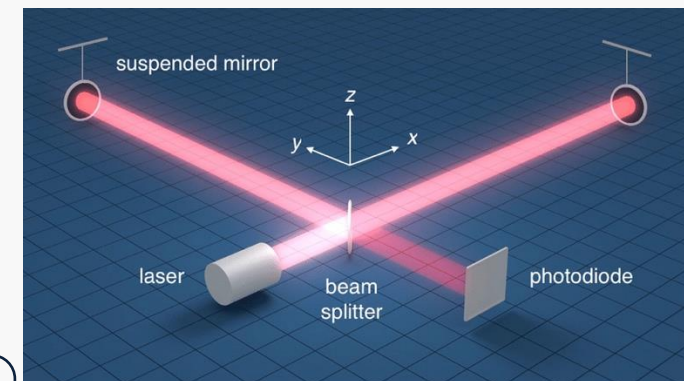
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Current GW detectors, LIGO, Virgo and KAGRA are Michelson interferometer with suspended mirrors

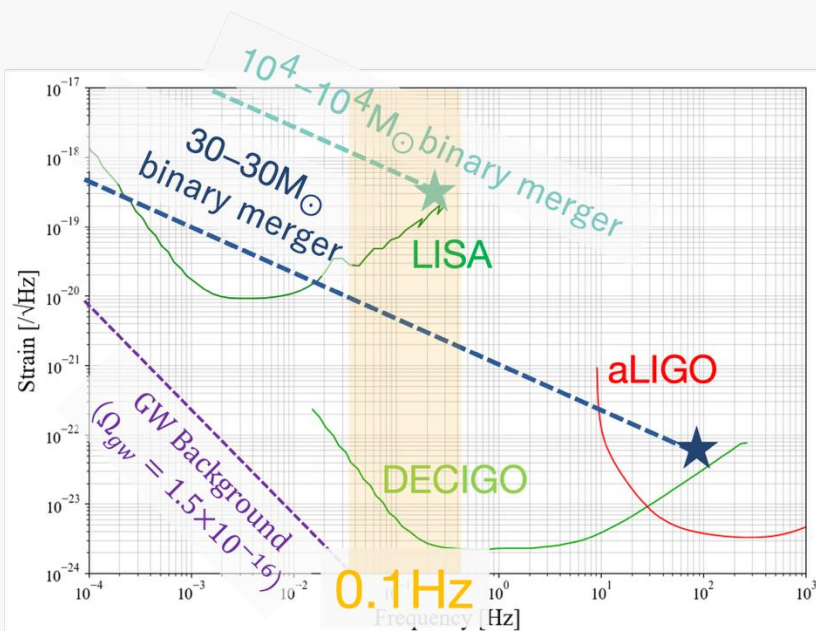
➡ Good sensitivity in 10Hz-1kHz, but not sensitive in lower freq. (below resonant freq. of pendulums)



$$f_0 = \frac{1}{2\pi} \sqrt{\frac{g}{l}}$$



<https://www.ligo.caltech.edu>

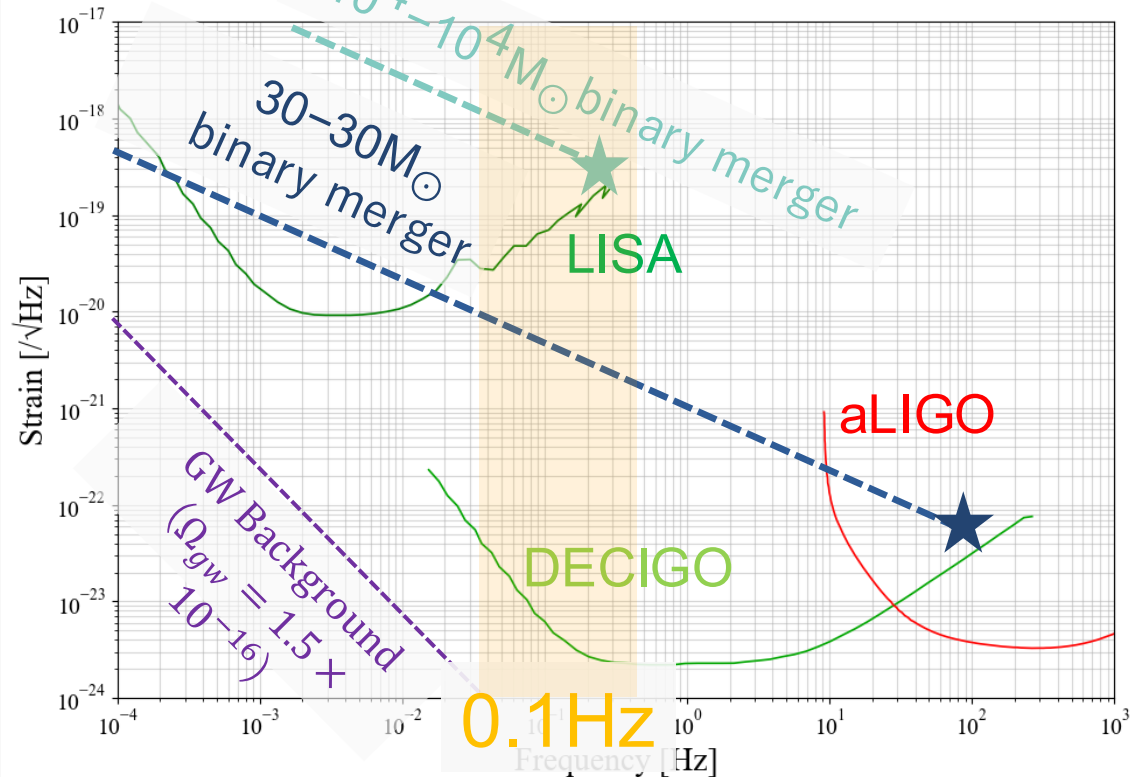


GW source at 0.1Hz:

- Solar Mass BH binary merger
➡ early alert for current detector
- Intermediate Mass BH binary merger
unveiling the formation process of SMBH
- GW background
➡ direct observation of the early universe

Low frequency GW and detectors

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Space born GW detector:
escape from seismic noise and
make TMs in all freq. range

Space GW detector

LISA: 10^{-5} - 10^{-1} Hz

DECIGO: 10^{-1} - 10 Hz

Scheduled for
launch in 2035

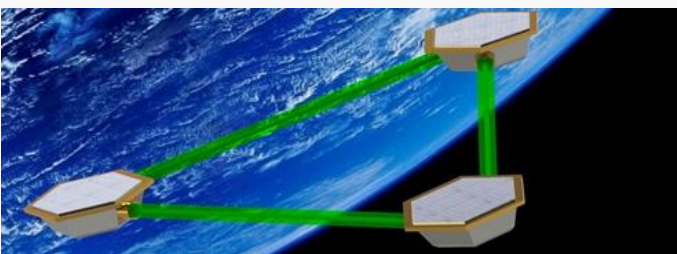
After 2040

High costs and long time...

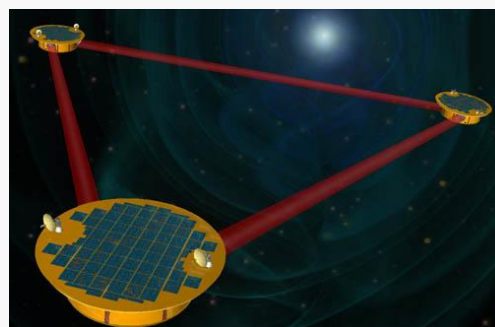
Low cost and easy maintenance

**ground-based detectors
are also important**

DECIGO



LISA



NASA

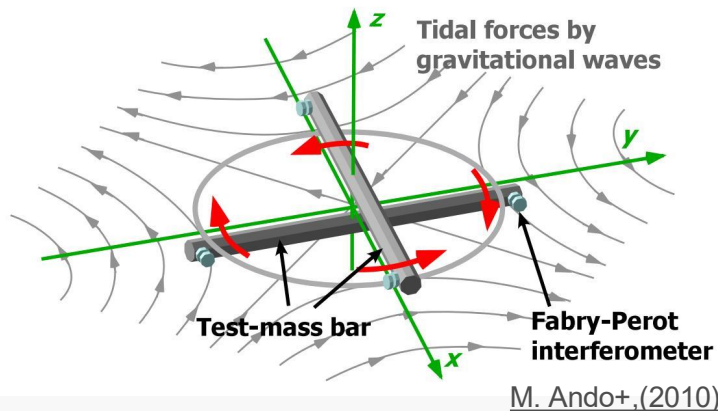
Low freq. Ground-based GW detectors

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Resonant freq. of normal pendulum is above 0.1Hz

↙ **New vibration isolation method** is required
for ground-based detector to 0.1Hz observation

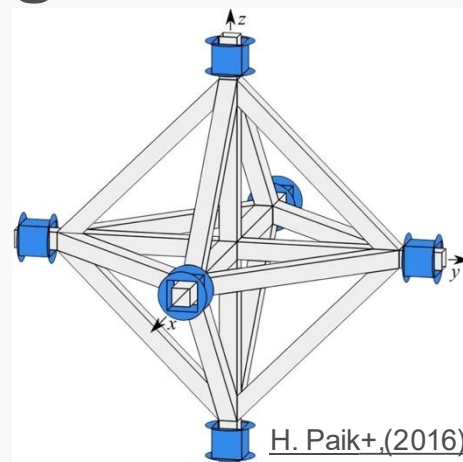
Torsion pendulums



共振周波数の低いねじれ振り子のねじれ回転を計測

- TOBA [M. Ando+, \(2010\)](#)
- TorPeDO [S. Chua+, \(2023\)](#)

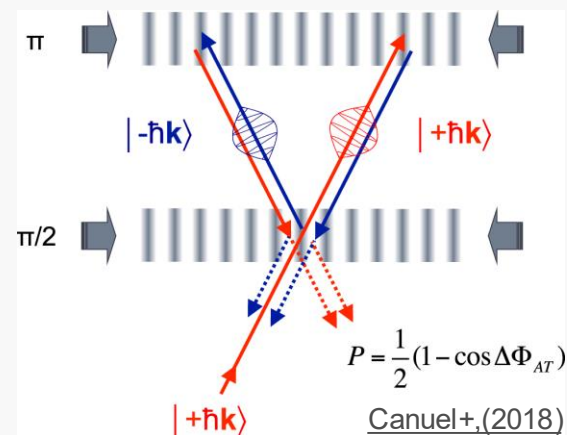
Superconducting magnetic levitation



超伝導磁気浮上したマス間の
相対運動を計測

- SOGRO [H. Paik+, \(2016\)](#)

Atomic interferometer

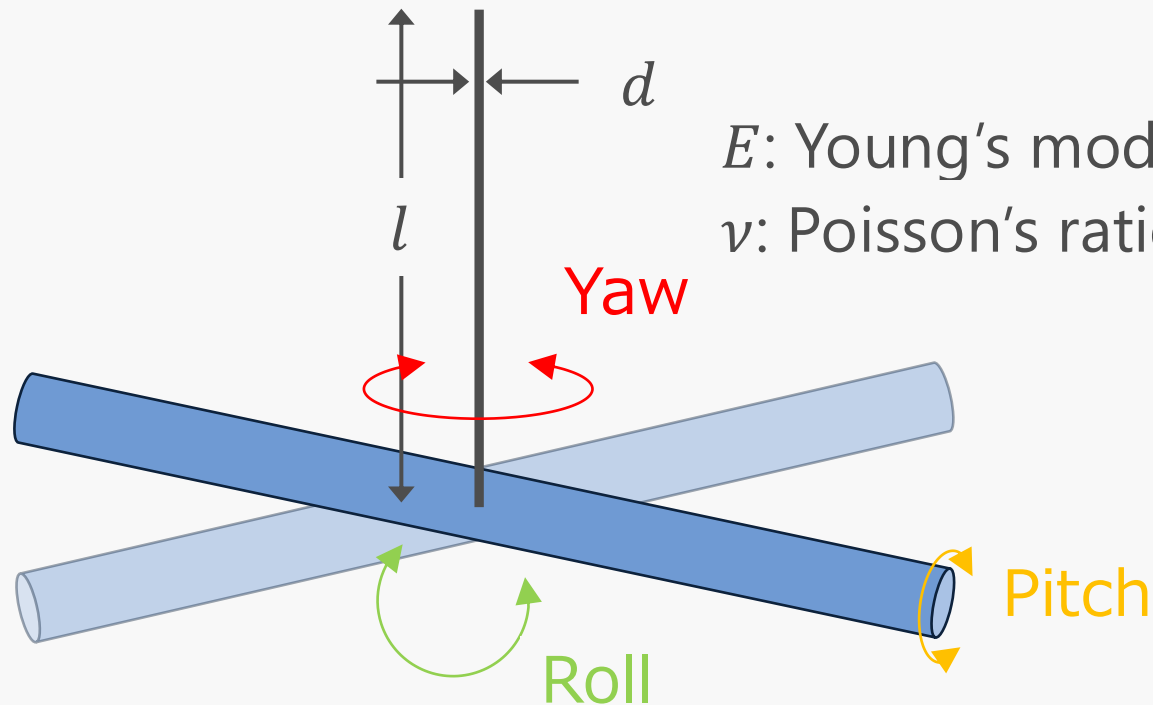


自由落下する原子の物質波の
干渉を計測

- ZAIGA [M. Zhan+, \(2020\)](#)
- MIGA [Canuel+, \(2018\)](#)

Torsion pendulum

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E : Young's modulus
 ν : Poisson's ratio

$$f_{\text{Yaw}} = \frac{1}{2\pi} \sqrt{\frac{\pi E d^4}{64 l (1 + \nu) I}}$$

$$f_{\text{Yaw}} = 20 \text{ mHz}$$

when $l = 50 \text{ cm}$, $d = 0.3 \text{ mm}$

$E = 411 \text{ GPa}$, $\nu = 0.28$,

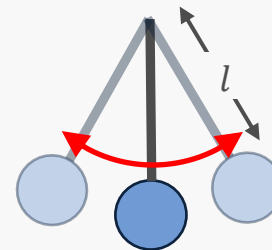
$I = 0.016 \text{ kg} \cdot \text{m}^2$

(tungsten wire and aluminum TM)

I : moment of inertia

Resonant freq. of torsional mode
can be lowered easily!

Normal pendulum



$$f_0 = \frac{1}{2\pi} \sqrt{\frac{g}{l}}$$

$$f_0 = 0.7 \text{ Hz when } l = 50 \text{ cm}$$

$$f_0 = 20 \text{ mHz when } l = 0.6 \text{ km}$$

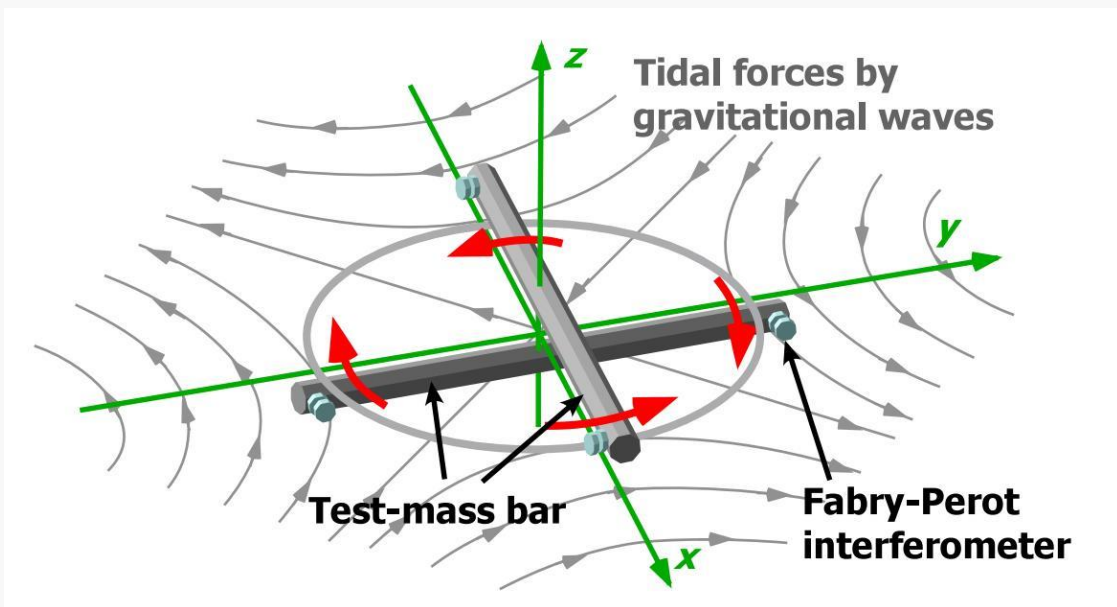
Response of torsion pendulum

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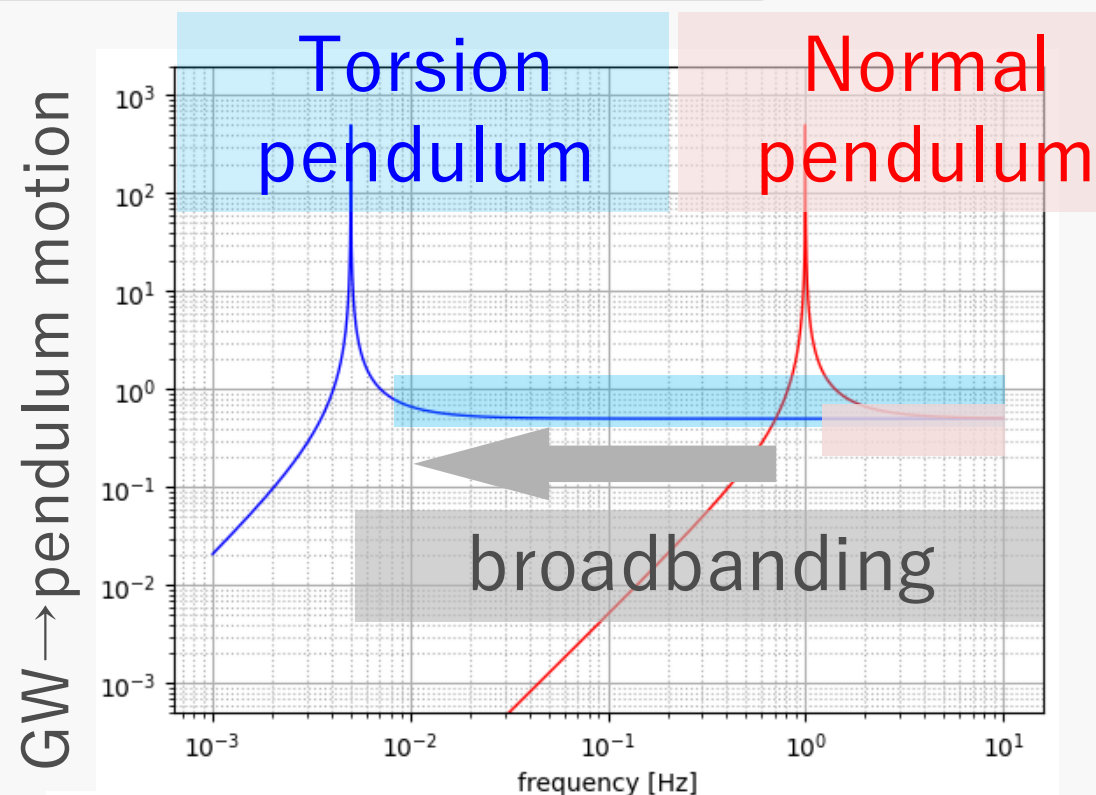
Test Masses have to be free masses to detect GW

Used normal pendulums, but
broadbanding toward low freq. is limited (resonant freq. $\propto \sqrt{1/l}$)

