

Search for New Physics in Electronic Recoil Data from XENONnT

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 - a dual-phase LXe TPC
- XENONnT
 - Radon removal system (RRS)
 - Background model B_0
 - 90% C.L. upper limit
- Summary

Search for New Physics in Electronic Recoil Data from XENONnT

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(XENON Collaboration)^{††}

The primary science goal of XENONnT:
Search for weakly interacting massive particles (WIMPs)

To reduce radon-induced backgrounds in liquid xenon detectors

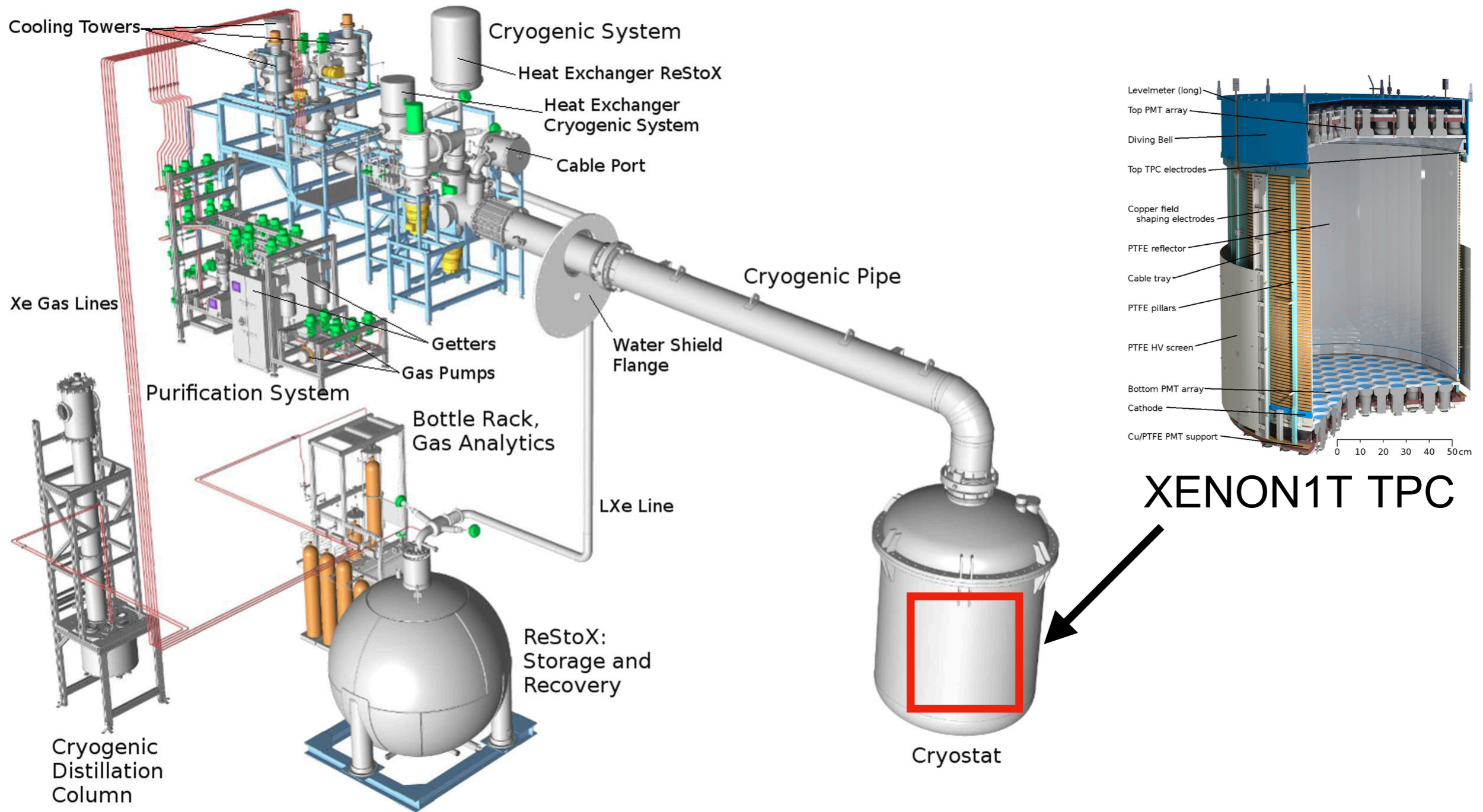


A high-flow radon removal system based on cryogenic

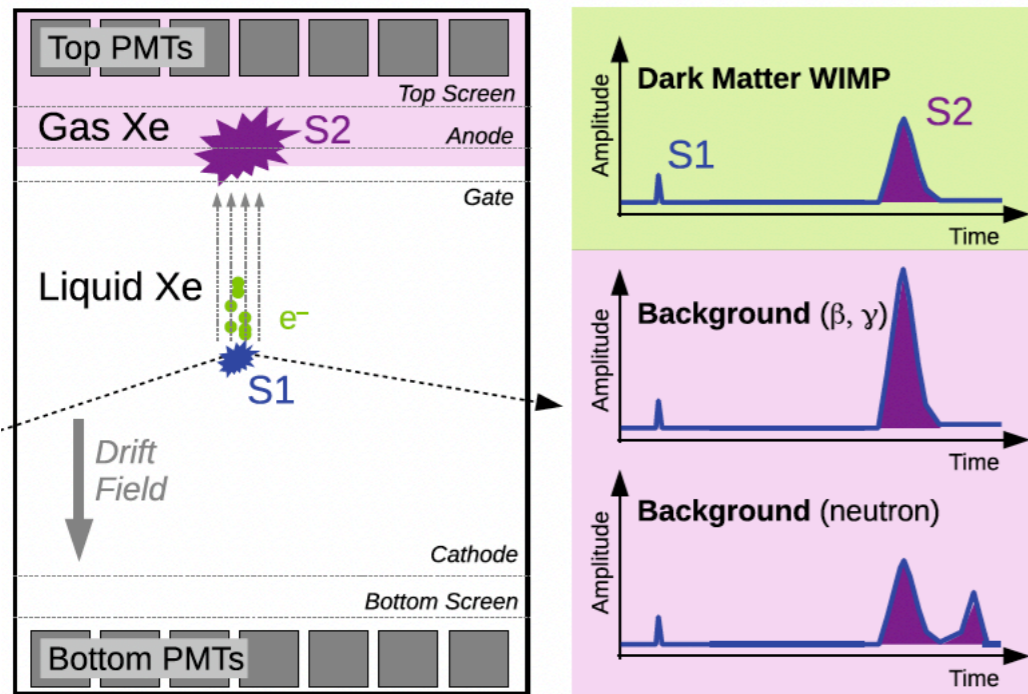


data analysis

- No excess above background
- Set stringent new limits
 - solar axions
 - an enhanced neutrino magnetic moment
 - boson dark matter



Working principle of a dual-phase LXe TPC



Particles scatter off xenon nuclei (WIMPs or neutrons)

→ nuclear recoils

Particles interact with atomic electrons (γ rays and β electrons)

