# **Angular Signal Amplification with a Coupled Cavity** for Torsion-Bar Antenna

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#### Abstract

Torsion-Bar Antenna (TOBA) is a ground-based GW detector using a torsion pendulum. The resonant frequency of torsional motion is  $\sim 1$ mHz, therefore TOBA has good design sensitivity in low frequency, specifically  $10^{-19}$  / $\sqrt{Hz}$  at 0.1 Hz. TOBA can detect intermediate-mass black hole binary mergers, Newtonian noise, and so on. A prototype detector Phase-III TOBA with a 35 cm-scale pendulum is under development to demonstrate noise reduction. The target sensitivity is set to  $10^{-15}$  / $\sqrt{Hz}$  at 0.1 Hz.

To achieve our target sensitivity, we need to measure the pendulum rotation precisely. We propose a wavefront sensor with a coupled cavity (Coupled WFS) as an angular sensor for Phase-III TOBA. In our method, an auxiliary cavity is used to compensate Gouy phase of a main cavity and enhance the first-order TEM modes in the main cavity. The experimental demonstration was successfully performed in 2021. Here we show the principle and demonstration results of a Coupled WFS.



### 1. Introduction of TOBA

#### **TOBA: Torsion-Bar Antenna** [1]

- Ground-based GW detector for low frequency
- Aim to detect the torsional rotation Test masses of test masses suspended horizontally
- The resonant frequency of torsional motion is low ( $\sim 1 \text{ mHz}$ )
  - $\rightarrow$  Good sensitivity in low frequency even on the ground
- Scientific targets: intermediate-mass black hole binary mergers, Newtonian noise, Earthquake early warning, etc.

#### **Development plan** [2]

Phase-I (2009)	Phase-II (2015)	Phase-III (Now)	Final (Future)
Principle test		Technical demonstration	GW observation
10 <sup>-8</sup> /√Hz 20 cm Room	at 0.1 Hz (Established) bars temp.	10 <sup>-15</sup> /√Hz at 0.1 Hz (Target) 30 cm bars Cryo. Temp. (4 K)	10 <sup>-19</sup> /√Hz at 0.1 Hz (Target) 10 m bars Cryo. Temp. (4 K)

### **Design of the coupled cavity**

- Parameters are designed to allow Gouy phase compensation
- The main cavity is folded to monitor the transmitted light
- Mirrors are fixed to a spacer rigidly to stabilize the alignment

#### Length control of the coupled cavity

- PDH technique with two modulation frequencies
- Hierarchical control for the main cavity



#### $\Rightarrow$ See Satoru's poster (#A14) for more details about TOBA

[1] M. Ando et al., Phys. Rev. Lett. 105, 161101(2010) [2] T. Shimoda et al., International Journal of Modern Physics D 29, 1940003 (2020)

### 2. Principle of Coupled WFS

#### **Coupled WFS:** Wavefront sensor with a coupled cavity

An improved WFS proposed

to measure the torsional rotation of TOBA test masses accurately

- An auxiliary cavity can compensate Gouy phase of a main cavity
  - $\rightarrow$  Both TEM00 and TEM10 are resonant in the main cavity
  - $\rightarrow$  Angular signal is amplified by the finesse of the main cavity



#### **Comparison of angular sensors**

	Michelson interferometer	WFS	Coupled WFS
Shot noise (Requirement: 5×10 <sup>-16</sup> rad/√Hz)	0	No signal amplification	Signal amplification
Frequency noise	Asymmetry of two light paths	•	2
Beam jitter noise	Non-parallel of two mirrors		No amplification of beam jitter
Thermal noise	0	Narrow range measurement	Narrow range measurement
Linear range	0	0	Trade-off with signal amplification

- + 0.0 + 100 20 80 Time [sec]
  - Cavities were successfully locked TEM00 and TEM10 simultaneously
- Signal amplification was successfully demonstrated by three times

### 5. Summary & Future plans

- We are developing TOBA to detect GW in low frequency
- We propose Coupled WFS as an angular sensor for TOBA
- We demonstrated angular signal amplification and locking scheme of Coupled WFS
- We plan to suspend the test mass to stabilize the cavity lock

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