Resonant freq. of torsion pendulums: ~mHz

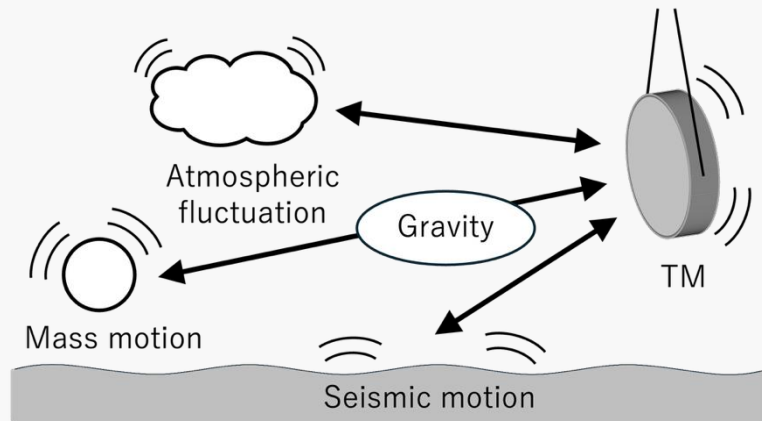


M. Ando+, (2010)



Geophysical science

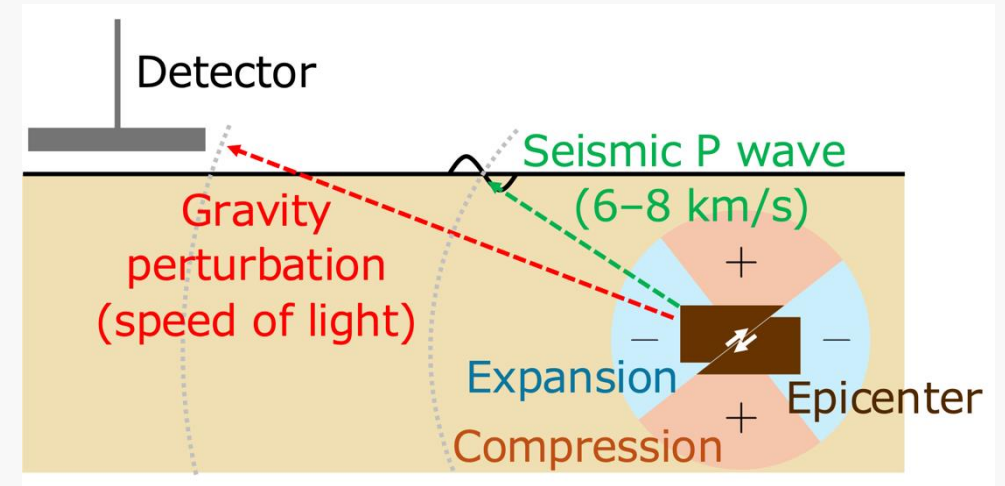
Newtonian noise



- 検出器付近の物体の運動による最も原理的な雑音
- 未観測だが、次世代検出器の感度を制限すると予想される

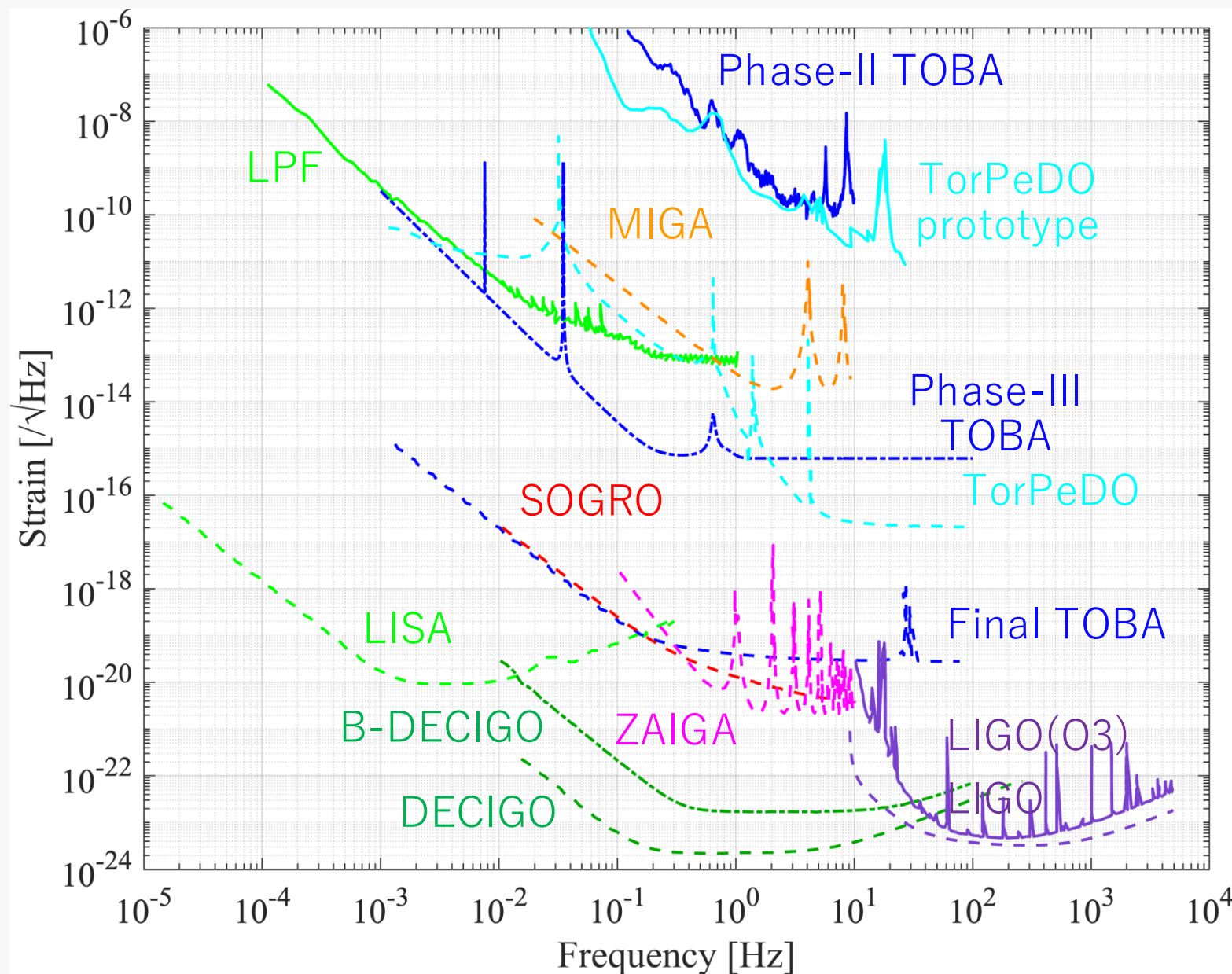
contribute
3G detectors

Earthquake alert system



- 地震による重力場変動が光速で伝搬
- 従来より速い地震速報を実現可能

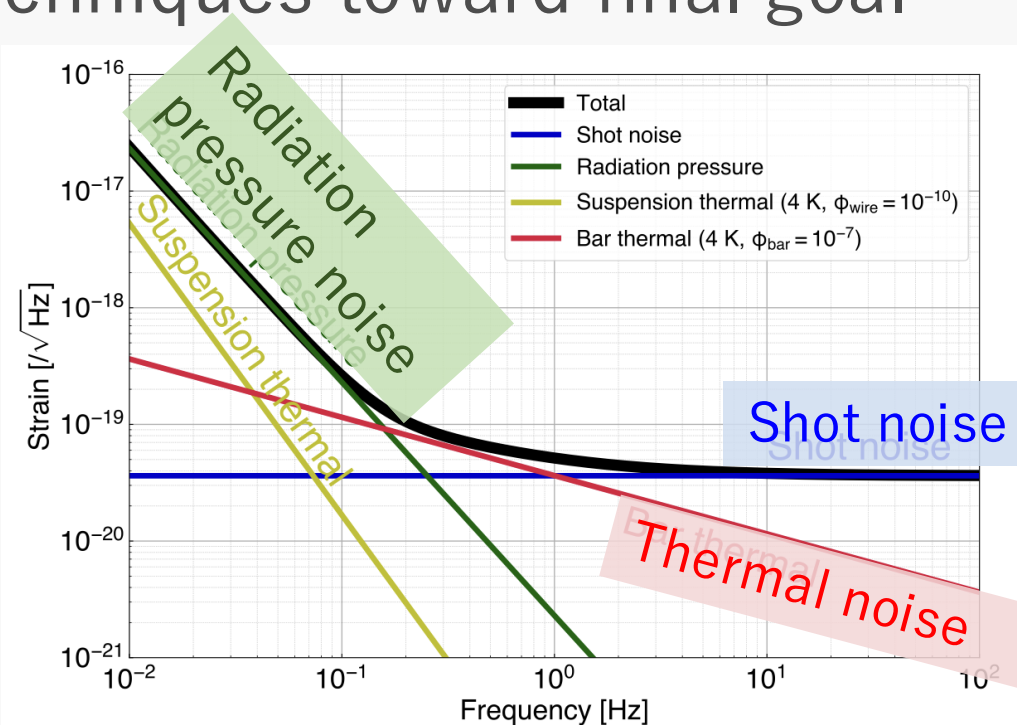
mitigate
disaster damage



実線: 達成
点線: 提案

$10^{-8}/\sqrt{\text{Hz}}$ @0.1Hz
くらいが最高

- TOBA; TOrsion Bar Antenna
- Low resonant freq. of torsion pendulum($\sim$ mHz) enables low freq. GW observation
- Developing prototype and component techniques toward final goal
- Designed sensitivity is limited by quantum noise and thermal noise



Demonstrate technique and improve sensitivity

Toward GW detection

Future

~Present

~2014

Phase
I, II

Principle test

demonstration of torsion bar

- 20 cm TMs
- Room temperature
- MI sensor
- Achieved $10^{-8}/\sqrt{\text{Hz}}$
- Set an upper limit on GW background

Phase
III

Development

Cryogenic demonstration,
noise reduction

- 30 cm TMs
- cryogenic(4K)
- Fabry-Perot cavity sensor

Target sensitivity:
 $10^{-15}/\sqrt{\text{Hz}}$ @0.1 Hz

Final

Final goal

upsizing

- 10 m TMs
- cryogenic(4 K)
- FP cavity sensor
- $10^{-19}/\sqrt{\text{Hz}}$

Low freq. GW

IMBH binary merger

- ~ 1 Mpc (Phase-III)
- ~ 10 Gpc (Final)

Unveiling the formation of
supermassive BH

GW background

Final

- $\Omega_{\text{GW}} \sim 10^{-7}$ @0.1 Hz

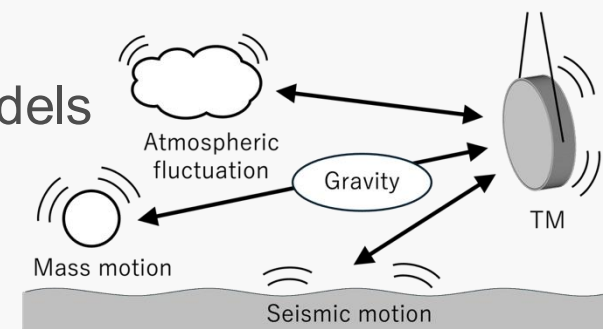
Direct observation of the early universe

Geophysics

Newtonian noise detection

- Expected to limit 3G detector's sensitivity
- not detected yet
- Testing current models

Contribute
3G detector



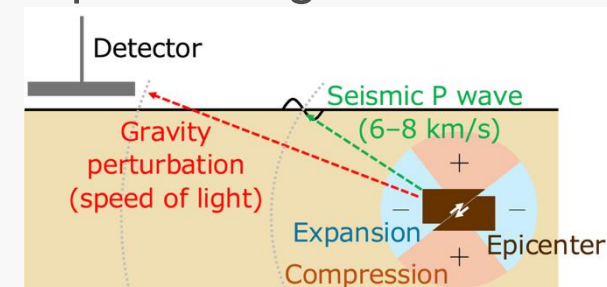
Phase-III

Earthquake alert

Phase-III

- Earthquake causes gravity gradient
- GG propagates at the speed of light

Mitigate
disaster damage



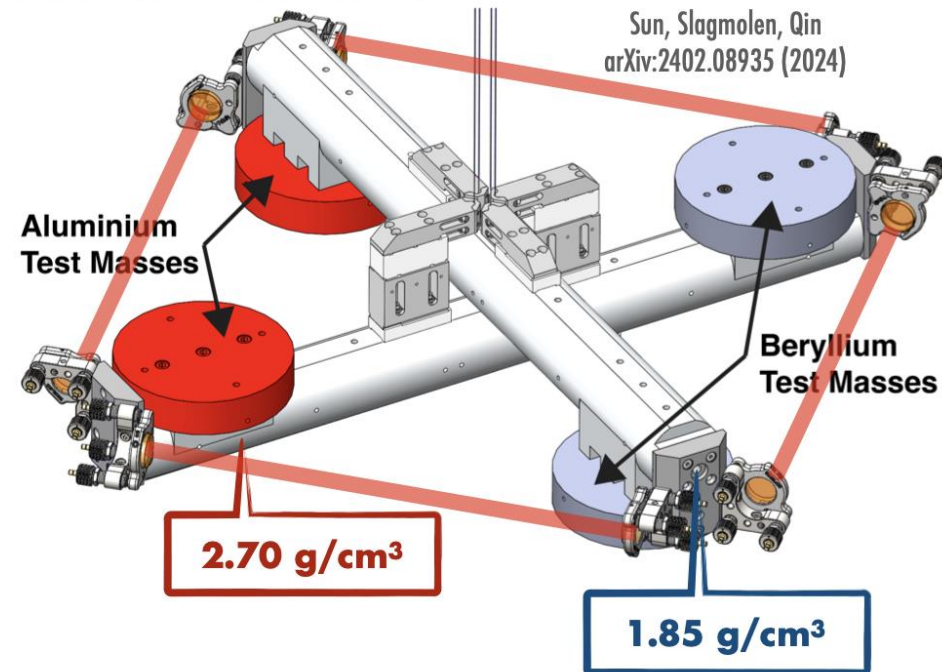
Differential Torsion Sensor for Dark Matter

- A new type of differential torsion sensor for vector dark matter detection

Test mass materials	Al / Be
Mass of each test mass	5 kg
sensor beam length	0.6 m
beam mass (w/ mounting equipment)	5kg
Mechanical lever arm	0.24 m
Mechanical resonant frequency	26 mHz
Suspension wires Q factor	3000

- ➔ Large common mode rejection
- ➔ Differential angular sensitivity
- ➔ Can achieve enhanced sensitivity without requiring significant upgrades in technology

[Also see relevant study – Okuma et al., arXiv:2404.00250 (2024)]



Require same moment of inertia
→ need mechanical engineering to make test masses with the same external dimensions

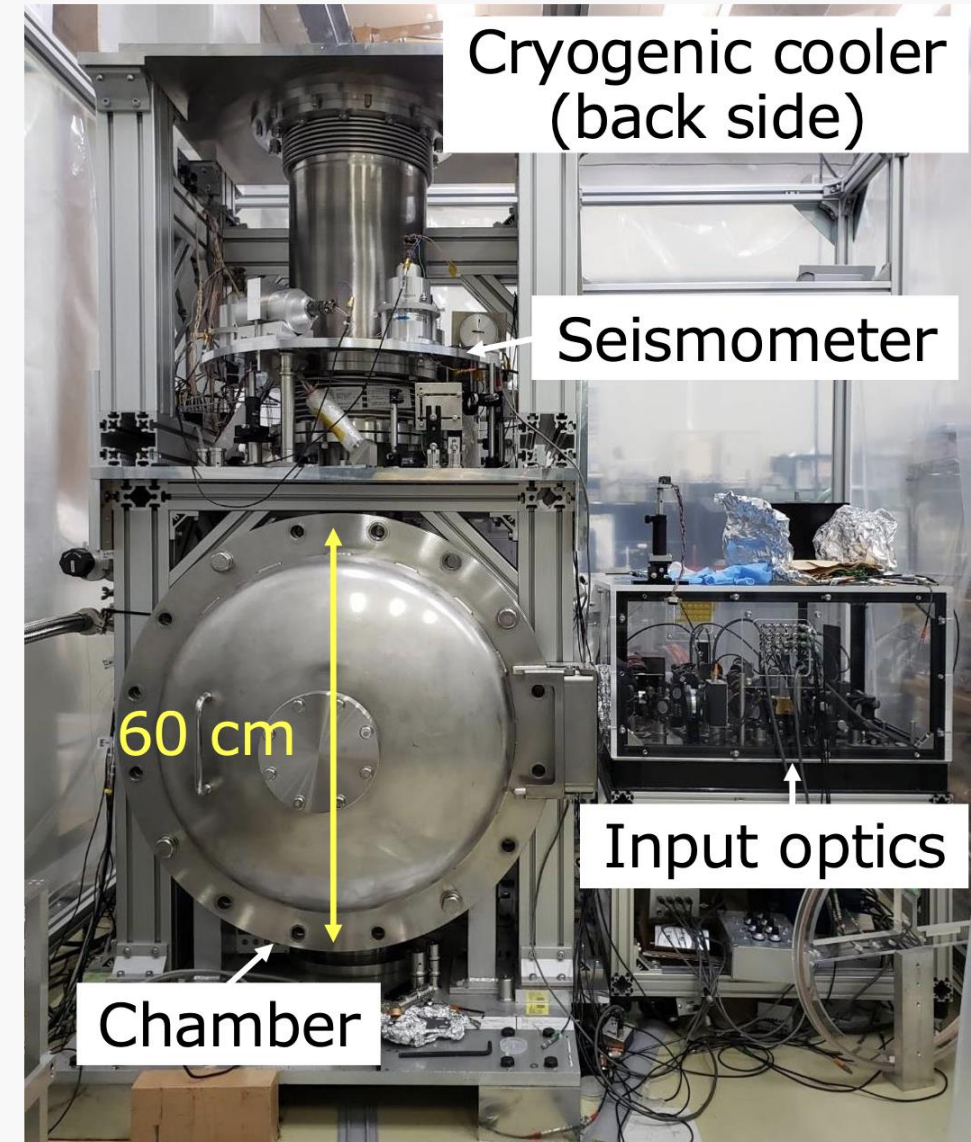
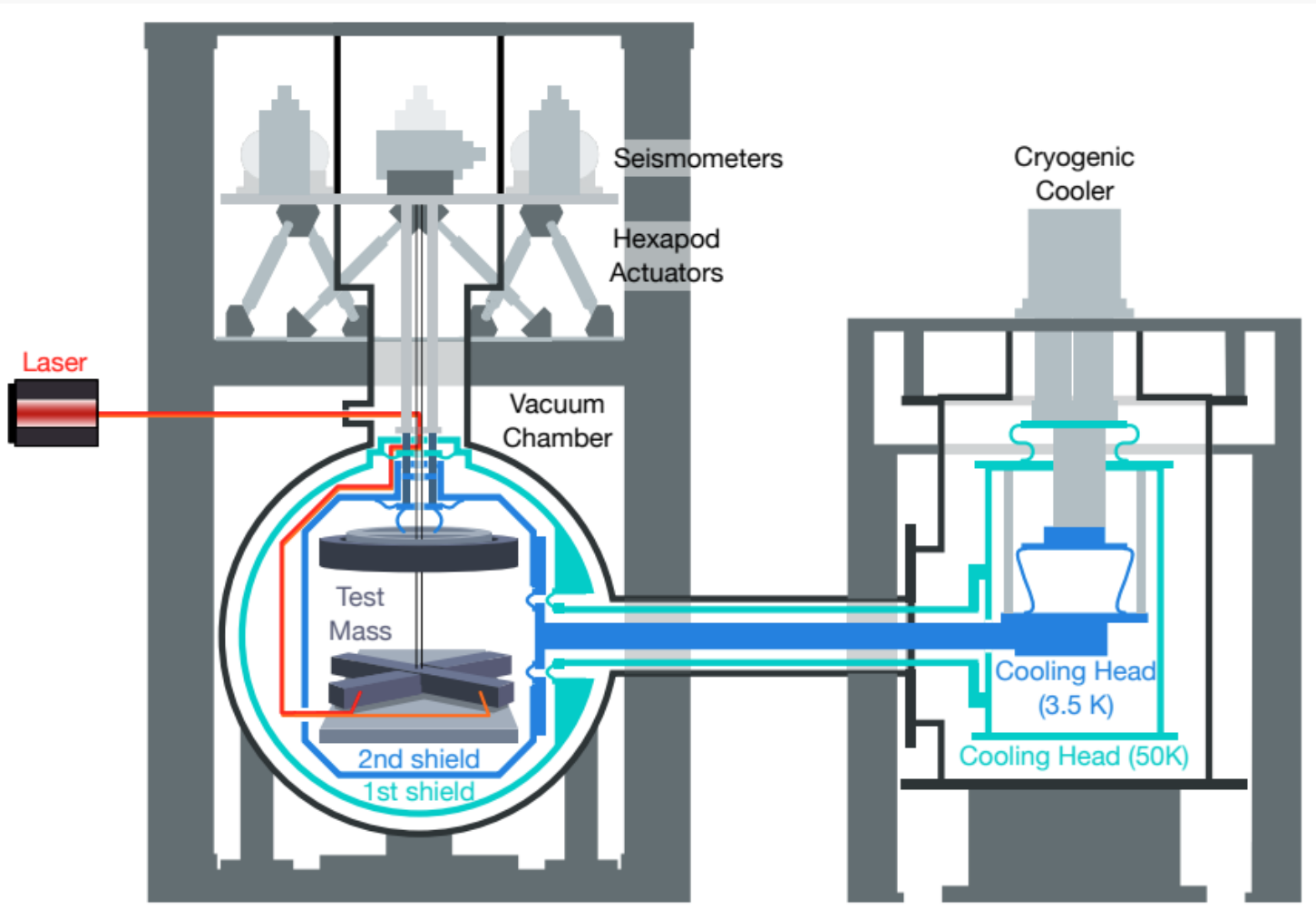
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Configuration of Phase-III TOBA