→ electronic recoils

The recoils excite and ionize the LXe

<https://link.springer.com/article/10.1140/epjc/s10052-017-5326-3>

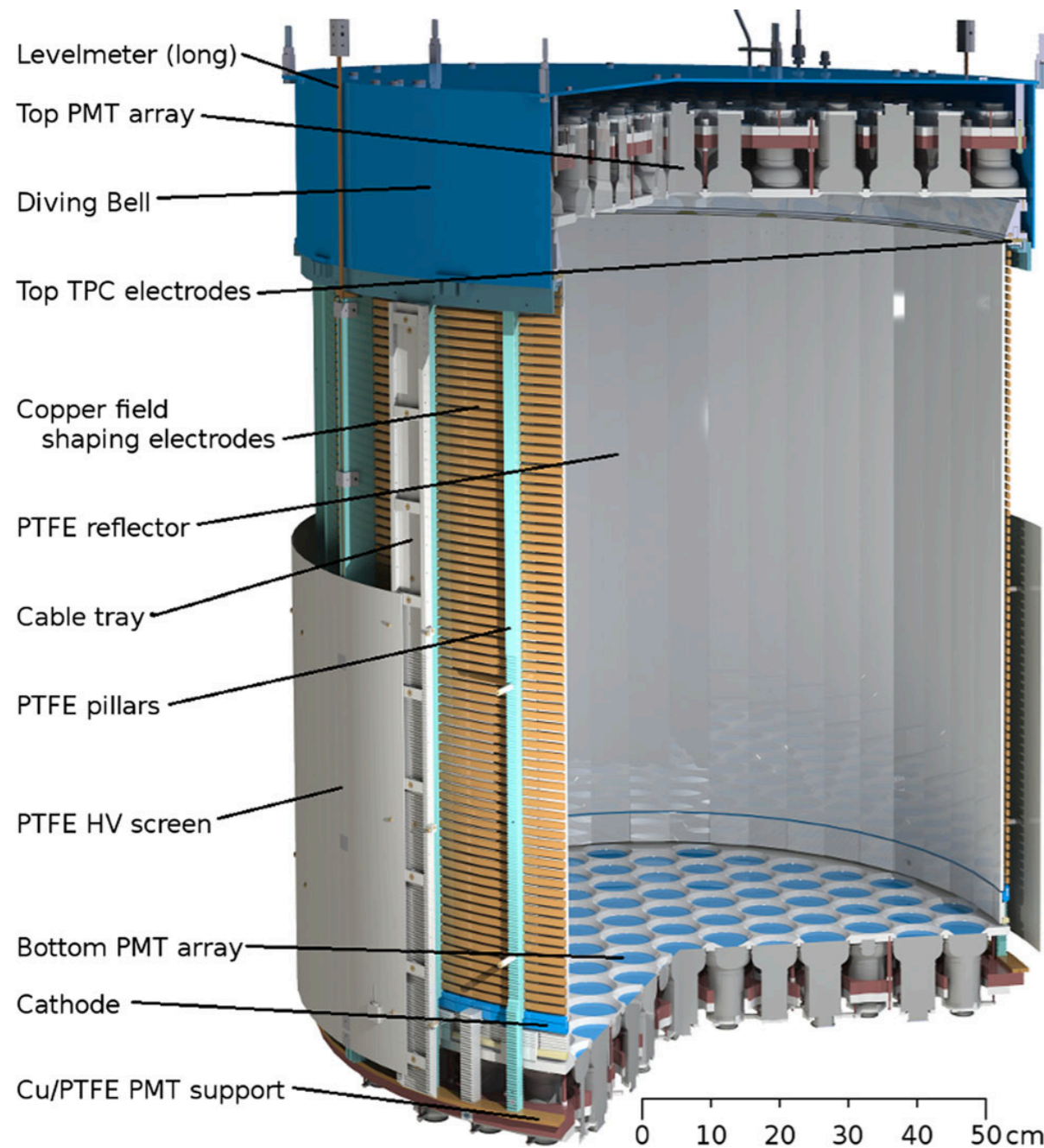
S1: scintillation photons

S2: secondary scintillation photons

The S1-S2 time difference

→ The ratio S2/S1 can be employed for electronic recoil background rejection

Illustration of a dual-phase LXe TPC

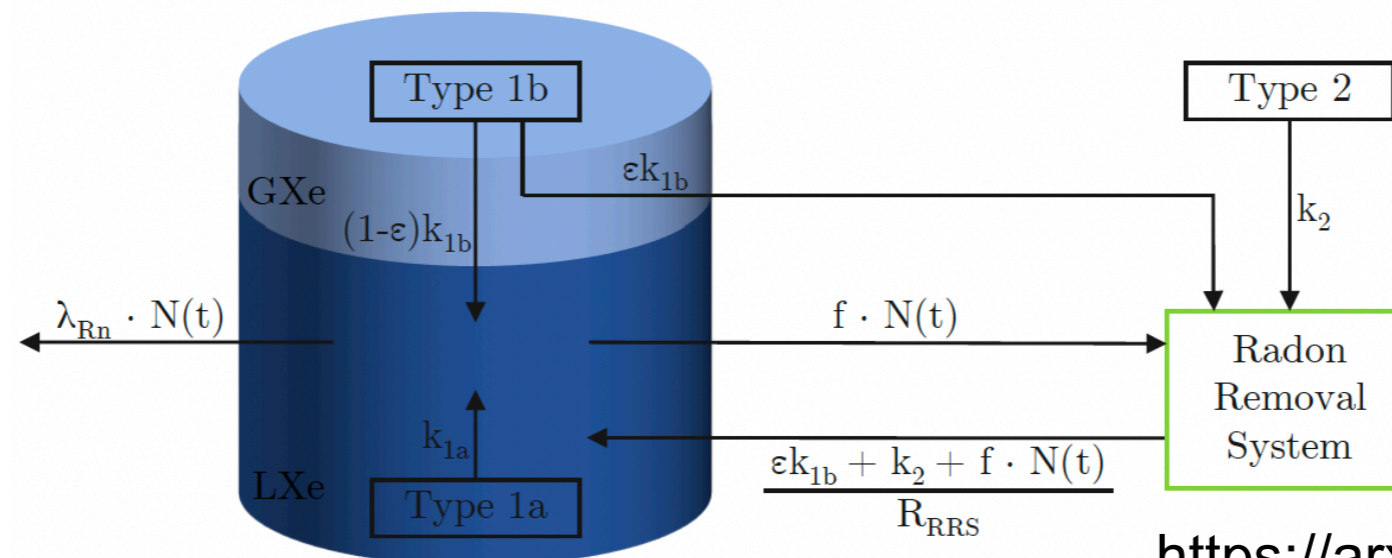


- an active LXe target of 2.0t
- built from materials selected for their low radioactivity
- enclosed by 24 interlocking and light-tight PTFE (polytetrafluoroethylene) panels

