# DANCE: Dark matter Axion search with riNg Cavity Experiment



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### **Dark Matter Mystery**

- Suggested in 1930s from galaxy rotation curves
- Accounts for ~80% of all the matter of the universe
- The nature remains mystery



drives an acceleration of the expansion of the universe

#### **Dark Matter Models**

- ~90 orders of magnitude
- Searches focused on WIMPs, but not detected yet
- Motivates new searches for other candidates



### Ultralight DM with Interferometers

- Bosonic ultralight field (<~1 eV) are well-motivated from cosmology
- Behaves as classical waves

$$f = 242 \text{ Hz} \left( \frac{m_{\text{DM}}}{10^{-12} \text{ eV}} \right)$$

 Laser interferometers are sensitive to such oscillating changes





#### Various Searches Possible



### **Axion Dark Matter**

 Many experiments to search for ALPs through axion-photon coupling, especially by using magnetic fields (but ours don't)



### **Polarization Modulation from Axions**

- Axion-photon coupling  $(\frac{g_{a\gamma}}{4}aF_{\mu\nu}\tilde{F}^{\mu\nu})$  gives different phase velocity between left-handed and right-handed circular polarizations
  - $c_{\rm L/R} = \sqrt{1 \pm \frac{g_{a\gamma}a_0m_a}{k}} \sin(m_a t + \delta_{\tau})$ coupling constant axion field
- Linear polarization will be modulated p-pol sidebands will be generated from s-pol
- Search can be done without magnetic field



## Optical Cavity to Amplify the Signal

- Polarization rotation is small for short optical path

  Laser
- Optical cavities can increase the optical path, but the polarization is flipped by mirror reflections

![](_page_7_Figure_3.jpeg)

• Bow-tie cavity can amplify the rotation

![](_page_7_Figure_5.jpeg)

#### **Estimated Reach**

\* Shot noise limited, Better than CAST below 10<sup>-10</sup> eV 1-year observation  $10^{-6}$  $|g_{a\gamma}|$  (GeV<sup>-1</sup> **PVLAS 2016 ALPS-I 2010**  $10^{-7}$ **OSQAR 2015**  $10^{-8}$ **SHAFT 2021** umico 2008 ABRA 10-cm 2021  $10^{-9}$ **KAGRA** trans 2024  $10^{-10}$ coupling **ĀLPS-II**  $10^{-11}$ **IAXO** 10-12 Ation NGC1275 2020 10-13 **KAGRA** axion-photon  $10^{-14}$ **LIGO** DANCE 10<sup>-15</sup> ABRACADABRA  $10^{-16}$ ADMX 2010+2018  $10^{-17}$  $10^{-18}$  $10^{-17} 10^{-16} 10^{-15} 10^{-14} 10^{-13} 10^{-12} 10^{-11} 10^{-10} 10^{-9} 10^{-8} 10^{-7} 10^{-6} 10^{-5} 10^{-4} 10^{-3} 10^{-2} 10^{-1} 10^{0} 10^{-10} 10^$ axion mass  $m_a$  (eV)

### **Cosmic Birefringence and DANCE**

- Same principle
- Two-axion model can explain both cosmic birefringence and dark matter in the mass range  $g_{\mu\nu}$ [Generation of DANCE and  $g_{\mu\nu}$ ]
- Dynamical Dark Energy can also be explained by axions

![](_page_9_Figure_4.jpeg)

### Status of DANCE

- First demonstration in 2021 10<sup>-1</sup>
  Y. Oshima+, 10<sup>-2</sup>
  PRD 108, 072005 (2023)
- 4 orders of magnitude improvement over the past years

![](_page_10_Picture_3.jpeg)

![](_page_10_Figure_4.jpeg)

#### **First Data Analysis Results**

- Used 24-hour data from 12-day run in May 2021
- 551 candidates found from initial analysis

H. Nakatsuka+, PRD **108**, 092010 (2023)

- Veto analysis
  - Consistency veto (Frequency should be the same for different set of 24-hour data)
  - Q-factor veto (DM signal must have Q of 10<sup>6</sup>)
  - Remaining 7 candidates (all multiples of ~40 Hz) are also found in laser frequency control, and thus rejected
- Placed upper limits

![](_page_11_Figure_9.jpeg)

#### Simultaneous Resonance

 Carrier pol and sideband pol needs to be enhanced simultaneously for improving the sensitivity

![](_page_12_Figure_2.jpeg)

### **Cavity Birefringence Tuning**

- Near 45 deg incidence on cavity mirrors create reflection phase difference, which leads to nonsimultaneous resonance
- Reflection phase can be tuned by tuning laser wavelength
   H. Takidera+, arXiv:2505.06770 Resonant peak (1063.5nm)

![](_page_13_Figure_3.jpeg)

### Axion DM Search with KAGRA

- Linear cavities can be sensitive when the round-trip time equals odd-multiples of axion oscillation period
- Polarization optics installed in KAGRA
- First data expected in 2025

![](_page_14_Figure_4.jpeg)

### Summary

- Laser interferometers open up new possibilities for dark matter search
- Axion-like particles can be searched with DANCE to search for polarization modulation of light with optical ring cavity (without magnets!)
- Prototype experiment DANCE Act-1 is underway
- First result from 24-hour data released
  Y. Oshima+, PRD 108, 072005 (2023)
- Further sensitivity improvement underway with wavelength tunable laser
   H. Takidera+, <u>arXiv:2505.06770</u>