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TOBA consists of VIS, Suspension system, Optical system, Cryocooler



Configuration of Phase-III TOBA

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Vibration isolation system

- Feedback control with seismometer and actuator

Active vibration isolation

Suspension system

- TMs made of silicon
- Suspended by low loss sapphire fiber

TMs respond to GW

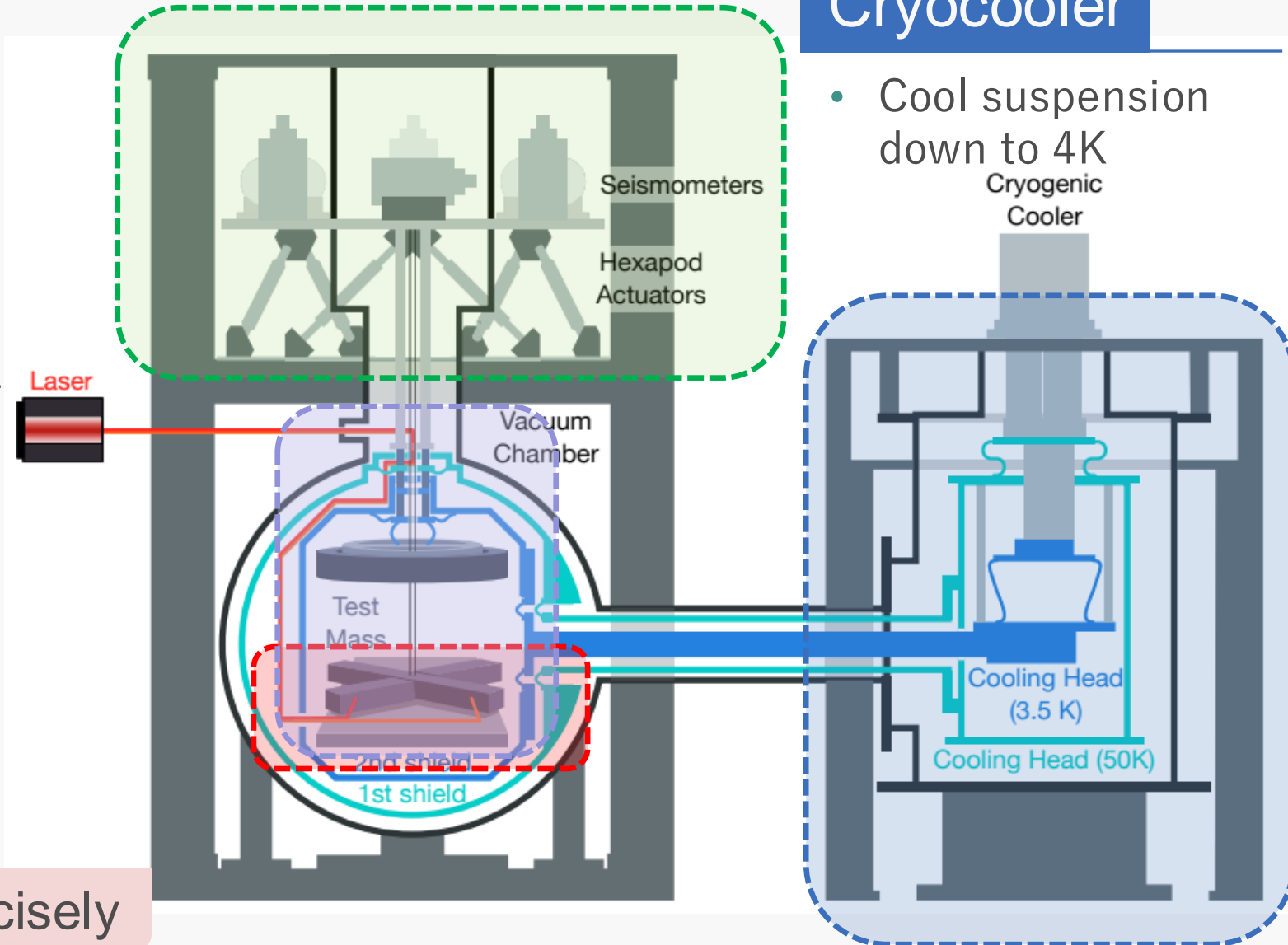
Optical system

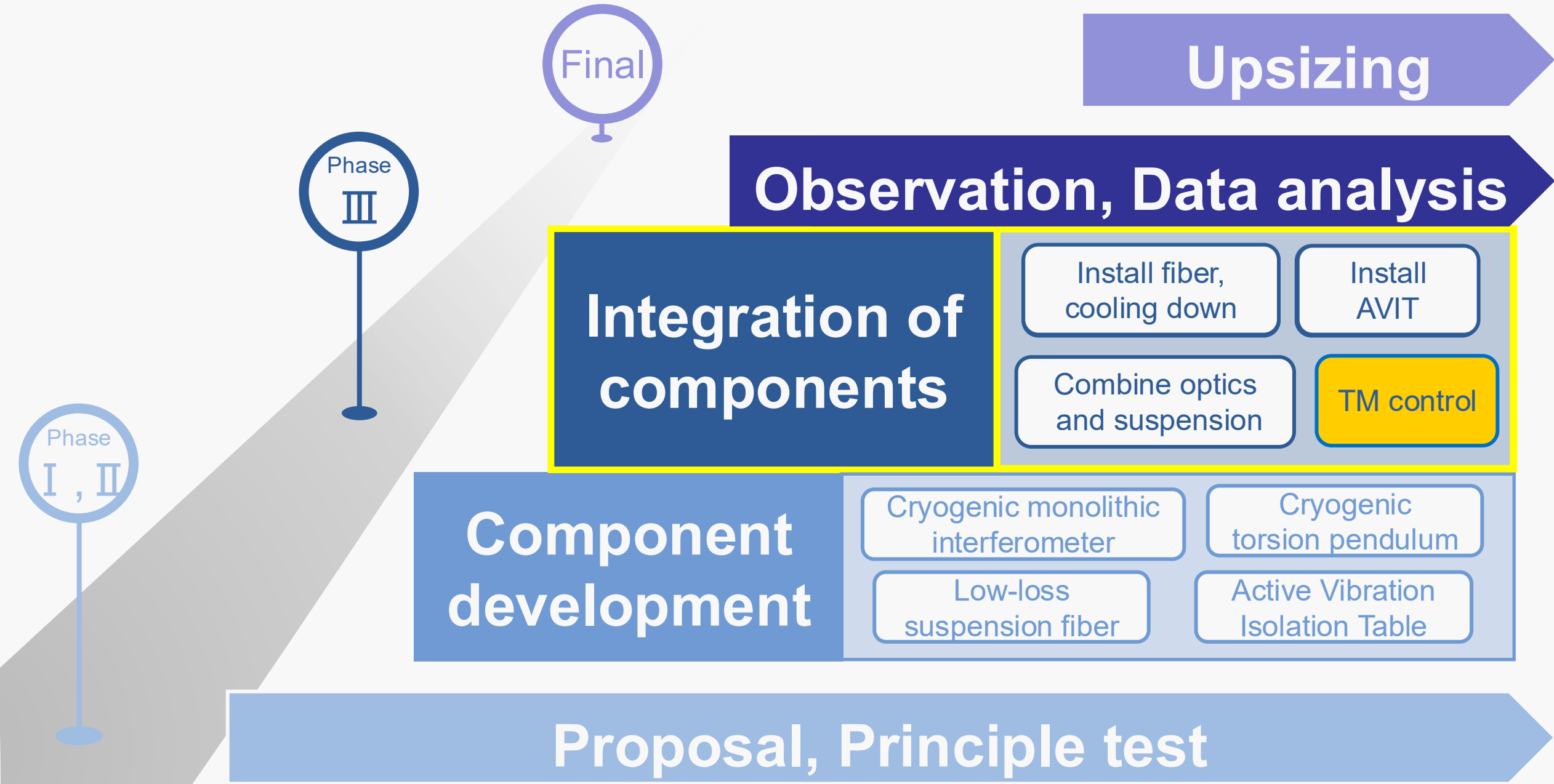
- Differential Fabry-Perot cavity

Readout the rotation of TMs precisely

Cryocooler

- Cool suspension down to 4K

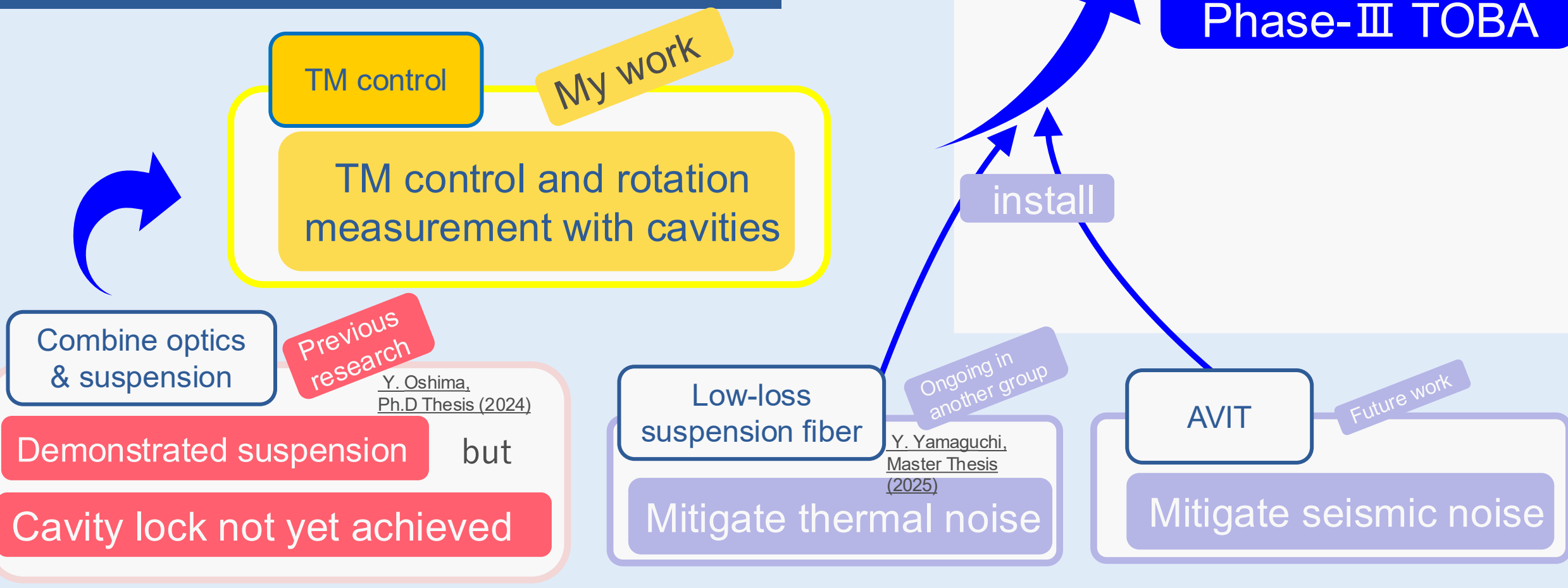




Integration of components

Observation, Data analysis

Completion of
Phase-III TOBA



Integration of components

Observation

Completion
of Phase-III

- Noise hunting
- Seismic coupling mitigation
- Operation in cryo. Temp.

install

Combine optics
& suspension

Previous
research

Y. Oshima,
Ph.D Thesis (2024)

Demonstrated suspension

but

Cavity lock not yet achieved

Low-loss
suspension fiber

Ongoing in
another group

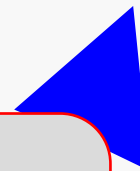
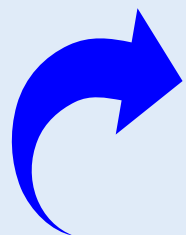
Y. Yamaguchi,
Master Thesis
(2025)

Mitigate thermal noise

AVIT

Future work

Mitigate seismic noise

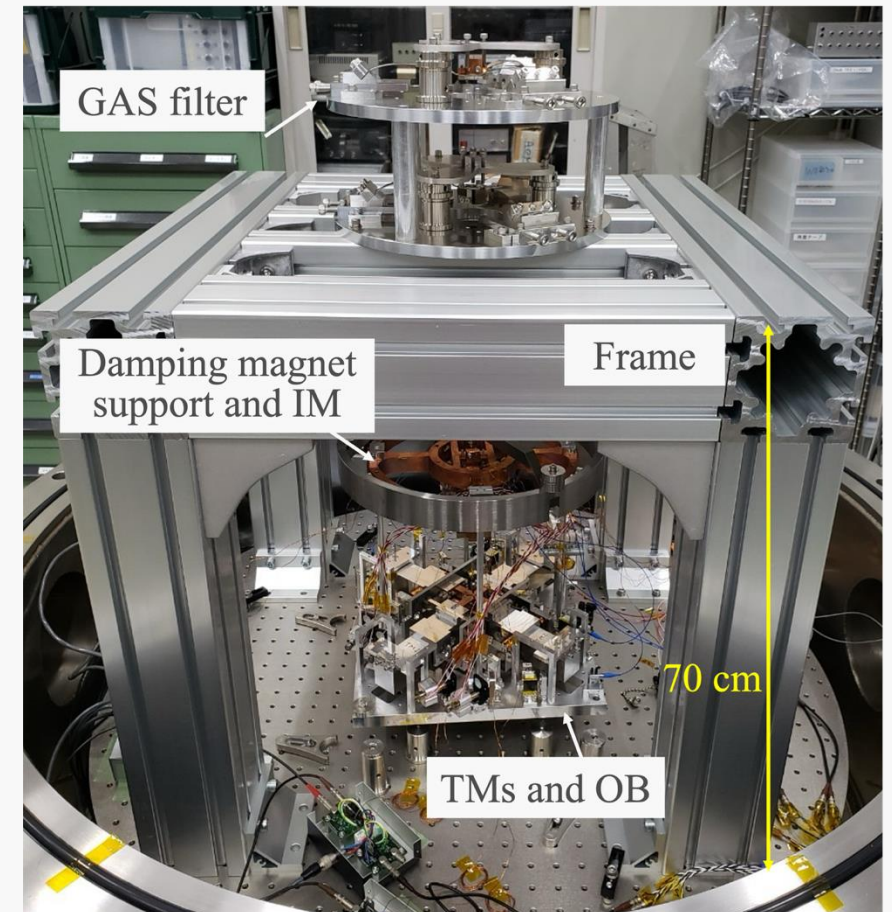
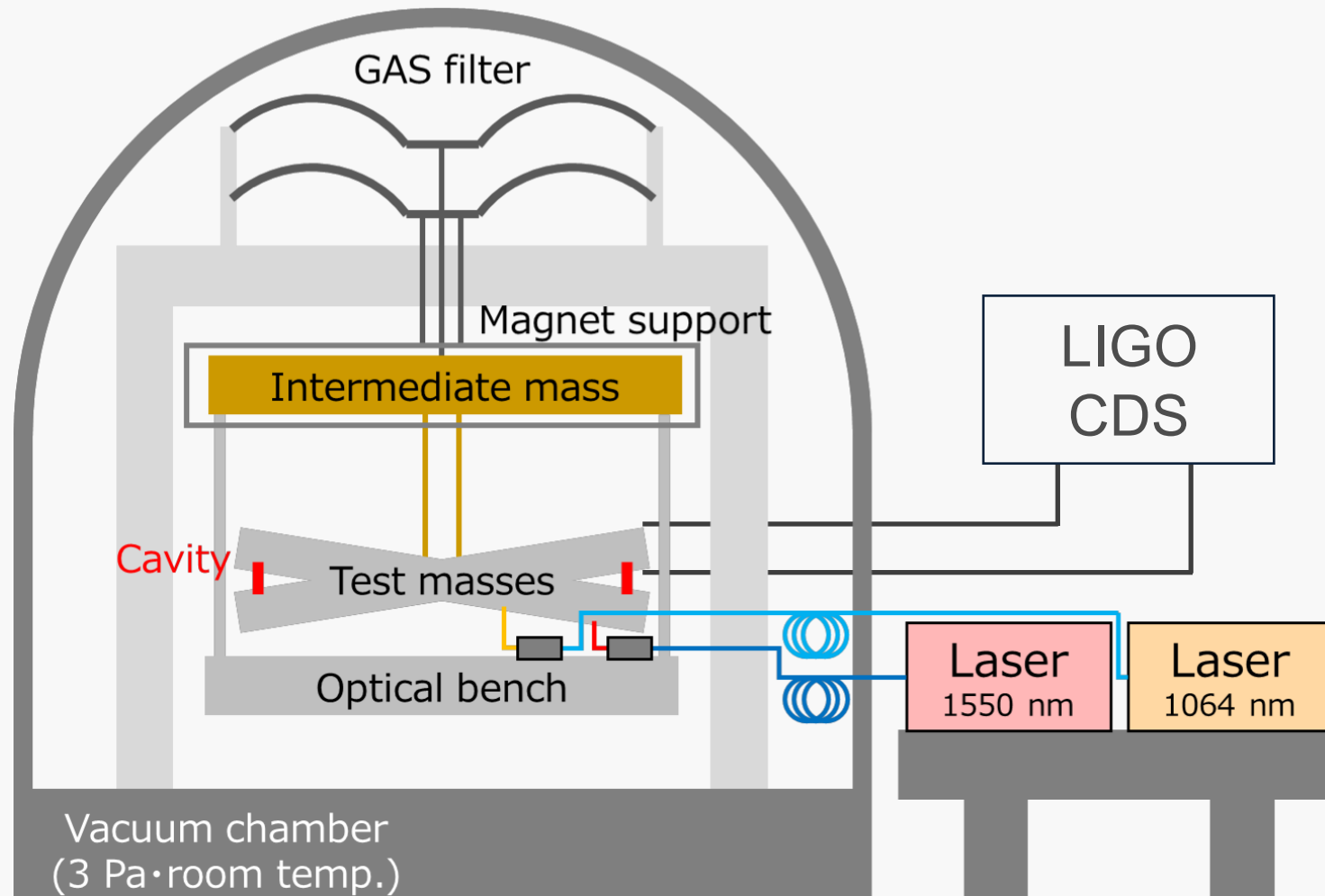


Experimental setup

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Suspension system is placed in the chamber