<https://link.springer.com/article/10.1140/epjc/s10052-017-5326-3>

Improvement from XENON1T

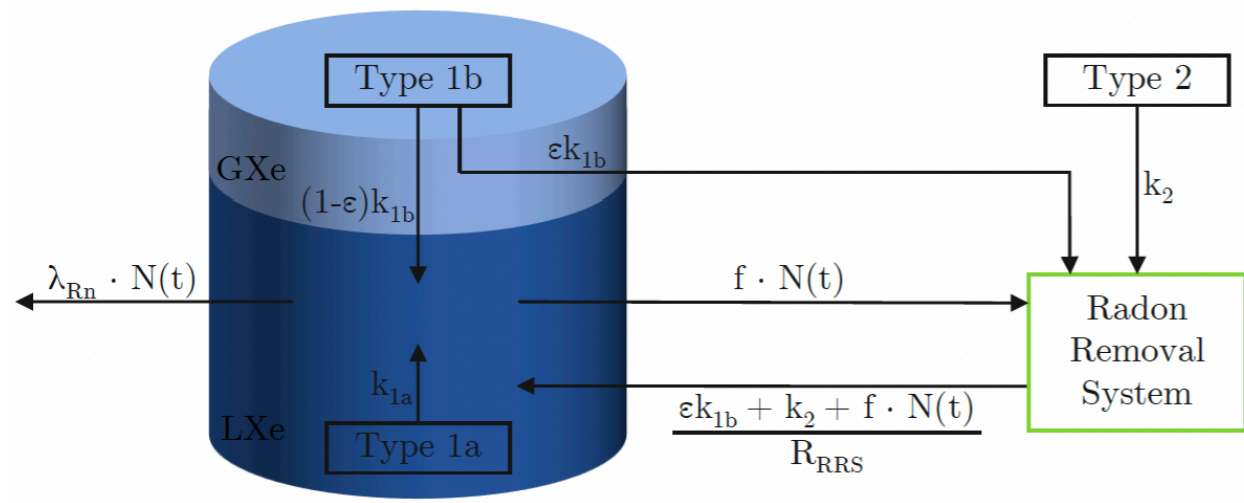
A high-flow radon removal system (RRS) based on cryogenic
 → further reduce background



<https://arxiv.org/abs/2205.11492>

Radon removal system (RRS)

Radon removal system (RRS)



<https://arxiv.org/abs/2205.11492>

The change in the number of radon particles

$$\frac{dN(t)}{dt} = k_{1a} + (1 - \epsilon)k_{1b} - \lambda_{Rn} \cdot N(t) - f \cdot N(t) + \frac{k_2 + \epsilon k_{1b} + f \cdot N(t)}{R_{RRS}}$$

the decay of ^{222}Rn

the effective radon particle flux

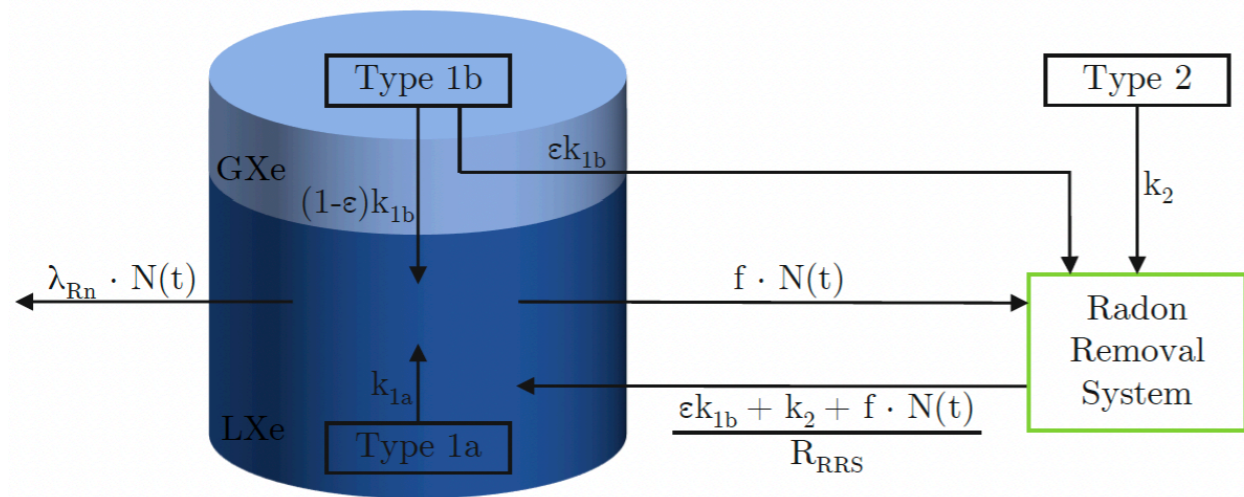
$$\lambda_{Rn} = 0.18[1/d]$$

$$f = \frac{F_{Xe}}{m_{Xe}} = \frac{1}{\tau_{ex}}$$

ϵ : extraction efficiency τ_{ex} : the detector's Lee volume exchange time

$$R_{RRS} \equiv \frac{c_F}{c_D}: \text{RRS's reduction factor}$$

Radon removal system (RRS)



<https://arxiv.org/abs/2205.11492>

Starting condition $N(t = 0) = N_0$ (N_0 : the number of radon atoms before starting the removal)

$$N(t) = \frac{K}{\Lambda} + \left(N_0 - \frac{K}{\Lambda} \right) \cdot e^{-\Lambda \cdot t}$$

$$K = k_{1a} + (1 - \epsilon)k_{1b} + \frac{k_2 + \epsilon k_{1b}}{R_{RRS}}$$

$$\Lambda = \lambda_{Rn} + f \cdot \left(1 - \frac{1}{R_{RRS}} \right)$$

In case a perfect RRS can fully remove radon

$$N_{\text{equi}} \stackrel{t \rightarrow \infty}{=} \frac{K}{\Lambda} = \frac{k_{1a} + (1 - \epsilon)k_{1b} + \frac{k_2 + \epsilon k_{1b}}{R_{\text{RRS}}}}{\lambda_{\text{Rn}} + f \cdot \left(1 - \frac{1}{R_{\text{RRS}}}\right)} \stackrel{R_{\text{RRS}} \rightarrow \infty}{=} \frac{1}{\lambda_{\text{Rn}} + f} \cdot (k_{1a} + (1 - \epsilon)k_{1b})$$

$N_{\text{equi}}(R_{\text{RRS}}, f, \epsilon)$

The reduction inside the detector's LXe volume

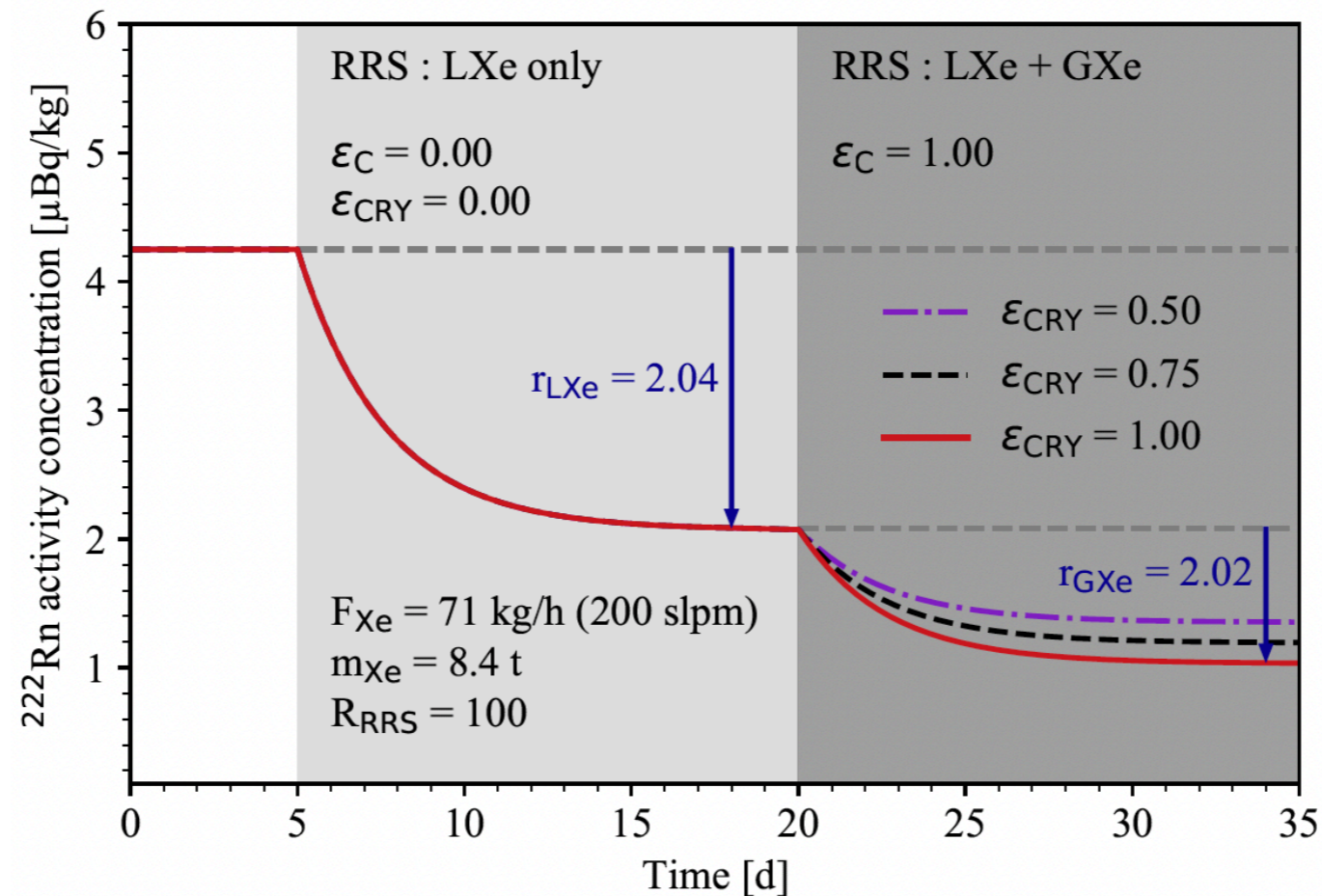
$$r(R_{\text{RRS}}, f, \epsilon) = \frac{N_{\text{equi}}(R_{\text{RRS}} = 1, f, \epsilon)}{N_{\text{equi}}(R_{\text{RRS}}, f, \epsilon)} \quad r(R_{\text{RRS}} \rightarrow \infty, f, \epsilon) = \frac{\lambda_{\text{Rn}} + f}{\lambda_{\text{Rn}}} \cdot \frac{k_{\text{tot}}}{k_{1a} + (1 - \epsilon)k_{1b}}$$