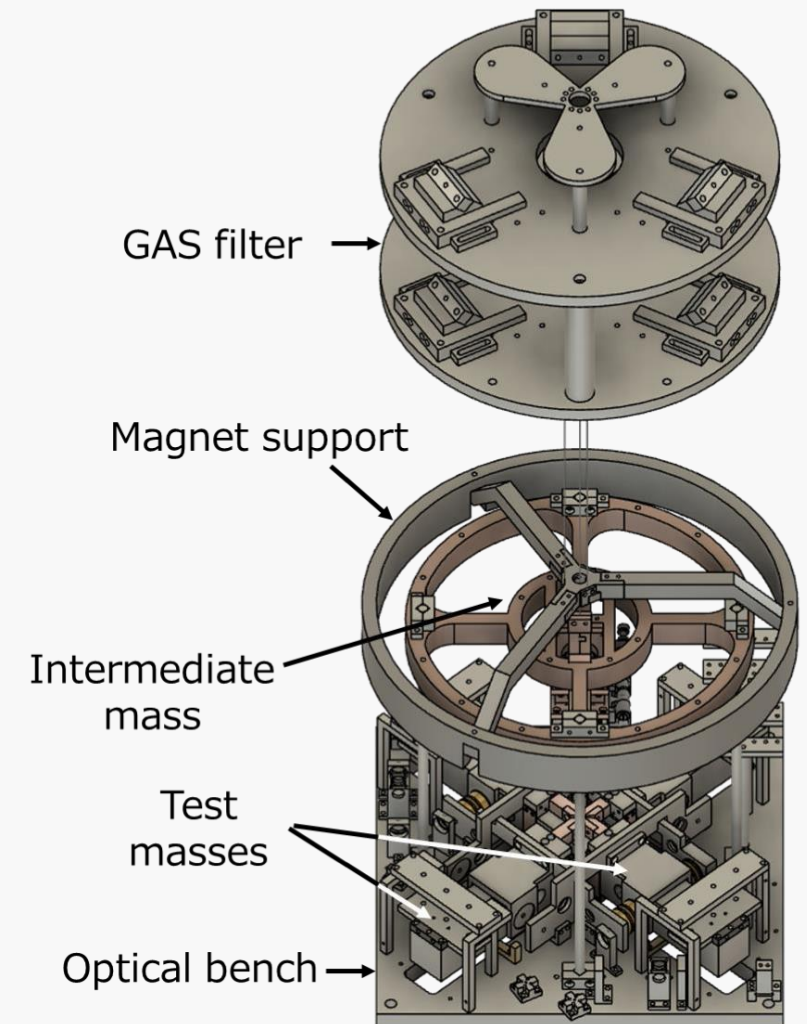
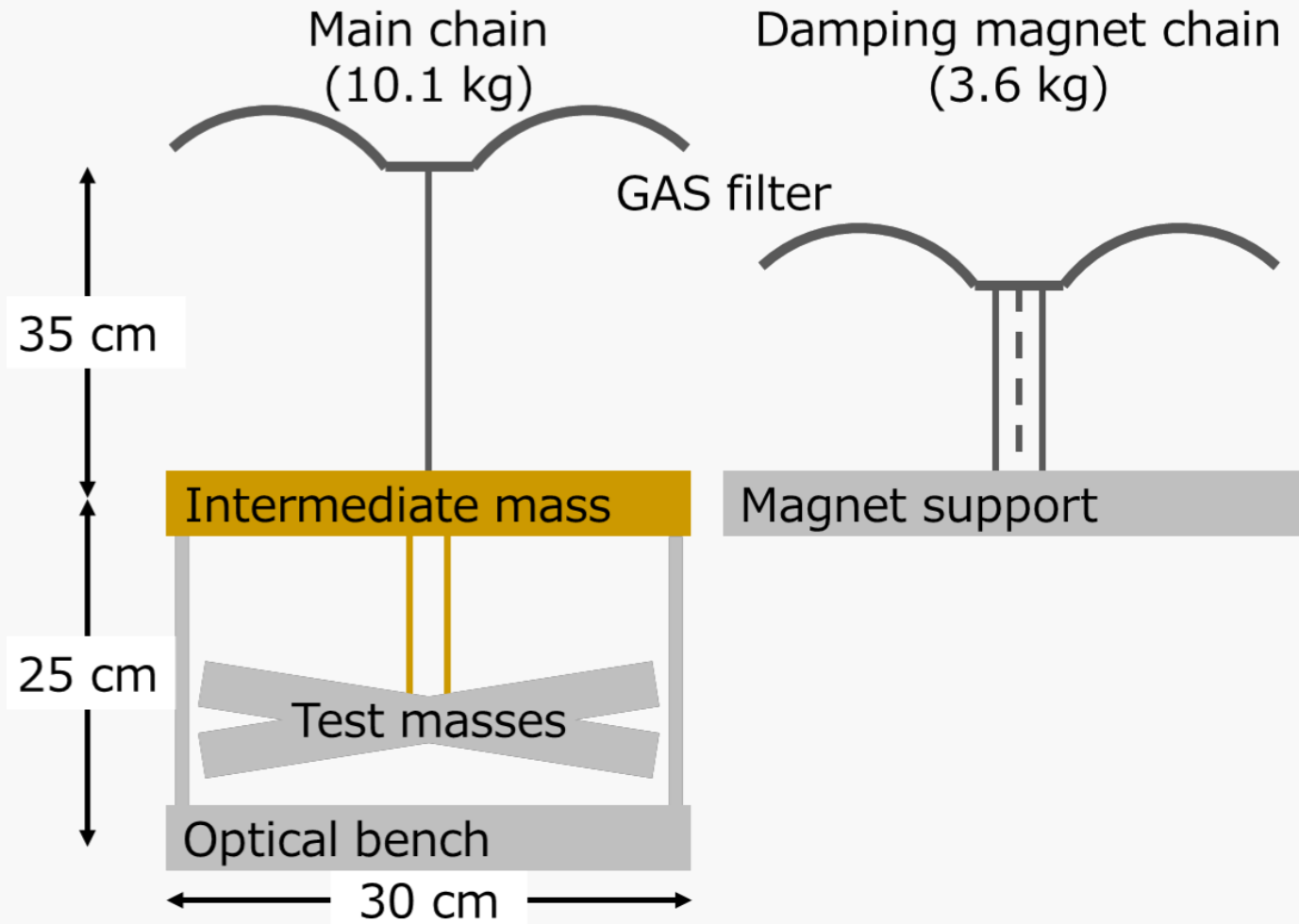
Laser beam is introduced to the chamber from the source via fibers



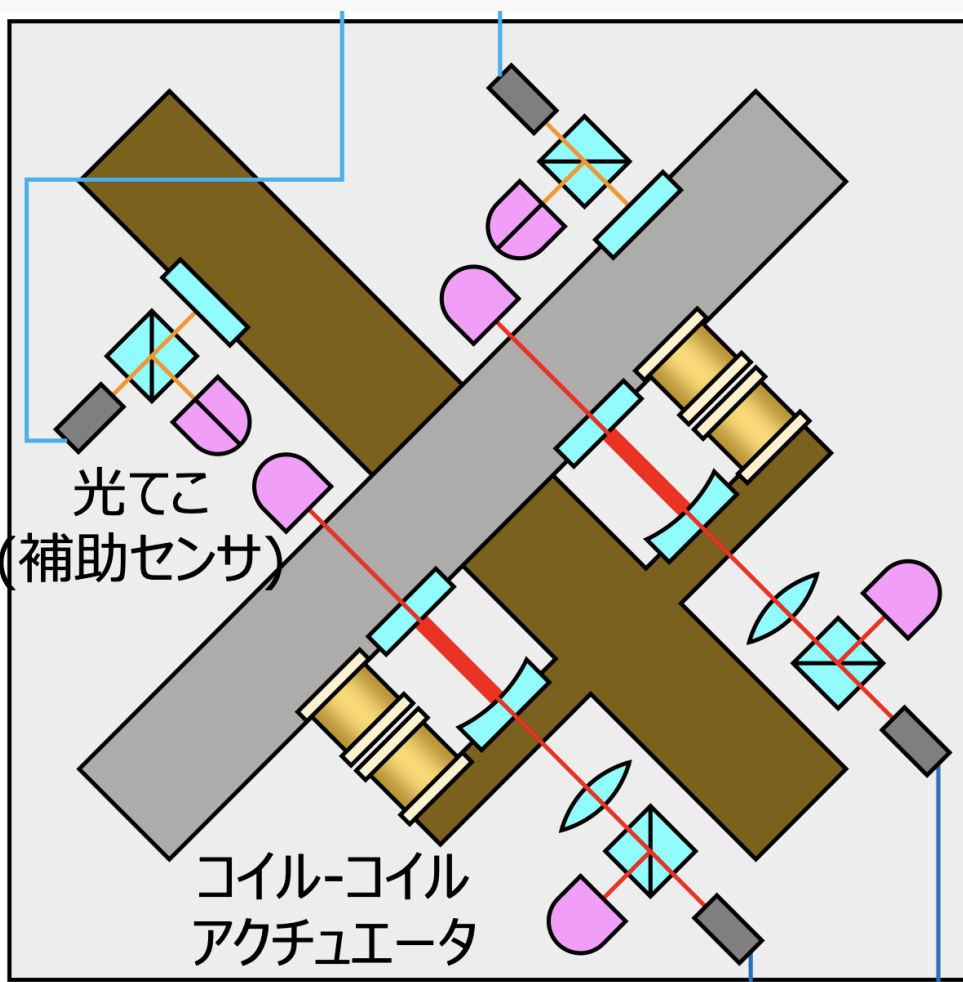
Design of suspension system

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Suspension system consists of two chains



Top view of OB



— レーザー 1550 nm
(高さ: 25 mm)

— レーザー 1064 nm
(高さ: 35 mm)

Differential FP cavities

- 2 FP cavities between TMs to detect differential rotation

Reduce noises from OB vibration

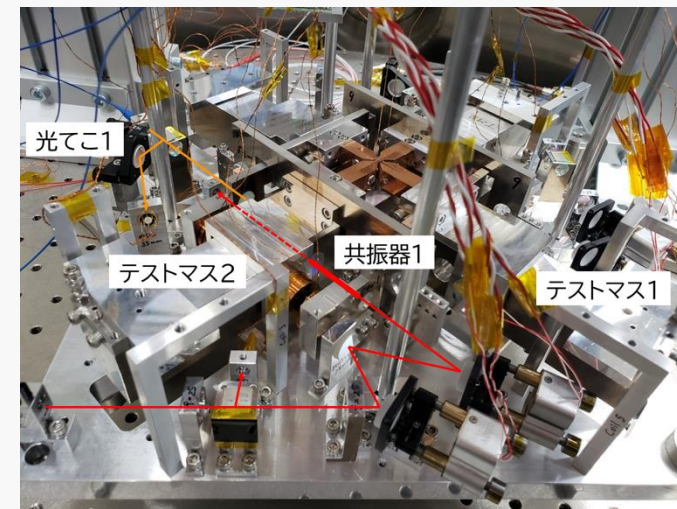
Actuator

- Actuate TMs with coil-coil actuator

Reduce magnetic field noise

OpLev

- Auxiliary sensor



Torsion pendulum control with FP cavities

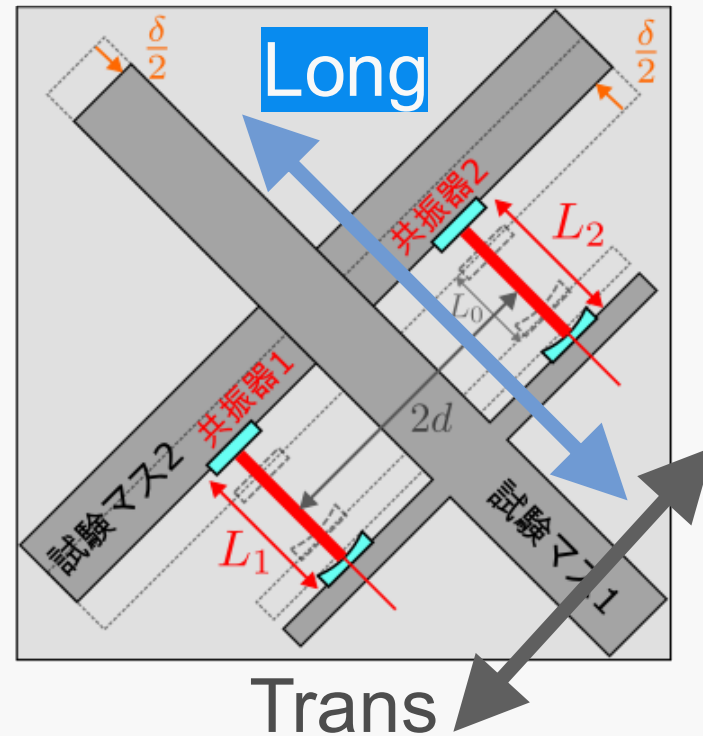
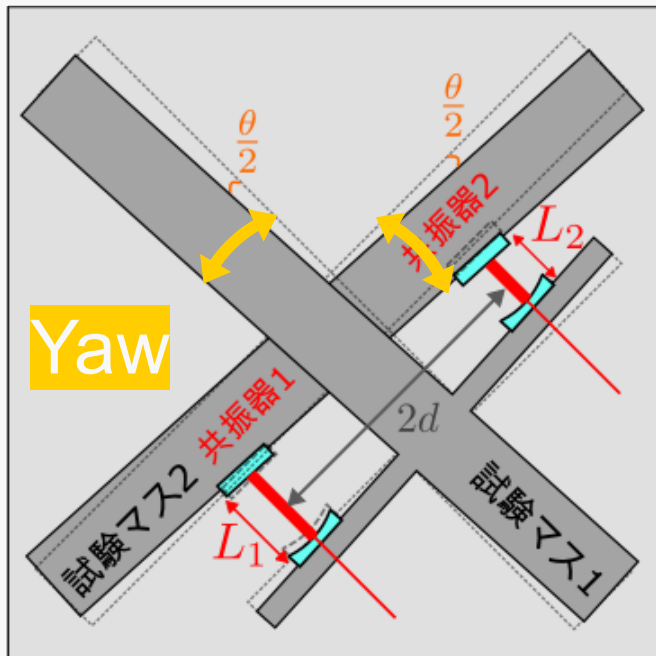
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Control TMs' rotational motion with FP cavities

Yaw rotation and Long movement changes cavity length

2DoF control is required

Non-obvious method is necessary



Yaw

differential

$$\begin{cases} L_1 = L_0 + d\theta \\ L_2 = L_0 - d\theta \end{cases}$$

Long

common

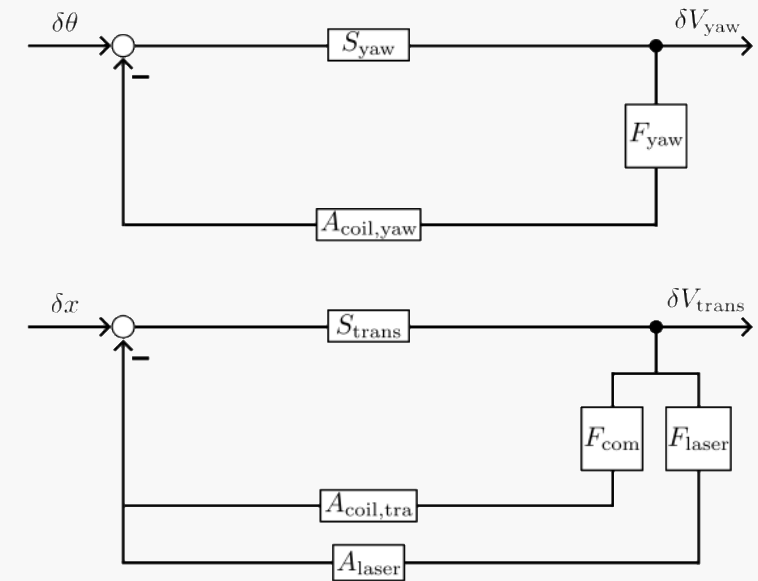
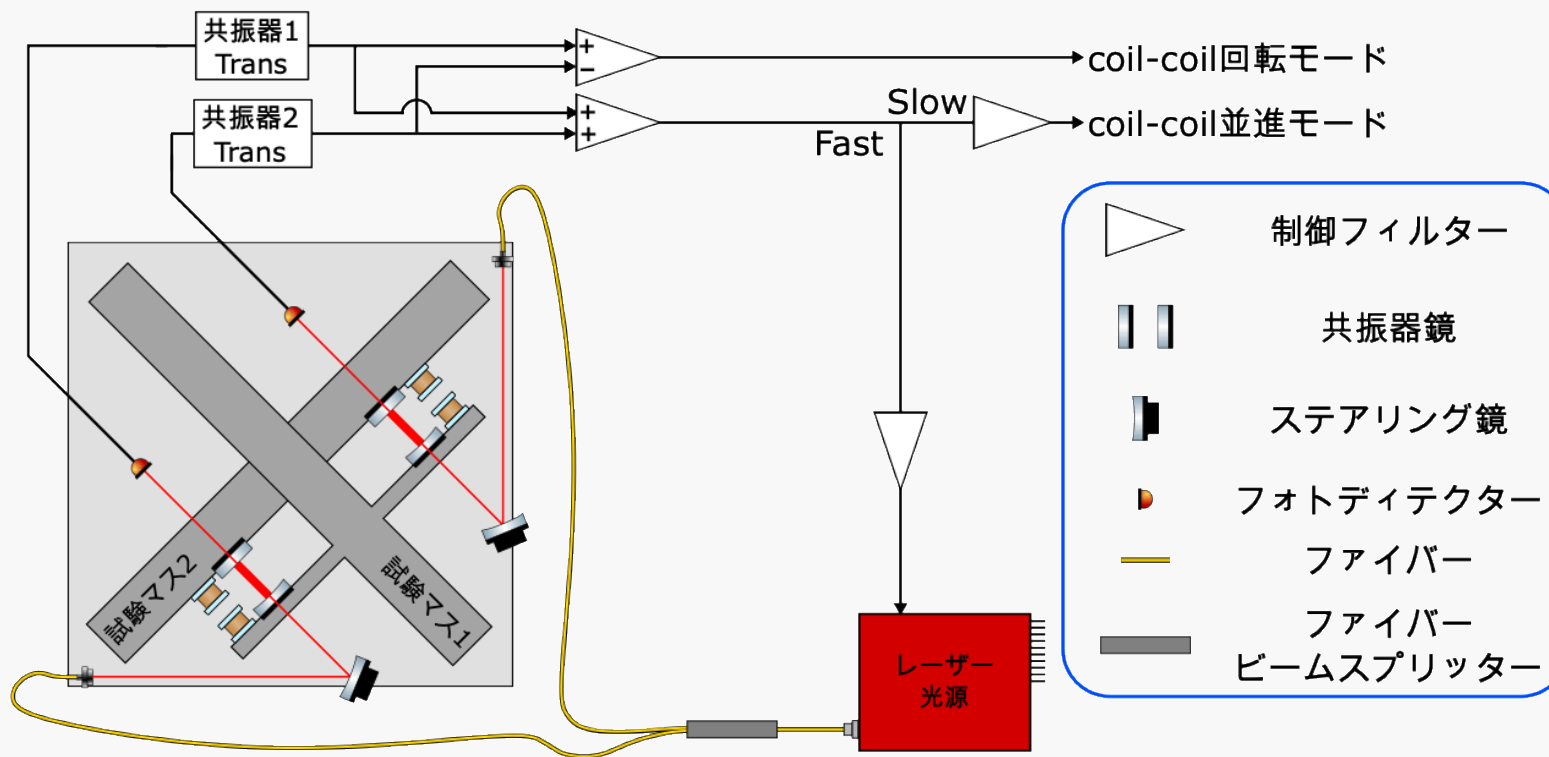
$$L_1 = L_2 = L_0 + \delta$$

Torsion pendulum control with FP cavities

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Feed back control for each mode, Yaw and Long

Feed back { Yaw signal to coil-coil actuator
Long signal to coil-coil and laser frequency actuator



Sensor (Fabry-Perot cavity)

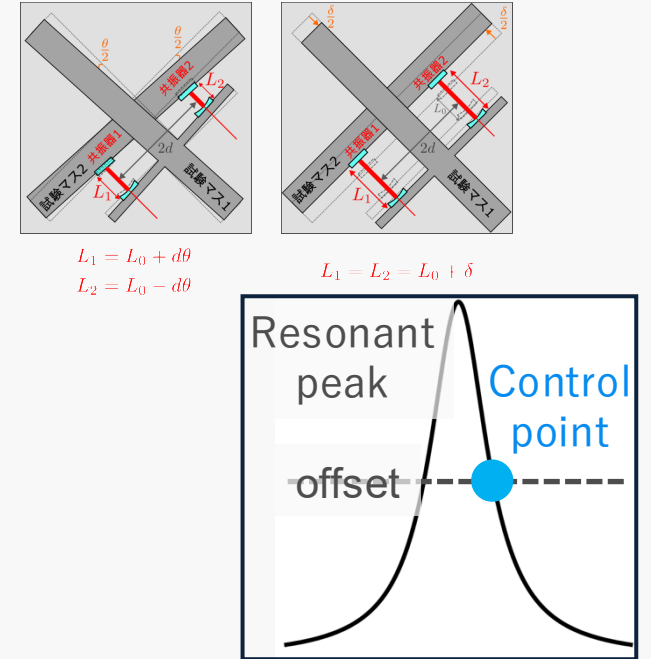
Yaw error signal from differential signal

Long error signal from common signal

Locking to the cavity transmission slope via DC offset

Offset lock

- Larger control range
- Don't need RF PD



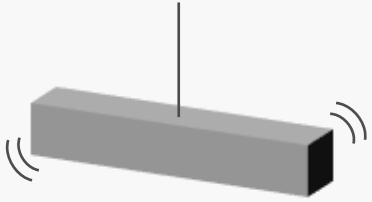
Actuator (coil-coil, laser frequency)

Yaw actuate with differential mode of coil-coil actuator

Long actuate with common mode of coil-coil actuator & frequency actuator

Frequency actuator works symmetrically on both cavities

Lock acquisition: the process of acquiring controlled state from un-controlled state



Acquiring controlled state from free-run(un-controlled) state requires actuator with **high authority & fast response**

In this work...