The combined LXe and GXe modes

The individual reduction capabilities

$$r_{\text{LXe}} = r(R_{\text{RRS}} \rightarrow \infty, f, \epsilon = 0) = \frac{\lambda_{\text{Rn}} + f}{\lambda_{\text{Rn}}} \cdot \frac{k_{\text{tot}}}{k_{1a} + k_{1b}}$$

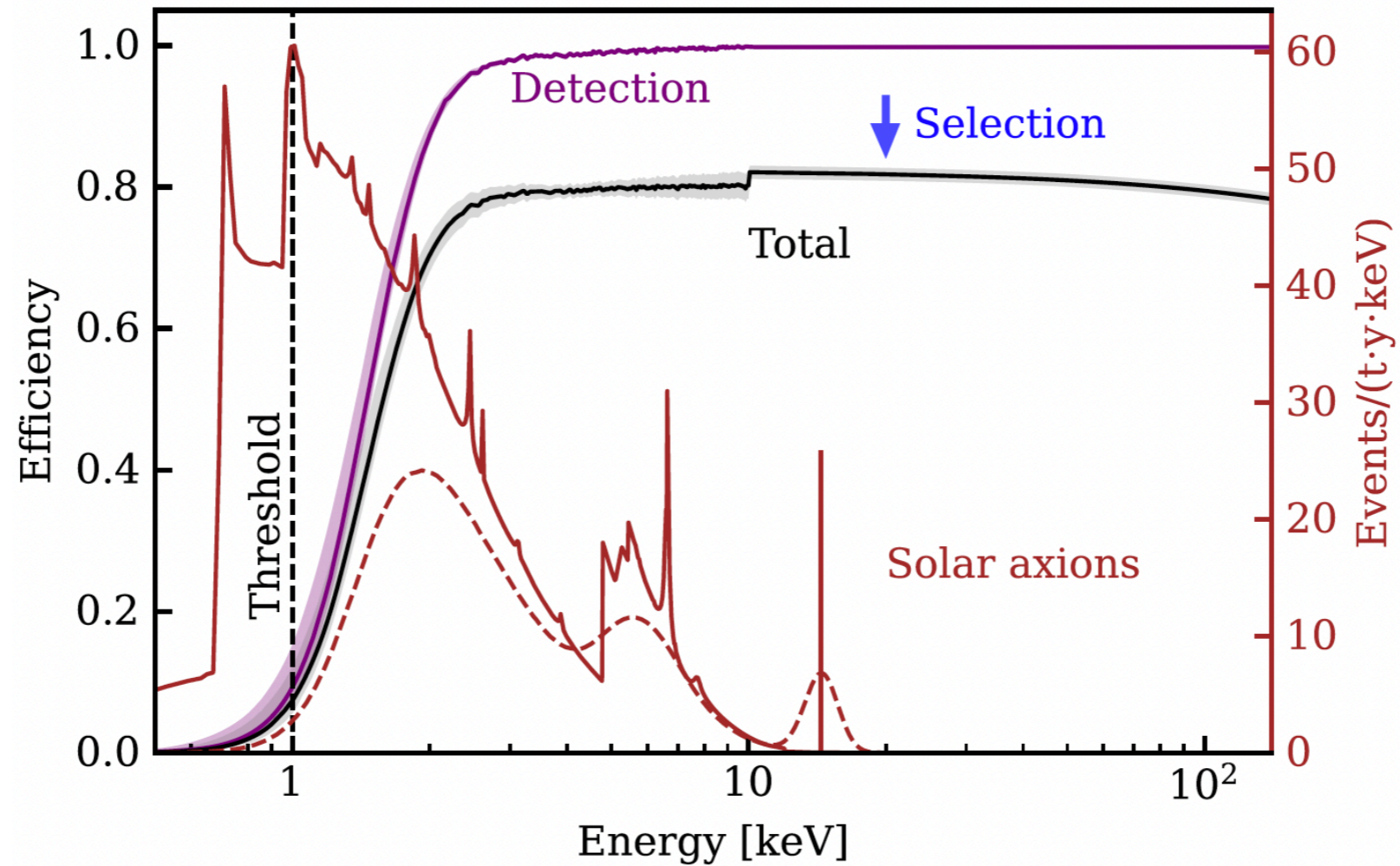
$$r_{\text{GXe}} = r(R_{\text{RRS}} \rightarrow \infty, f = 0, \epsilon) = \frac{k_{\text{tot}}}{k_{1a} + (1 - \epsilon)k_{1b}}$$