Requirement

- ① Bandwidth ~1kHz
- ② 2DoF lock acquisition

solution

Bandwidth~5kHz

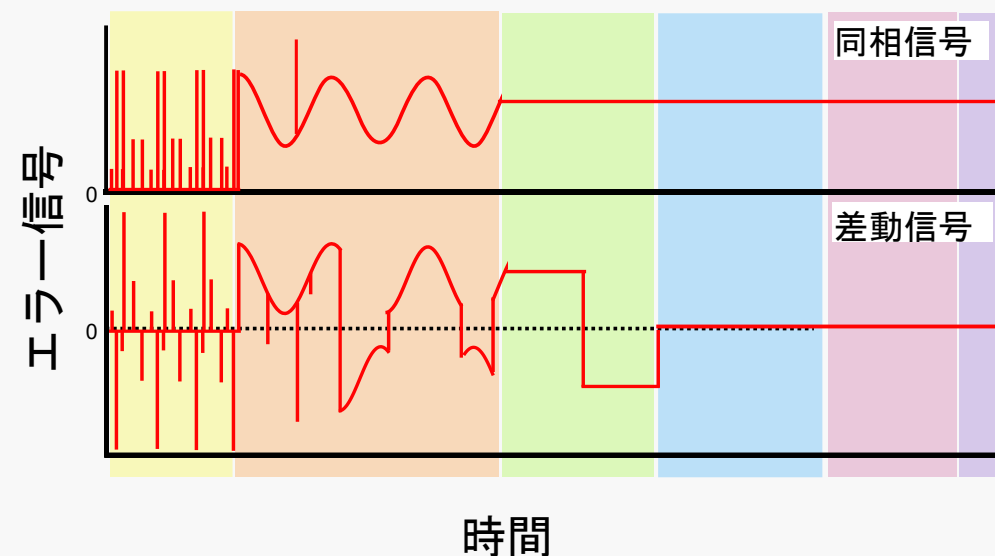
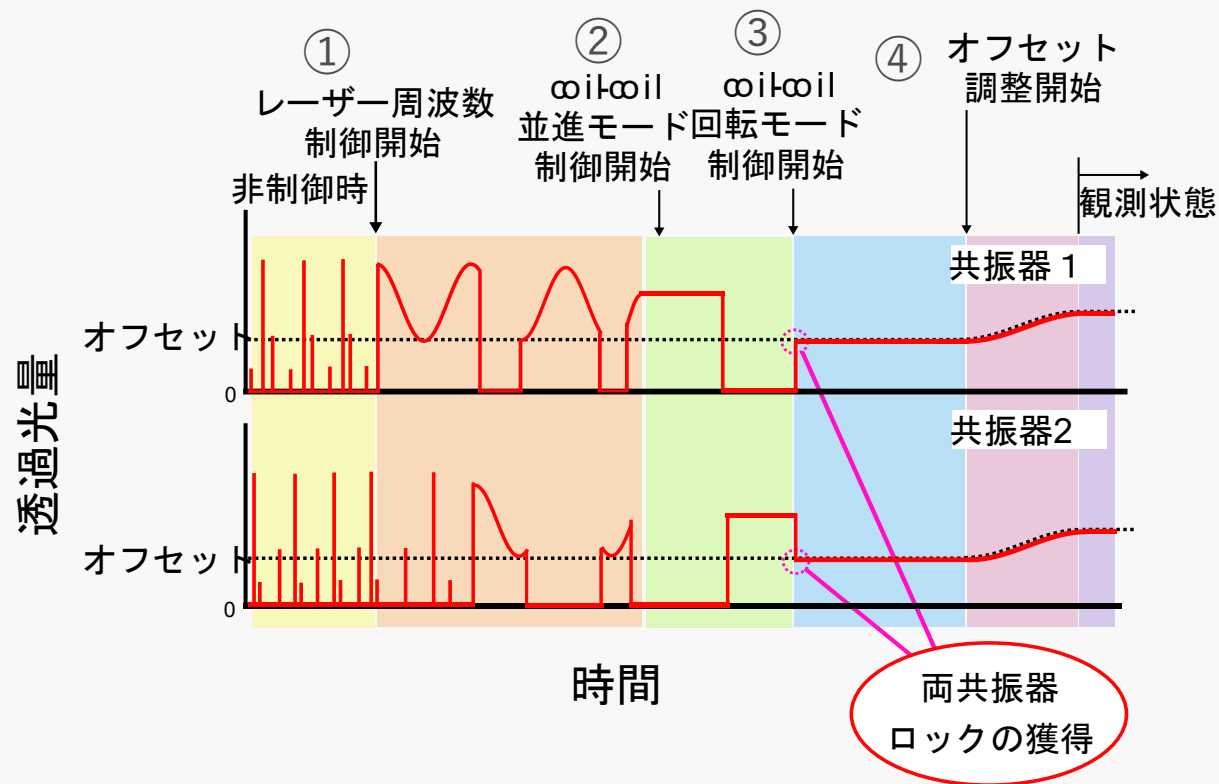
① **laser frequency actuator**

Bandwidth of coil-coil actuator is limited ~50Hz by CDS delay

② **Stepwise** lock acquisition(DoF by DoF)

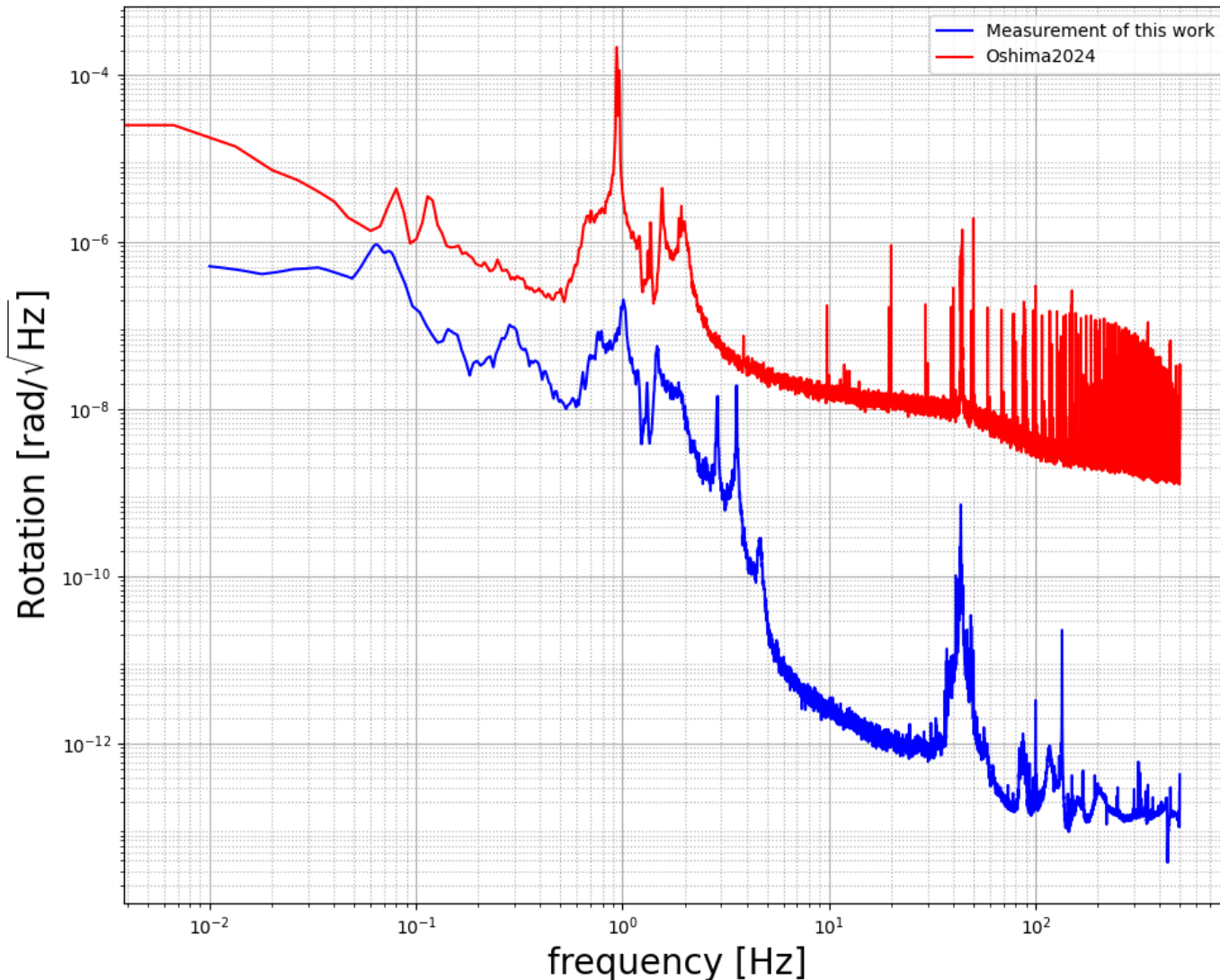
Lock common mode with freq. actuator first.
After that, lock differential mode.

Method of lock acquisition



- ① locking common mode with freq. actuator
- ② stabilize common mode with coil-coil common mode
- ③ locking differential mode with coil-coil differential mode
- ④ adjusting offset levels

Rotation sensitivity



— OpLev in previous research

Y. Oshima,
Ph.D Thesis (2024)

— This work

Measured TMs' rotation with
main sensor; FP cavities

Compared to previous result,
Sensitivity improved

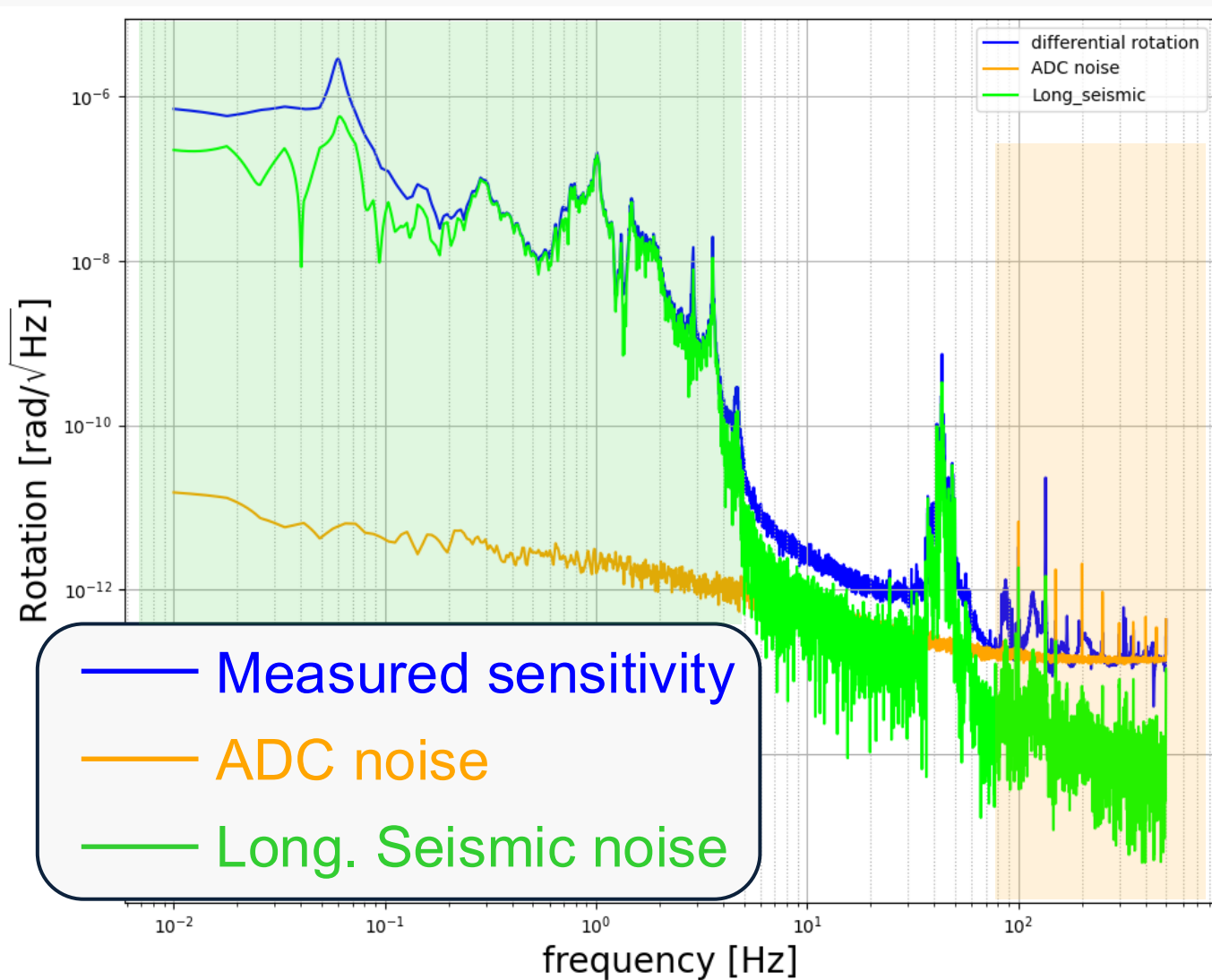
~2 orders below 4Hz

~4 orders above 4Hz

Mainly due to

- Reduction of sensor noise
- Mitigation of OB vibration noise

Rotation sensitivity and Noise budget



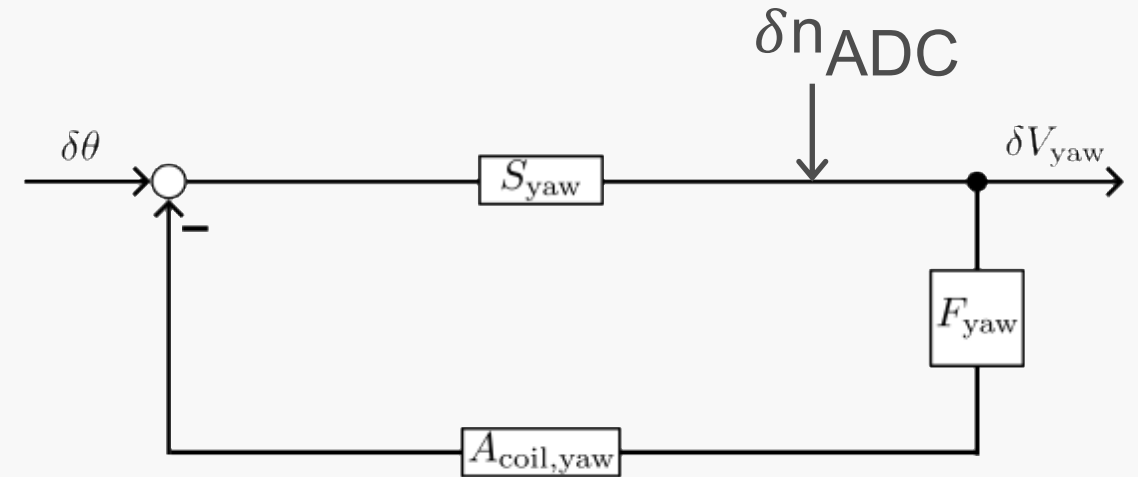
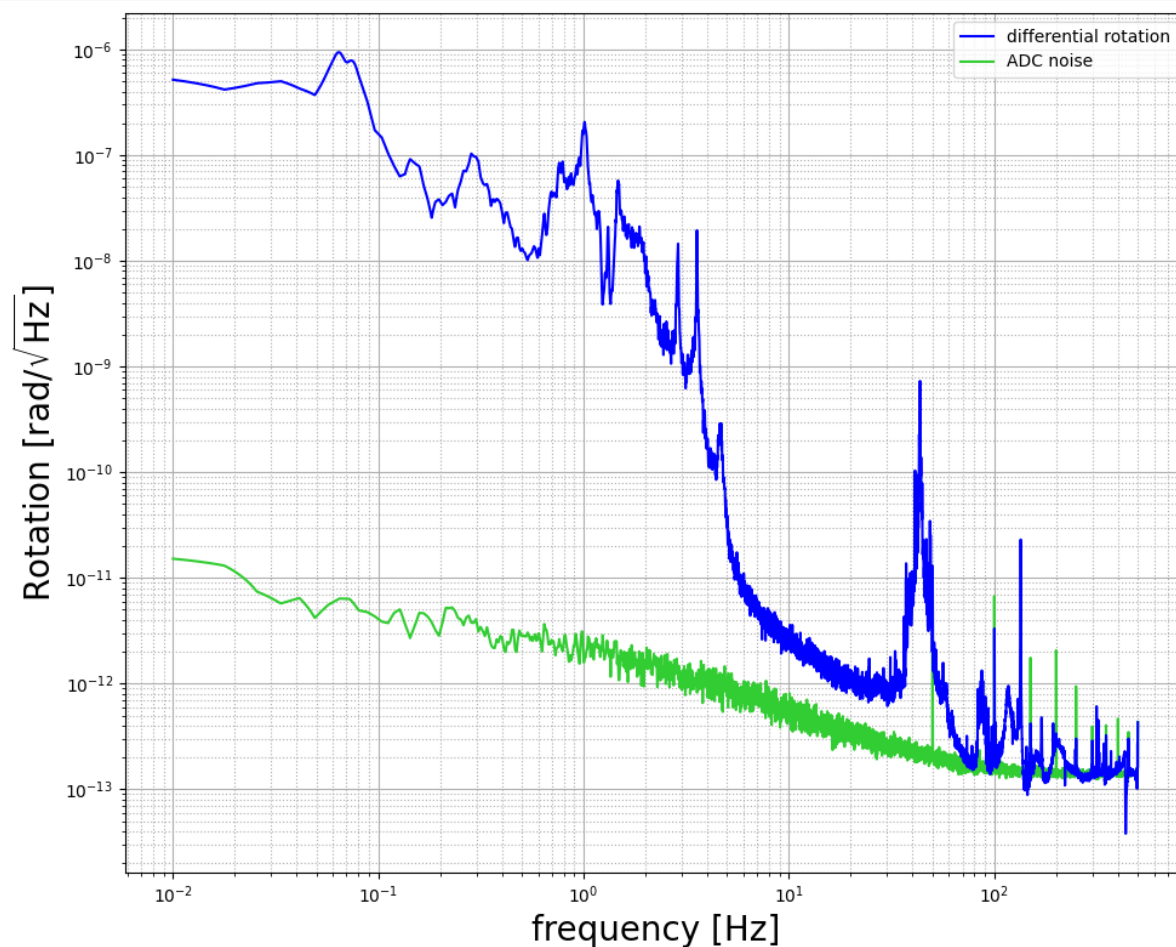
0.2-5Hz: seismic noise

above 60Hz: ADC noise

UGF of cavities control is ~20Hz

Digital system noise

Noise floor above 60Hz is limited by **ADC noise**

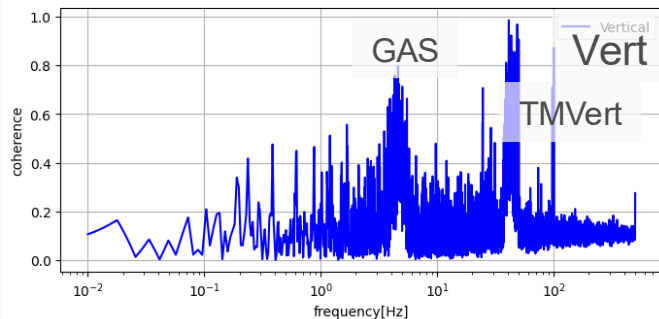
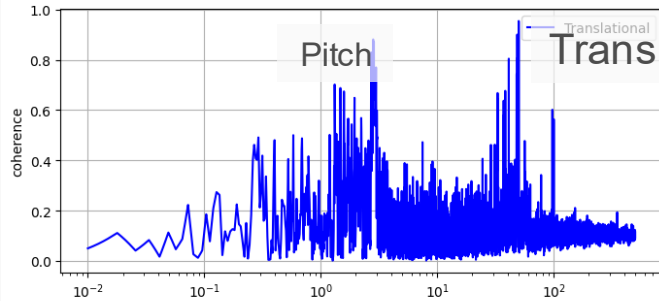
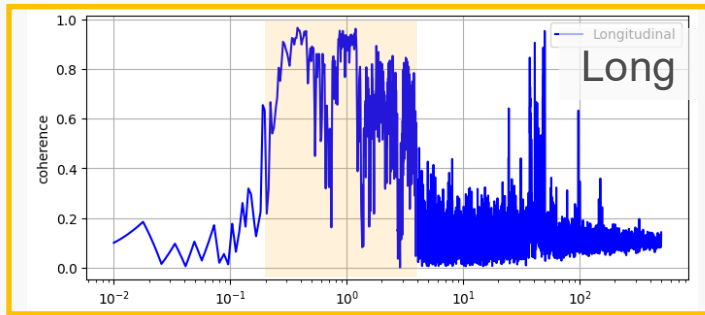


Amplification of the signal before CDS
mitigates ADC noise

Discussion of noise source ② seismic noise 32/49

In principle, **Translational seismic motion** is not a noise source, but it couples to imperfection of system(ex: TM tilt) to be rotational noise

Estimated seismic motion contribution from coherence with seismometer signal

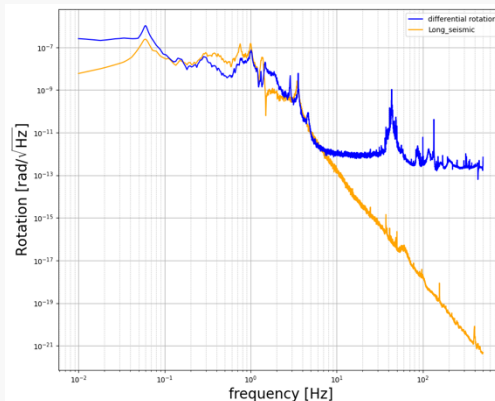


Significant coherence with **Long. seismometer**

Trans, Vertとも各モードの共振ピークでは
コヒーレンスあり

Long. seismic noise is dominant

Couples to TM1 Pitch tilt,
TM2 Roll tilt



Long地面振動から回転角への伝達関数を測定し、
地面振動雑音のスペクトルを推定した

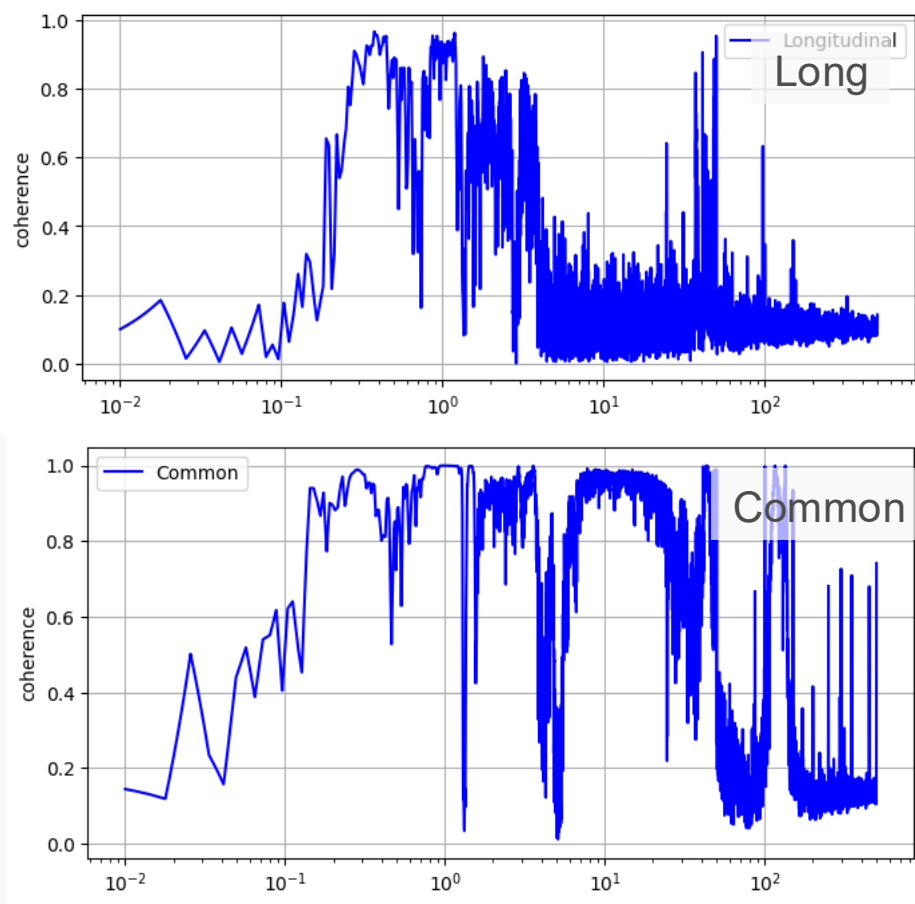
Adjustment of TM DC tilt

mitigates Seismic noise

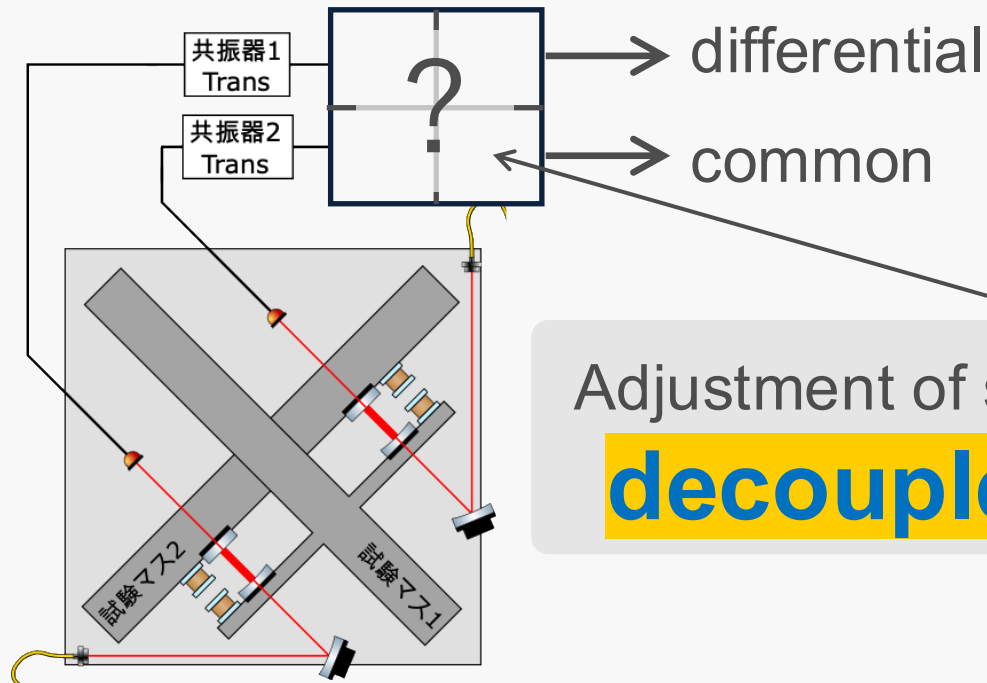
Discussion of noise source ③ signal coupling 3/49

If **common/differential decoupling** is insufficient,
longitudinal motion leaks into the rotation signal

As well as coherence with Long., coherence with common signal is also significant

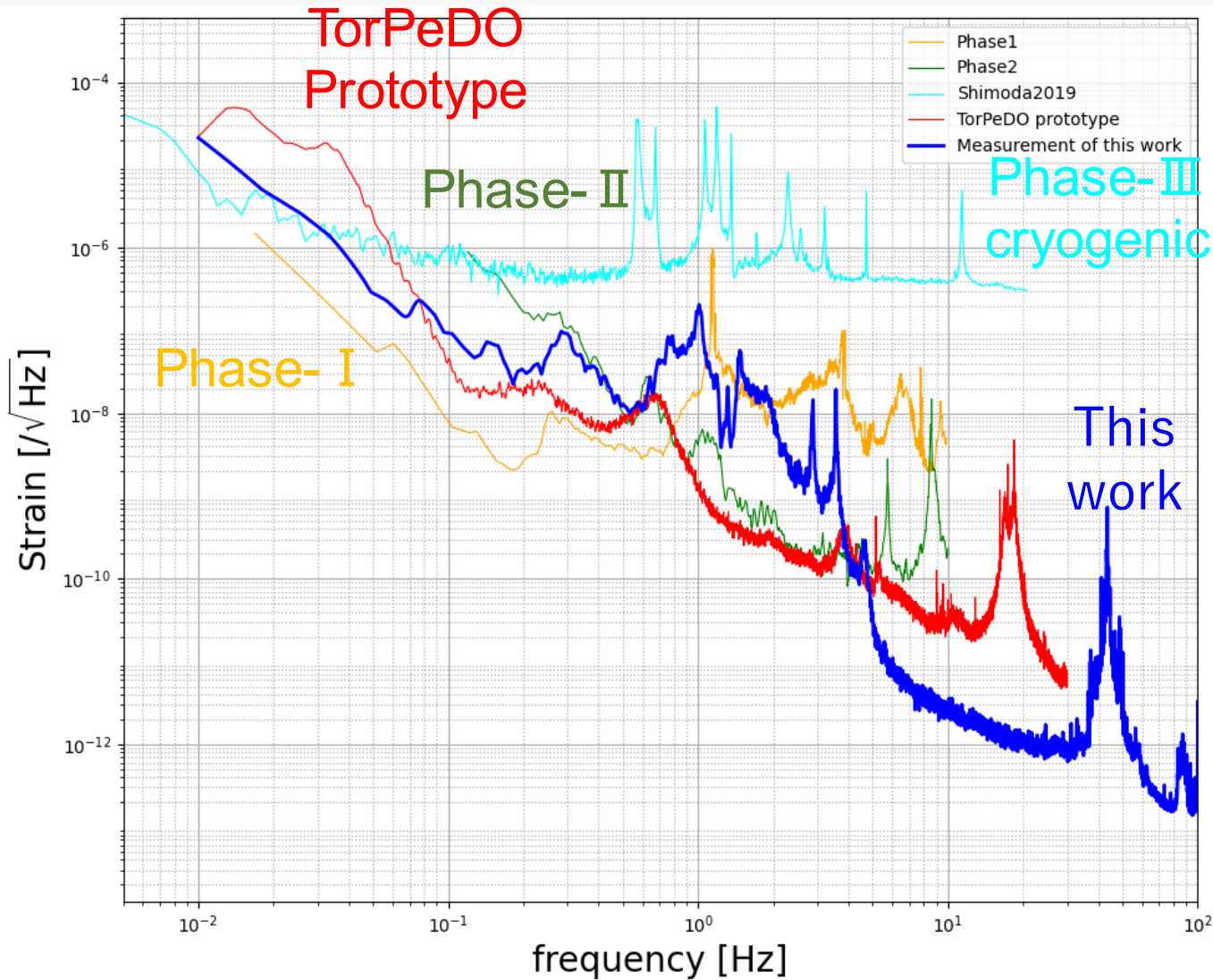


Overestimation of
seismic noise



Adjustment of sensor balance
decouples signal

Strain sensitivity



Best sensitivity in
Phase-III TOBA

Phase- I is still best at 0.1Hz

Tension in Sensor efficiency of cavity

Measured value

光てこの効率、
coil-coilアクチュエーターの
効率から順に逆算

gap

~3.5

Theoretical value

共振器スキャンのピーク値と
オフセット量から、
制御点での傾きを算出

larger smaller

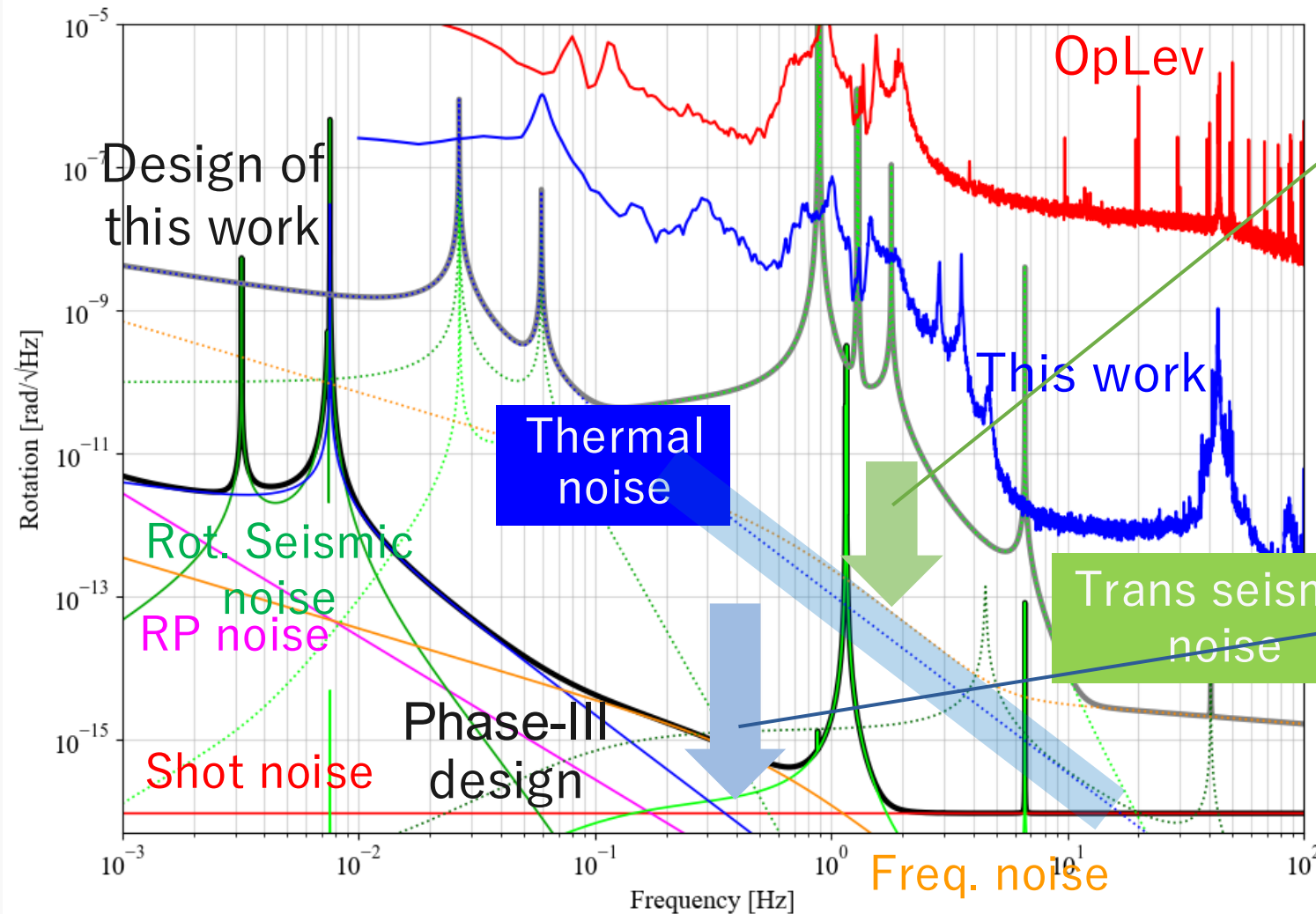
Generally, measured value can't be larger than theoretical value

(ミスアラインによるオプティカルゲインの低下から、理論値より低い値となる場合はある)



Due to overestimation of sensor efficiency,
overestimation of sensitivity?

Another calibration method is required



② seismic noise reduction

- TM tilt adjustment

Reach **thermal noise level**

③ thermal noise reduction

- Cooling down

- Common / Differential decoupling
 - Injection into frequency actuator(common actuator in principle)
 - Depends on the consistency of two cavities
 - After sensor decoupling, also decouple coil-coil actuator
- Revise calibration method
 - Injection into frequency actuator(Transfer function is known)
 - Not relying on oplev.

- ミニマムサクセス(キャビティロック)は達成できた
 - 本当は修論でシリコンテストマス導入まで行きたかった
- M1でももう少しTOBAについて勉強しておくべきだった
- 既知の話や細かい部分に時間を割き過ぎてしまった
 - 並進カップリングについては下田さんの修論で確立されていた
 - 実験系の要素が多過ぎて本質を見極めるのが難しかった
 - やはり博士学生(orポスドク)について実験できるのが理想
- 学会や研究会にたくさん行った
 - 行くたびに友達が増えて面白い議論もできた
 - TOBAの宣伝
 - 集中して実験する時間を取るのも大事

- Introduction of TOBA
- Summary of my master's work
- Minor update plan for next suspension design
- Toward Phase-III, VIS & Cryogenic

Why local sensors are required?

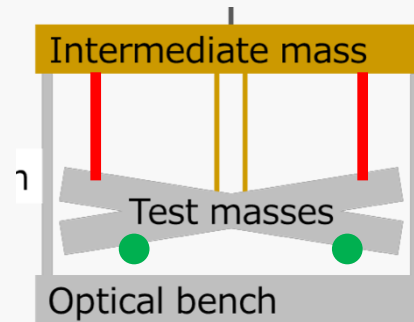
- Damping for lock acquisition→eddy current damping for TM currently
- Monitor of DC tilt? Maybe not sensitive enough

Severe space constraints

Desired modes

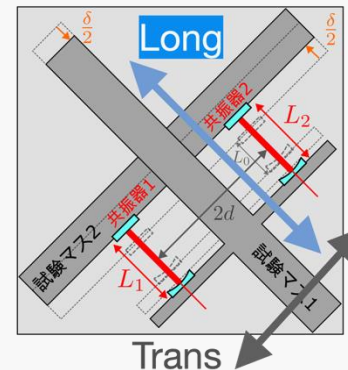
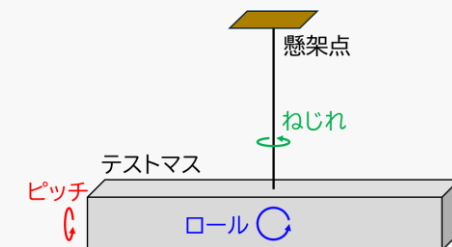
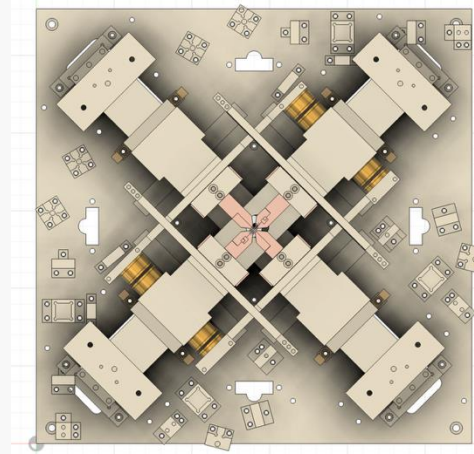
- Roll(Current oplev.s)
- Pitch(oplev. from IM?)
- Longitudinal/Trans(photo sensor from OB?)
- IM?

光学ベンチは結構ギチギチ...



New oplev.