Expected radon reduction in XENONnT ($R_{\text{RRS}} \rightarrow \infty$)



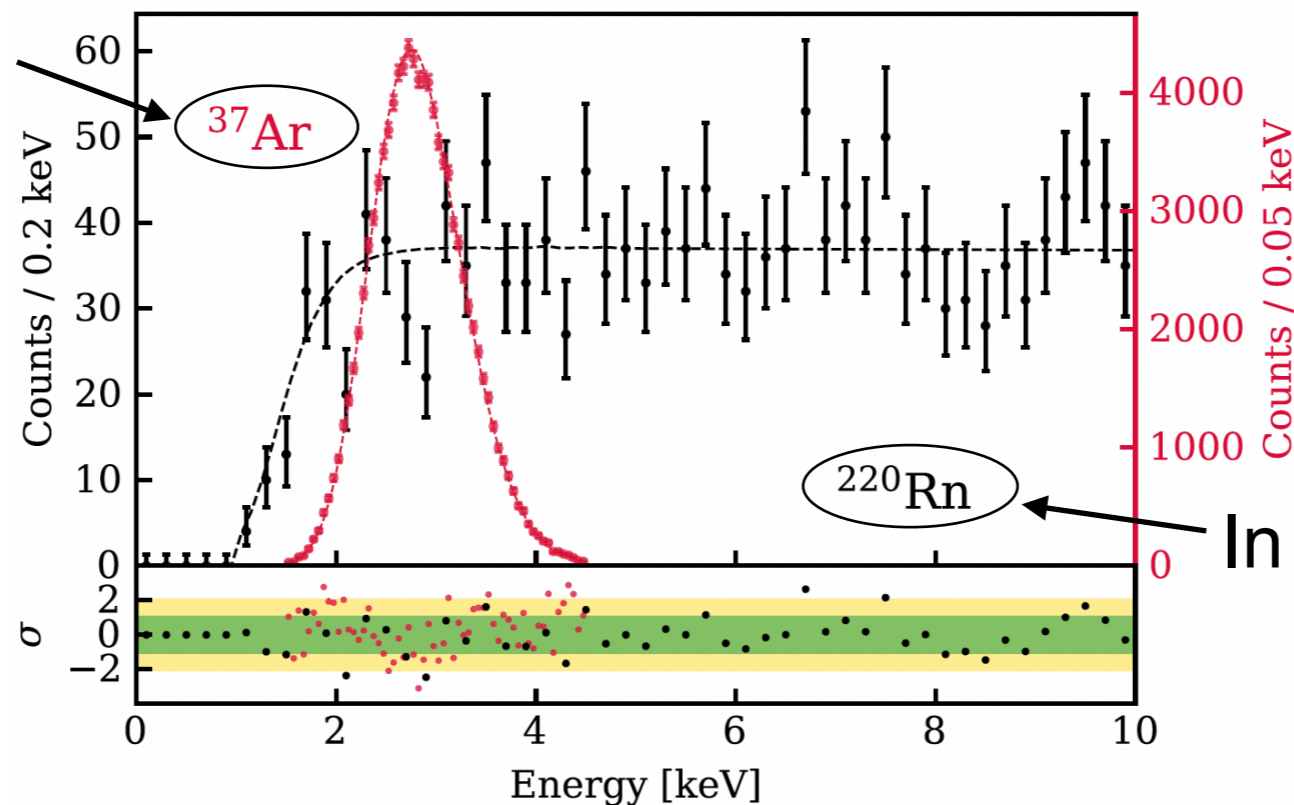
XENONnT can reach ^{222}Rn activity concentration of 1 $\mu\text{Bq/kg}$



Efficiencies in reconstructed energy and the solar axion signal

consider ER interactions only

In December 2021



In June 2021

^{220}Rn → the energy threshold, selection efficiency, energy reconstruction

^{37}Ar → validate the energy reconstruction and skew-Gaussian smearing model

Background model B_0

Consider 9 components in the background model B_0

| | Component | Constraint | Fit |
|---|--------------------------|-----------------|-----------------|
| The dominant background at low energy → | ^{214}Pb | (570, 1200) | 960 ± 120 |
| | ^{85}Kr | 90 ± 60 | 90 ± 60 |
| | Materials | 270 ± 50 | 270 ± 50 |
| The dominant component above 40 keV → | ^{136}Xe | 1560 ± 60 | 1550 ± 50 |
| | Solar neutrino | 300 ± 30 | 300 ± 30 |
| | ^{124}Xe | ... | 250 ± 30 |
| accidental coincidence → | AC | 0.70 ± 0.04 | 0.71 ± 0.03 |
| | ^{133}Xe | ... | 150 ± 60 |
| | $^{83\text{m}}\text{Kr}$ | ... | 80 ± 16 |

constraints between (39, 44) keV are excluded

Further the possibility of tritium as an explanation for the XENON1T excess
→ operated XENONnT in a different mode for 14.3 days after the SR0 data collected (SR0: from July 6, 2021 to November 10, 2021)



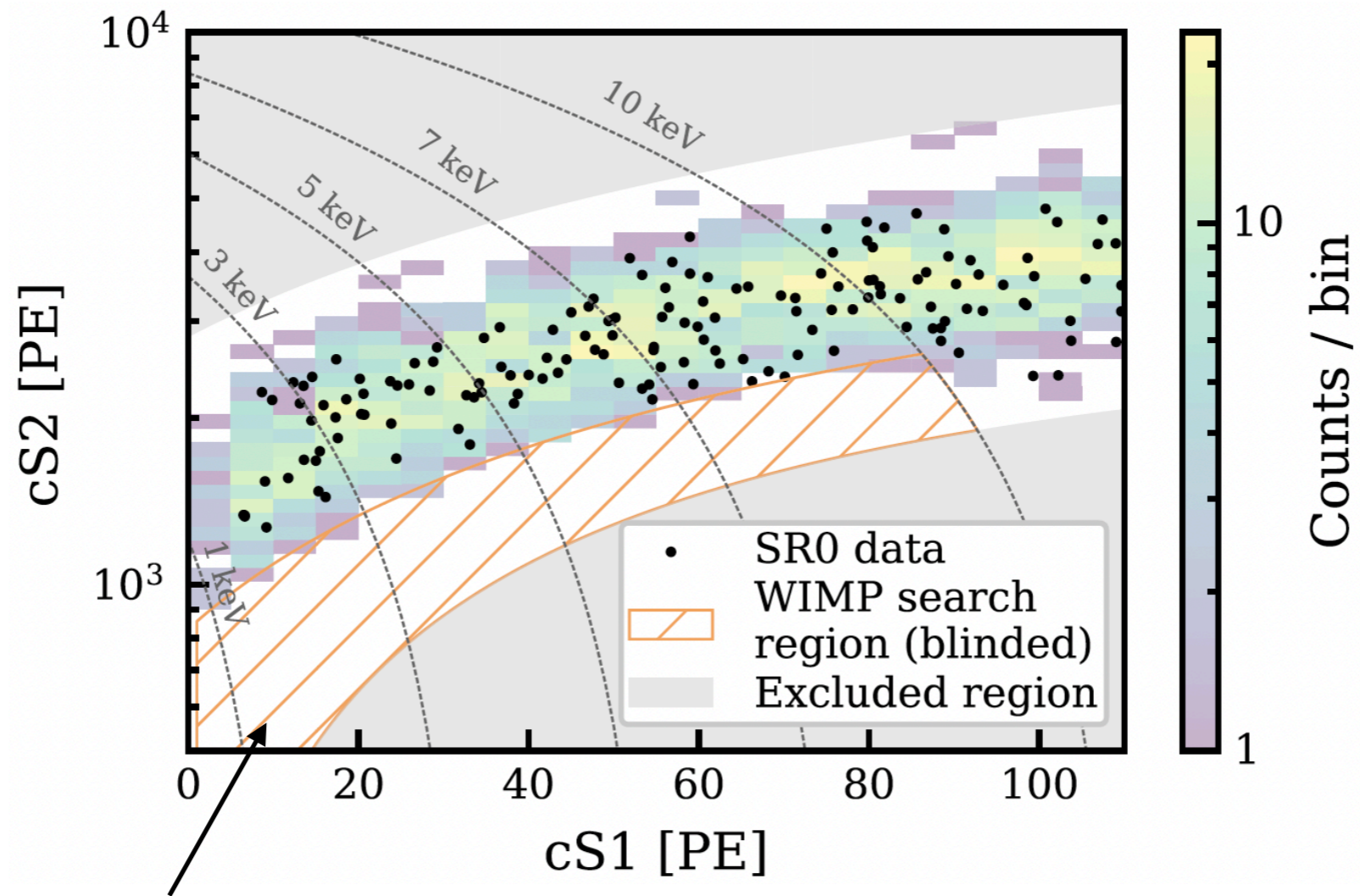
Tritium dataset showed no evidence for a tritium like excess



Tritium is not included in the background model !