Photo sensor



RMS is dominant in Roll & Long. Resonant mode
これらだけでもダンブしたい

Actuators for Roll & Pitch mode

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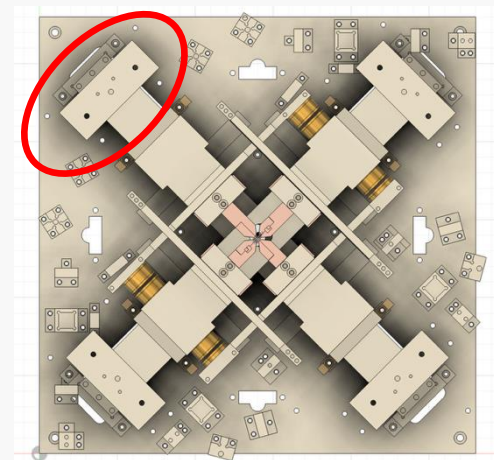
Roll & Pitch actuation for tilt adjustment is required

Slots for Roll & Pitch coil-coil actuators

DC actuation from OB

Yawの感度を汚さないか？

Also useful for damping in lock acquisition



Piezo driven setscrew

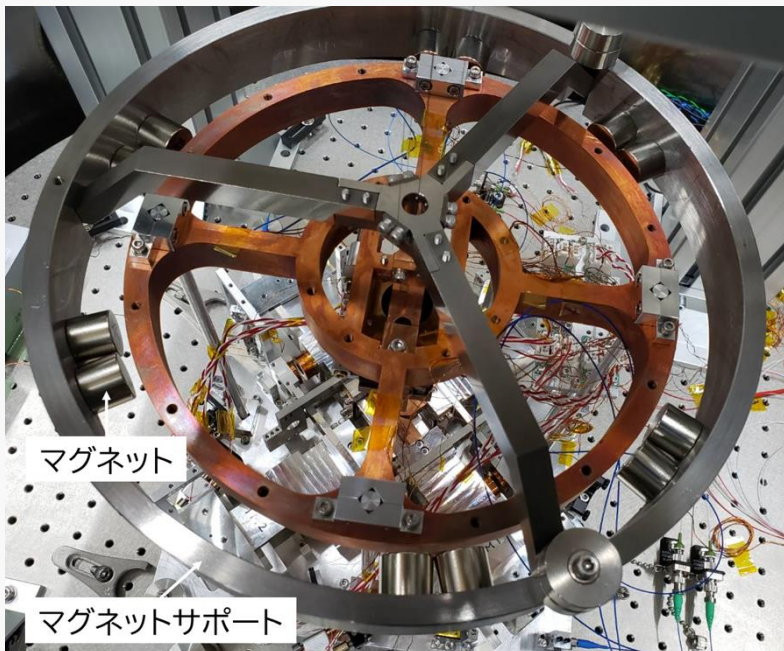
Purely adjust center of mass and tilt

Cryogenic compatible

TMにものはあまり載せたくない



JPE



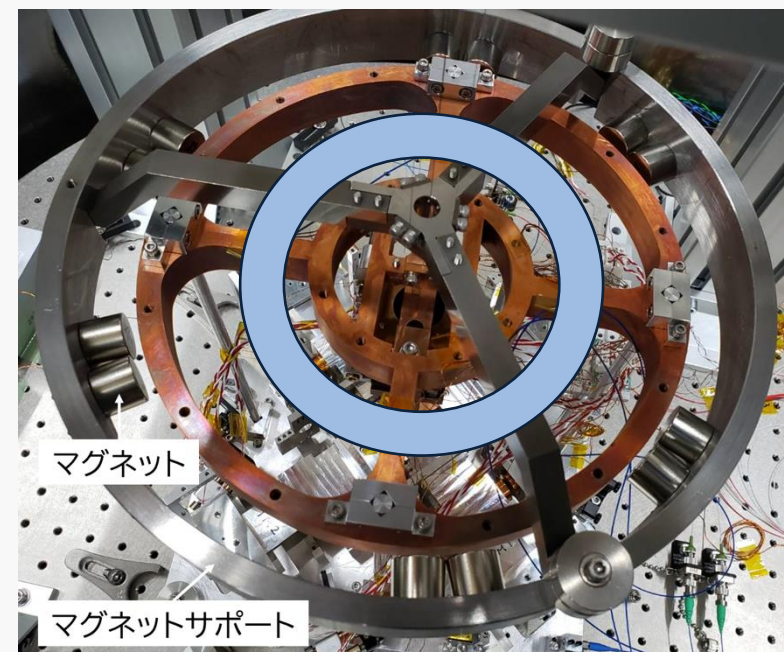
Current damping is overdamping for Yaw mode, but trans. Mode should be damped more

Damping from outside of IM tends to be overdamp for Yaw

➡ より内側でTrans. をダンプしたい

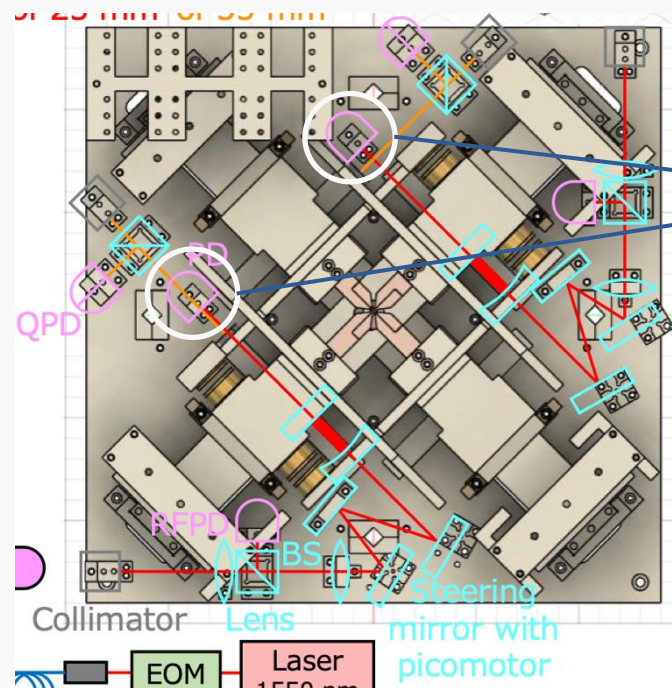
New damping magnets for Trans.

もっと内側の方が良い？



そもそも信号読み出しをどうするのか、
反射光でのPDHは低温でのRF読み出しが厳しい

➡ ~100kHzの変調で透過光PDH?
一旦は肩ロック継続、いずれにせよ透過光を使う



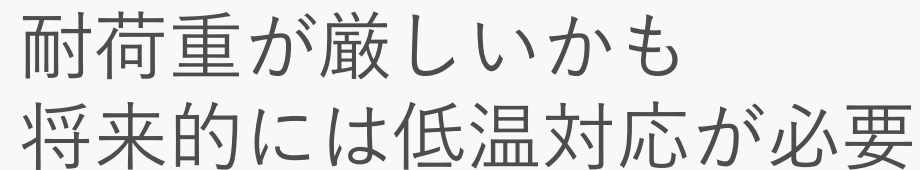
PDs for Transmitted light, no steering

どうにかステアリング鏡をねじ込みたい...

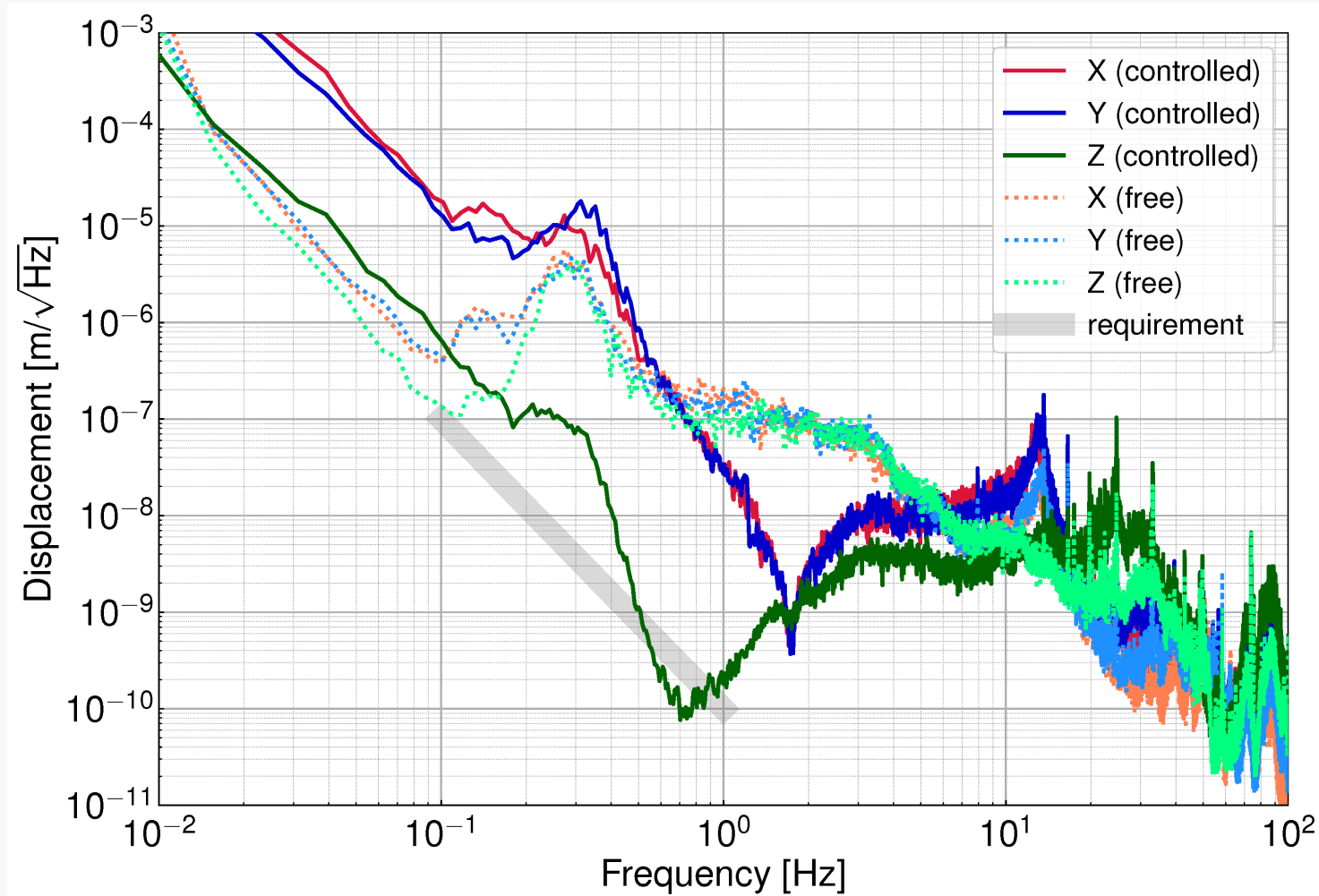
From Oshima-san's slide



 360° rotation stage is desirable



- Introduction of TOBA
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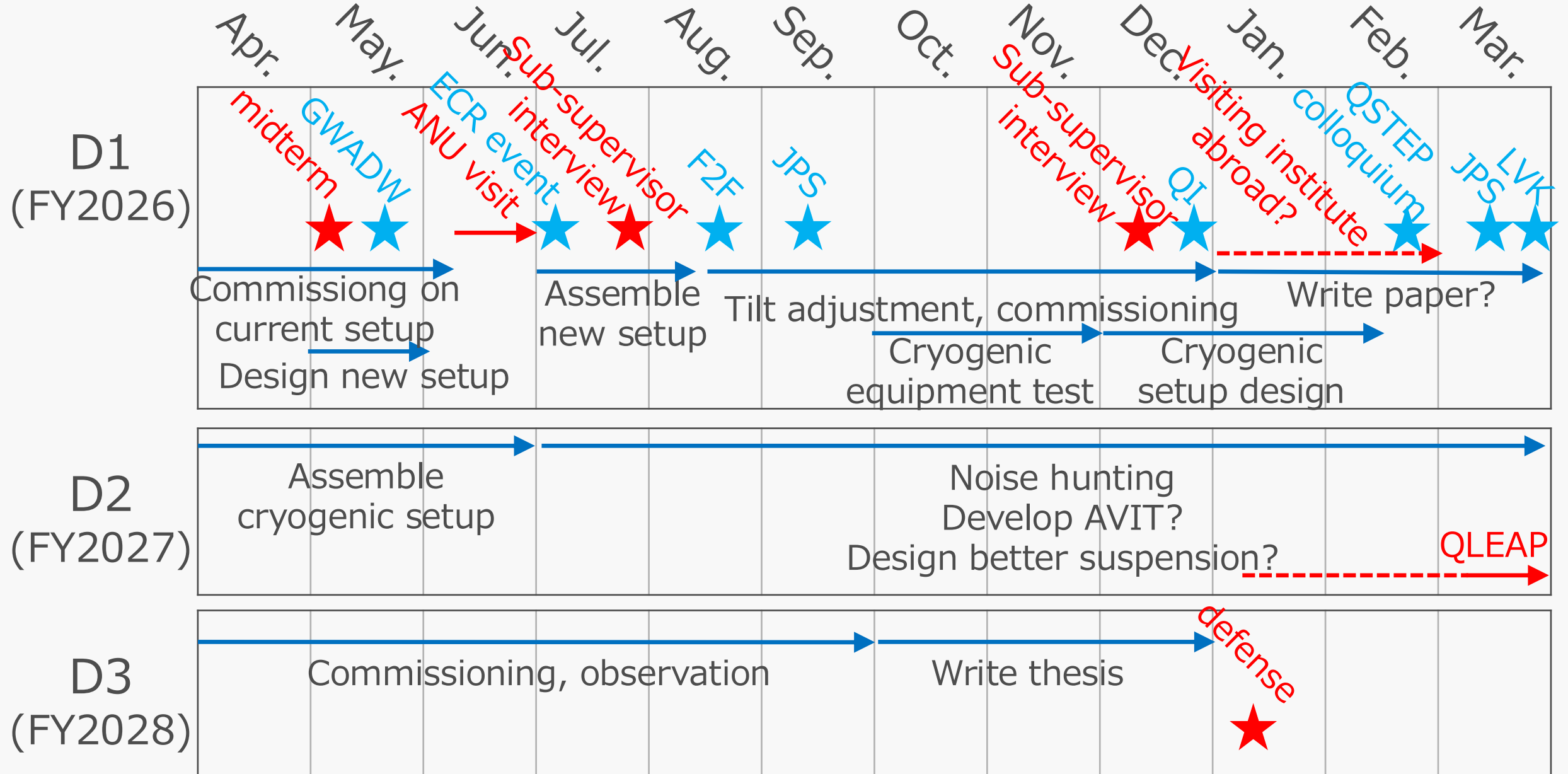


復旧はしておきたい

- Steering mirror
- Rotation stage
- PD circuit

Plans toward my master thesis

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- TOBA is a ground-based GW detector for low frequency using torsion pendulums
- Currently Phase-III TOBA is developed in Ando Lab.
- I am mainly developing the suspension and optics
- I managed to lock cavities
- Now I'm trying to reduce seismic noise and thermal noise by tilt adjustment and cooling down

Bonus slides

AVIT; Active Vibration Isolation Table

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- Actively isolate vibration on the suspension point(table)
- 3 DoF(x, y, z) control with geophones and piezo actuators

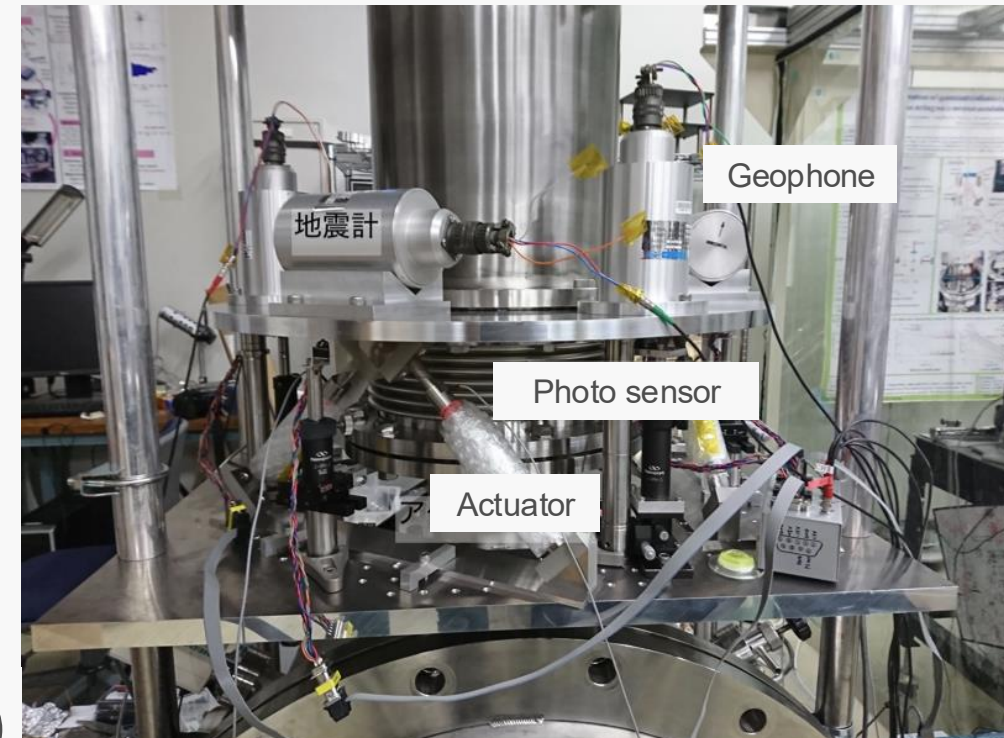
Geophone
× 6

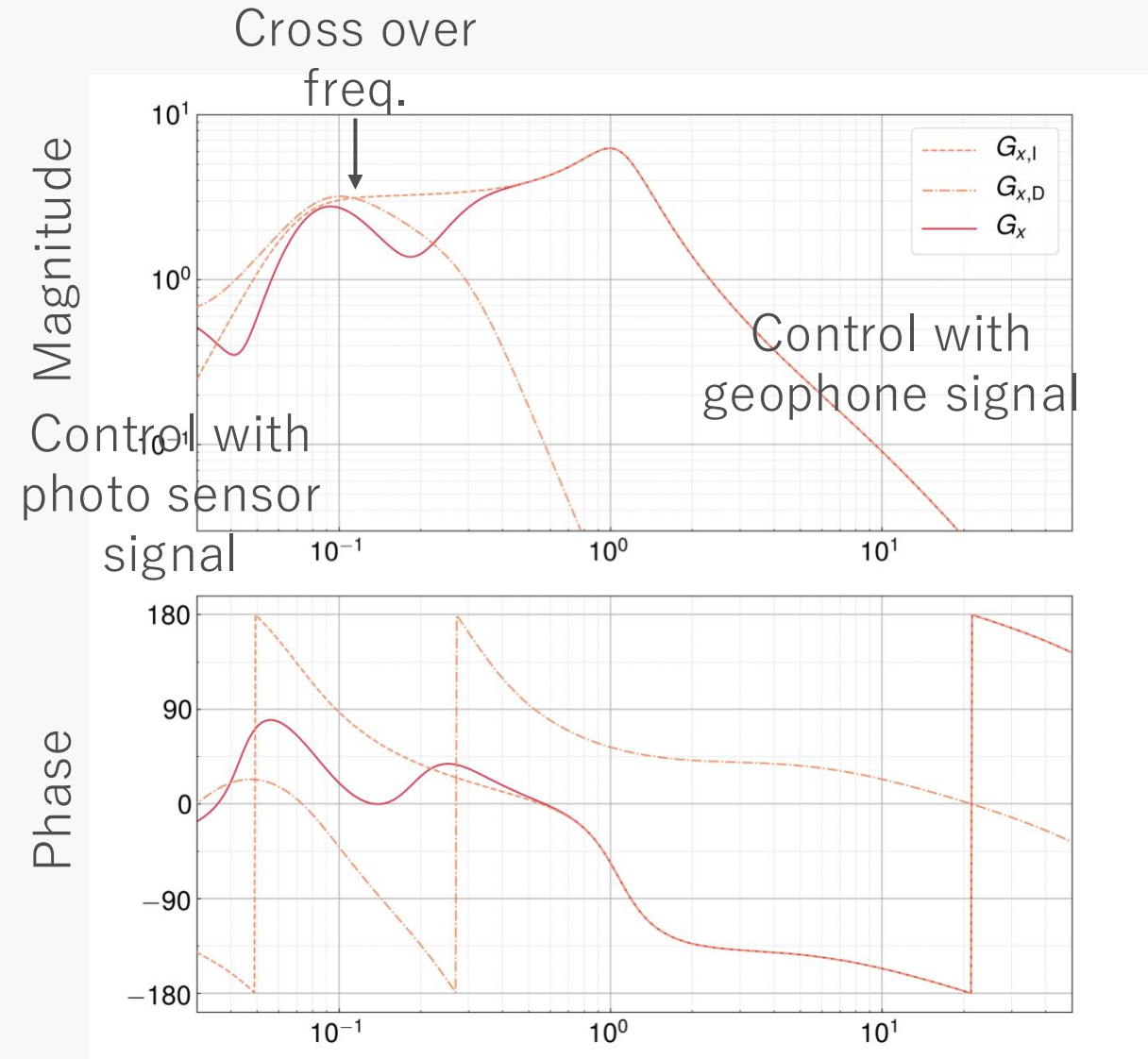


Table

Actuator × 6

(Photo sensor × 6)

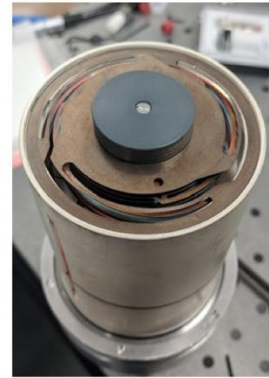




S. Takano(2019)

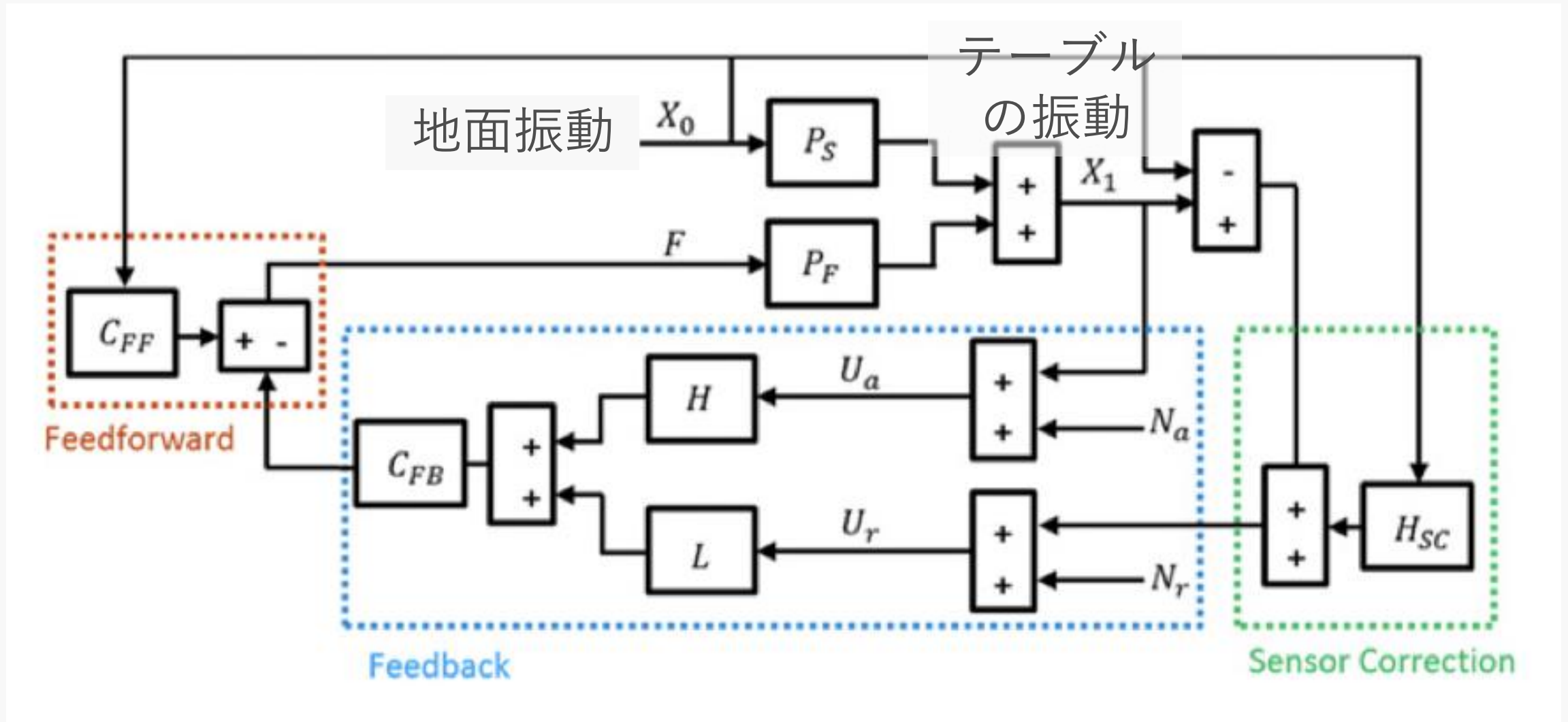
- Sensor noise of geophone increases in low freq.
 - High pass the signal from geophones
- Blend the signal from photo sensor(as relative sensor) to compensate DC drift
 - Just follow the ground motion in low freq.

- フレームの共振
 - 特注する？
- センサーの感度
 - L-4C自体の感度は十分だが、信号が小さくてアンプの雑音でリミット
 - 高性能なオペアンプが廃盤に？ KAGRAでも困っているらしい
 - LIGOではどうしているか調べてみる
 - 干渉計でL-4Cを読み出す？
- 傾きカップリング
 - 低周波で問題になる
 - 地面振動の傾き成分自体は小さいはず
→ テーブルの歪みが見えている？
 - テーブルを硬くする
 - 荷重との兼ね合い
 - 地面の地震計からフィードフォワードもあり？



L-4C with
HOQI readout

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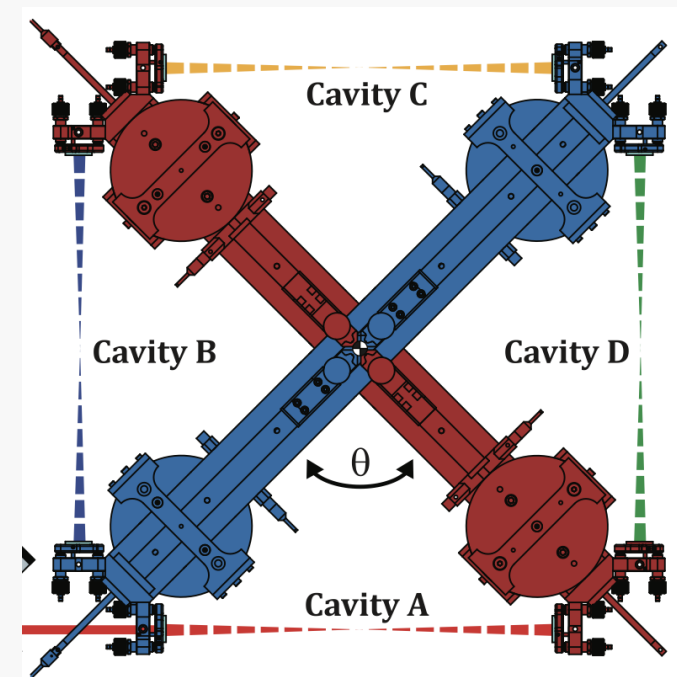
地面の地震計とのコヒーレンスが十分なら、フィードフォワードだけでも良いかも？



フレームの共振もテーブルの歪みも気にしなくて良い

手が回ればやりたいが、人手が欲しい

- 今の干渉計構成は最適なのか
 - 羽根は仕方ない…? テストマスが対等でないのも気持ち悪い
 - TorPeDOはcavity4本、レーザー4つでPLL
 - TMのYaw, Trans., Long.相対位置とレーザー周波数の4自由度あるので、原理的にはレーザー1つで制御できるはず
 - TM表面を鏡として使いたい、並進カップリングとの兼ね合い
 - TorPeDOは大きいサスペンションで受動的に防振する思想
- もっと受動的に防振したい
 - Inverted PendulumをAVITの上に入れる?
 - ベローズにスペースはあるはず
 - Ipに関してほとんど何も知らない…
 - TorPeDOのIPを見てきます



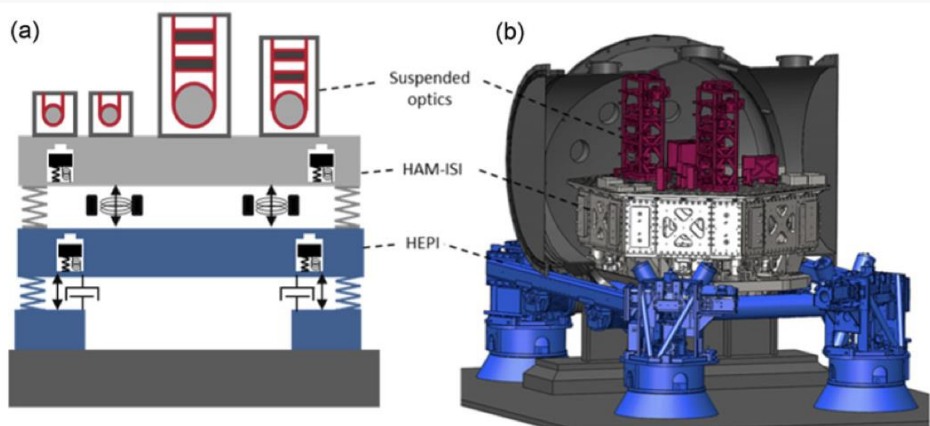
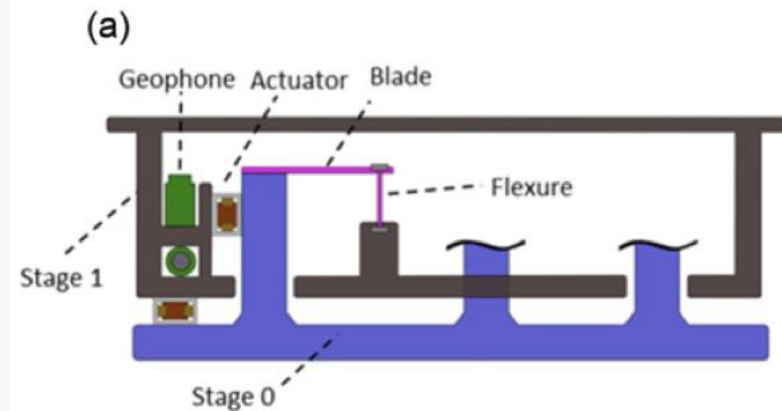
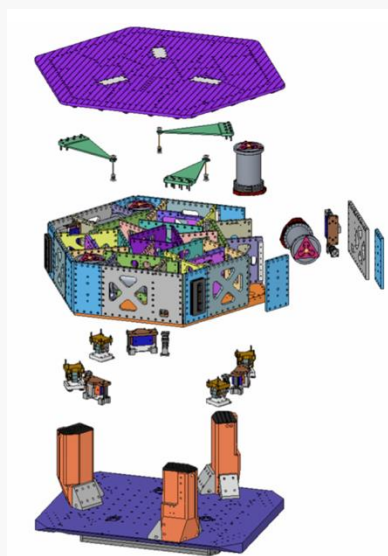
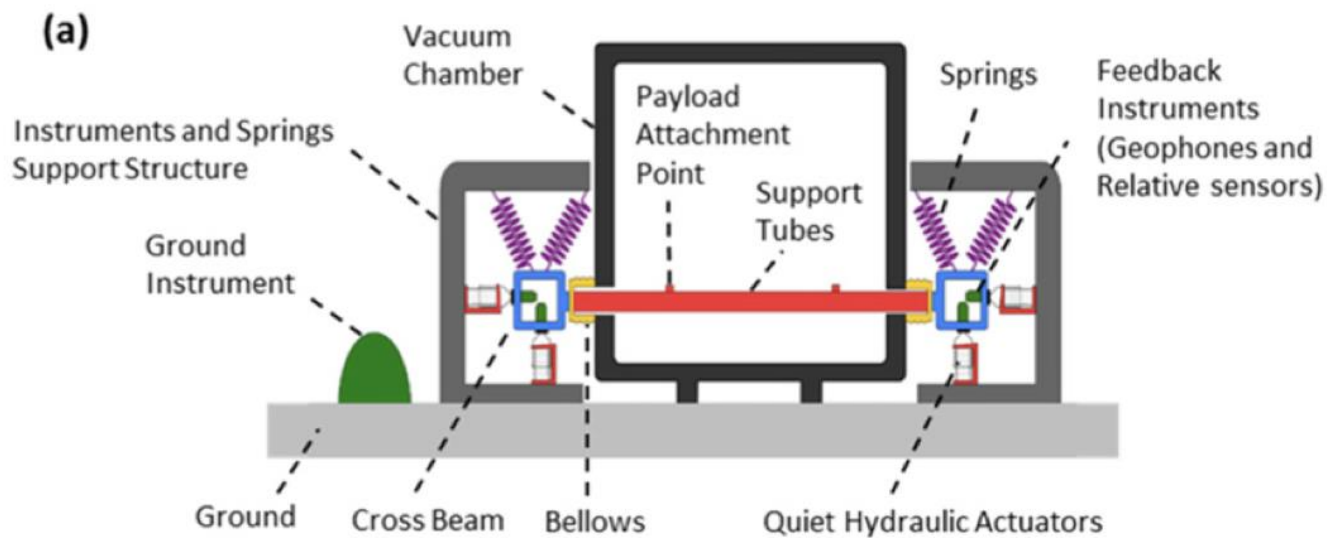
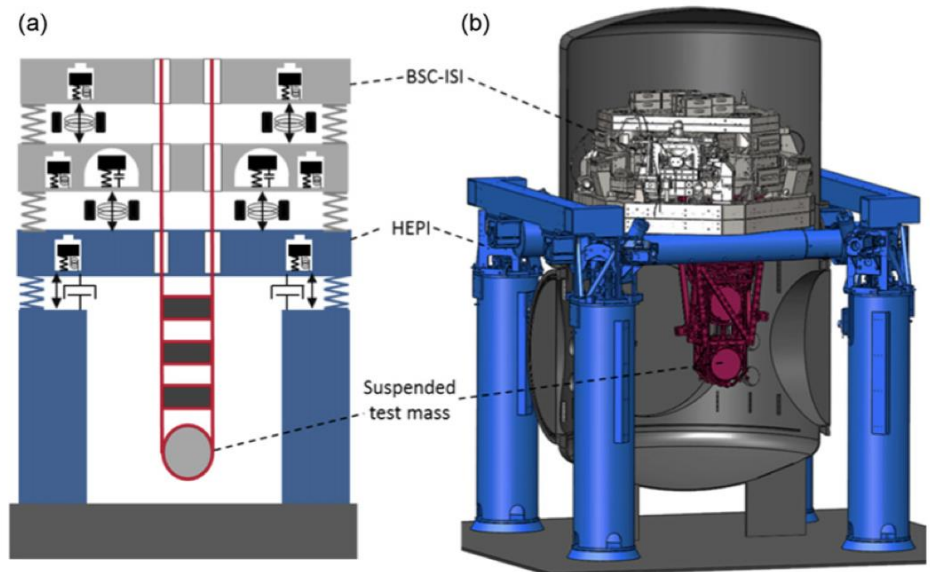
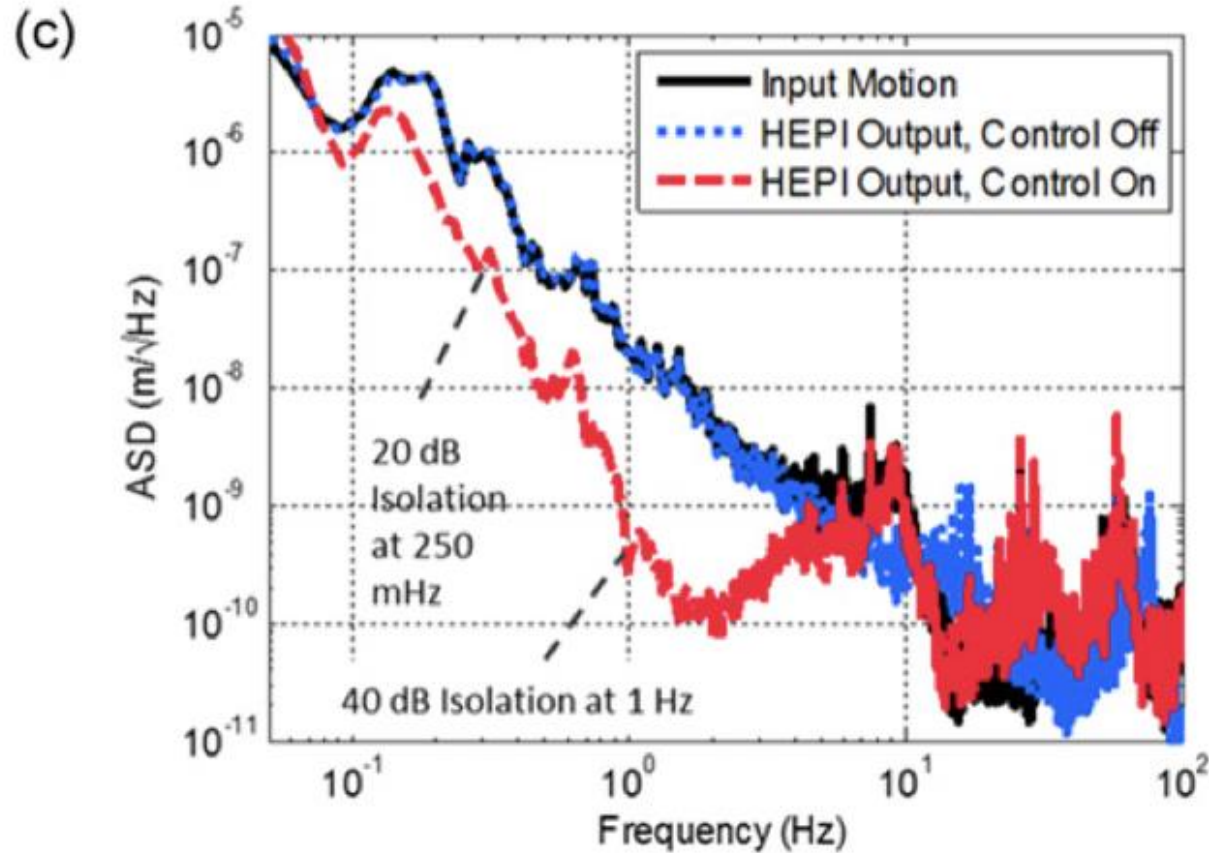
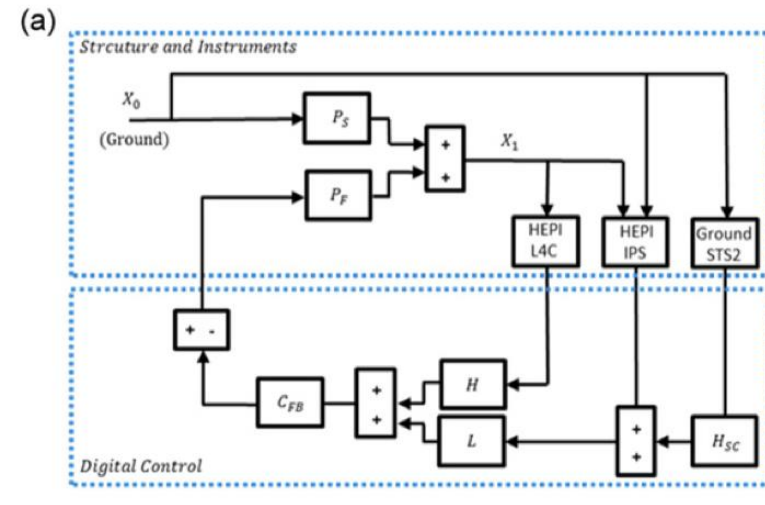


Figure 2. (a) Schematic and (b) CAD model of the isolation systems supporting the auxiliary optics in the HAM chambers.





HEPIは完全に能動的な防振装置
AVITでもまずはこれくらいを
目指したい



HEPIの防振性能

やっぱり地面の地震計で
センサーコレクション

地震計で6自由度制御しているわけではない

Self Introduction

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- **Tatsuya Sugioka/杉岡達哉**

- **Time line**

2002.2 :born in Hyogo

-2020.3 :Rakusei(洛星) high school in Kyoto

-2024.3 :Hokkaido Univ. applied physics

2024.4- :Ando Lab.

- **My Hobby : traveling, watching baseball, stargazing**