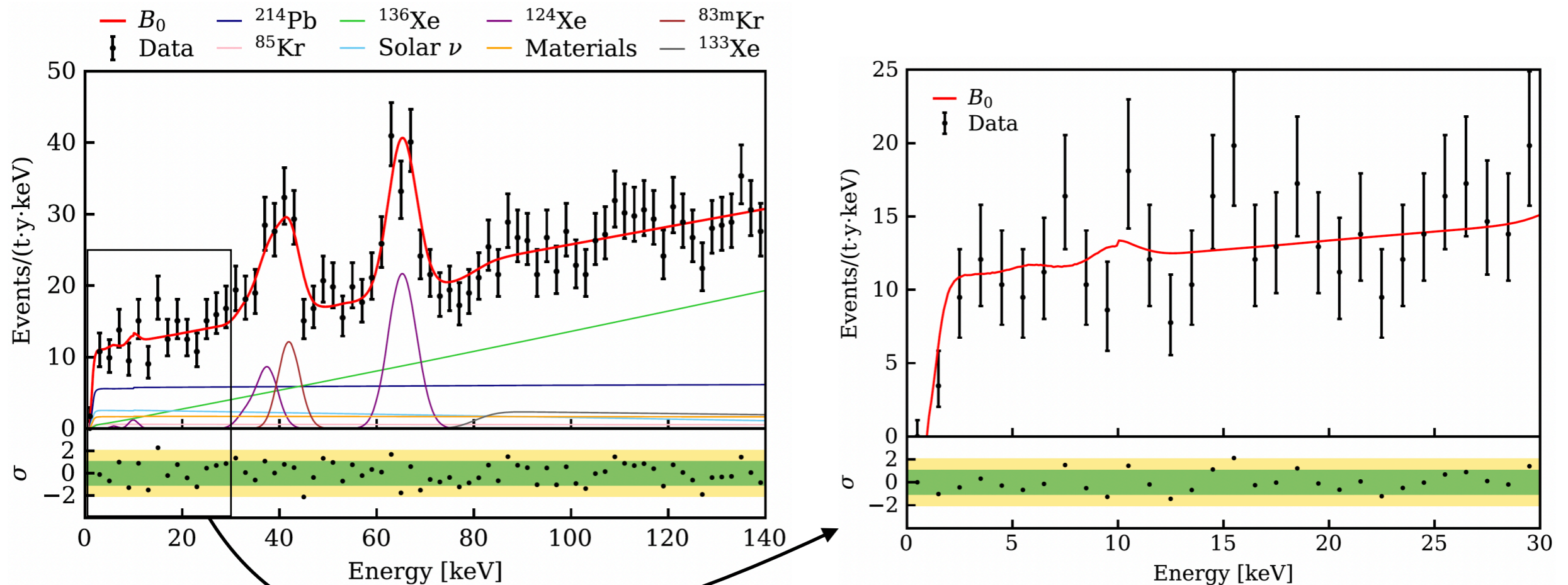


Tritium may be the cause of the excess observed in XENON1T



The WIMP search region is not used in this search

Fit to SR0 data using the B_0 model

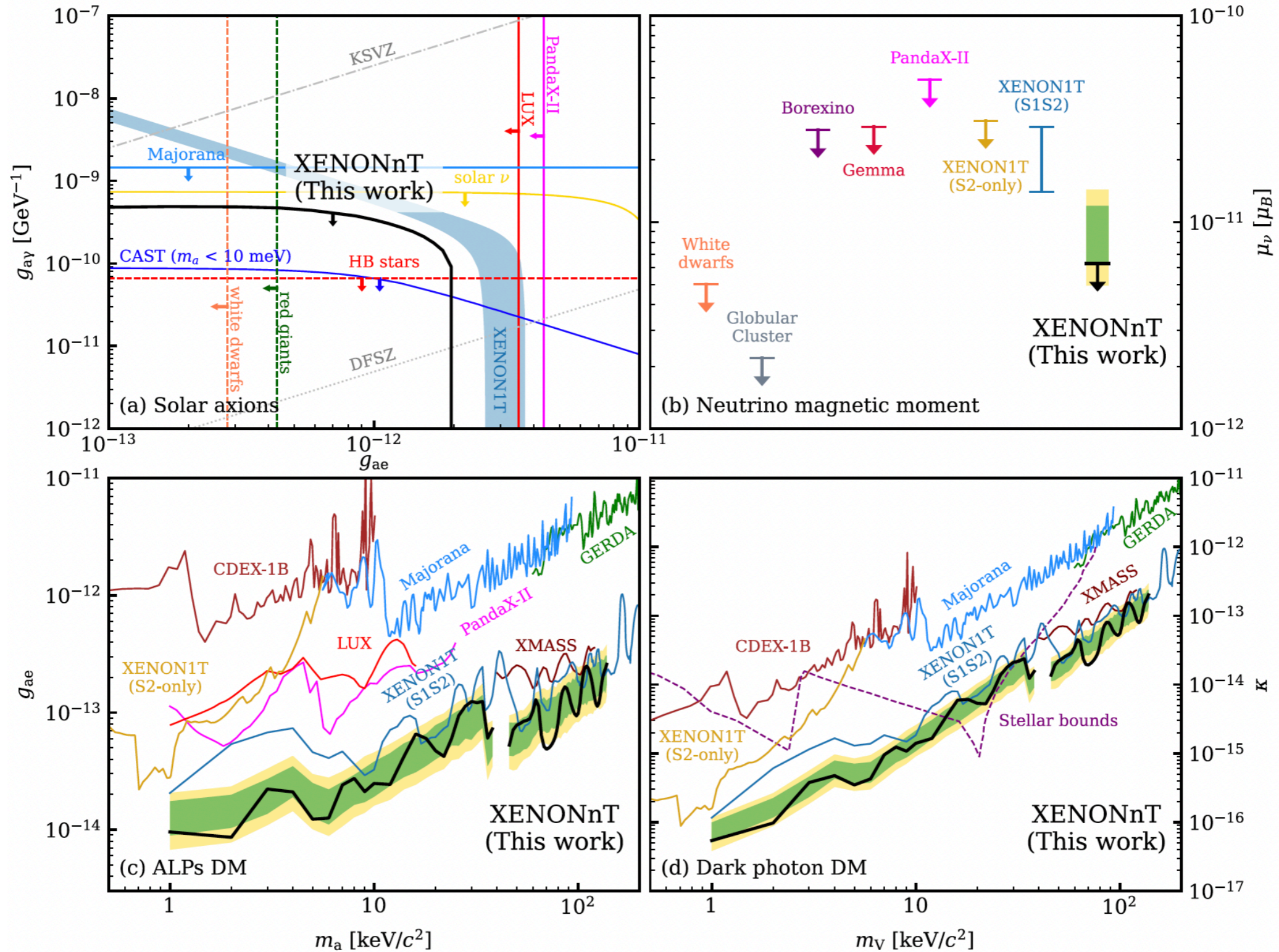


ER background rate within (1,30) keV \rightarrow $(15.8 \pm 1.3)\text{events}/(\text{ton}\times\text{year}\times\text{keV})$



The lowest background rate ever achieved !

90% C.L. upper limit



- The blind analysis shows no excess above the background.
→ Tritium may be the cause of the excess observed in XENON1T.
- The average ER background rate of (15.8 ± 1.3) events/(ton×year×keV) in the (1,30) keV energy region is the lowest ever achieved in a DM search experiment.