

Report on the Test of Lorentz Invariance

2017/11/17 ANDO GROUP SEMINAR

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Abstract

- Measuring Lorentz violation signal since Oct. 31st
- Data analysis in Mansouri-Sexl framework
 - ✓ I used a week-long data (Oct. 31st - Nov. 7th)
 - ✓ LV signal after Nov. 8th is very noisy (*why?*)
- Lorentz violation parameter:

$$\alpha + 1/2 = (-0.5 \pm 2.5) \times 10^{-12} \quad \left(\frac{\delta c}{c} \sim 6 \times 10^{-15} \right)$$

- ✓ the same order of magnitude as world record

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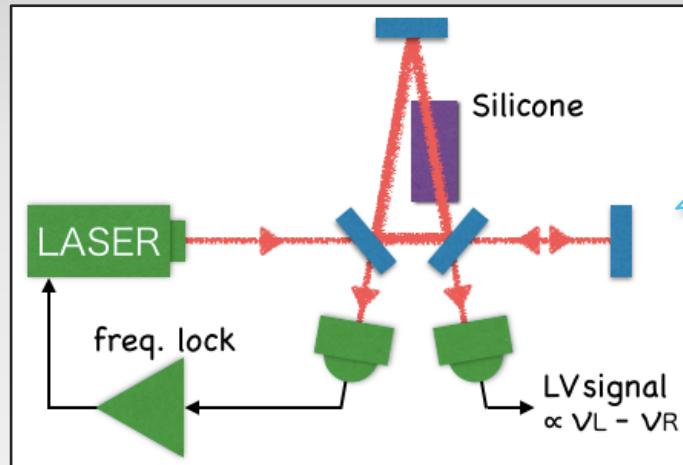
1. Mansouri-Sexl Theory
2. Optical Ring Cavity
3. Data Analysis
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Introduction

- Purpose of our experiment: Test of Lorentz invariance



Especially,
we search for Anisotropy
in the **one-way speed of light**
using optical ring cavity.

One-way speed of light

$$\begin{array}{c} \xrightarrow{c(\theta)} \\ \xleftarrow{c(\theta + \pi)} \end{array}$$

$$\frac{\delta c}{c} \lesssim 6 \times 10^{-15}$$

Y. Michimura (2013)

Round-trip speed of light

$$\begin{array}{c} \xrightarrow{c(\theta + \frac{\pi}{2})} \\ \xdownarrow{c(\theta)} \\ \xleftarrow{c(\theta)} \end{array}$$

$$\frac{\delta c}{c} \lesssim 1 \times 10^{-18}$$

M. Nagel (2015)

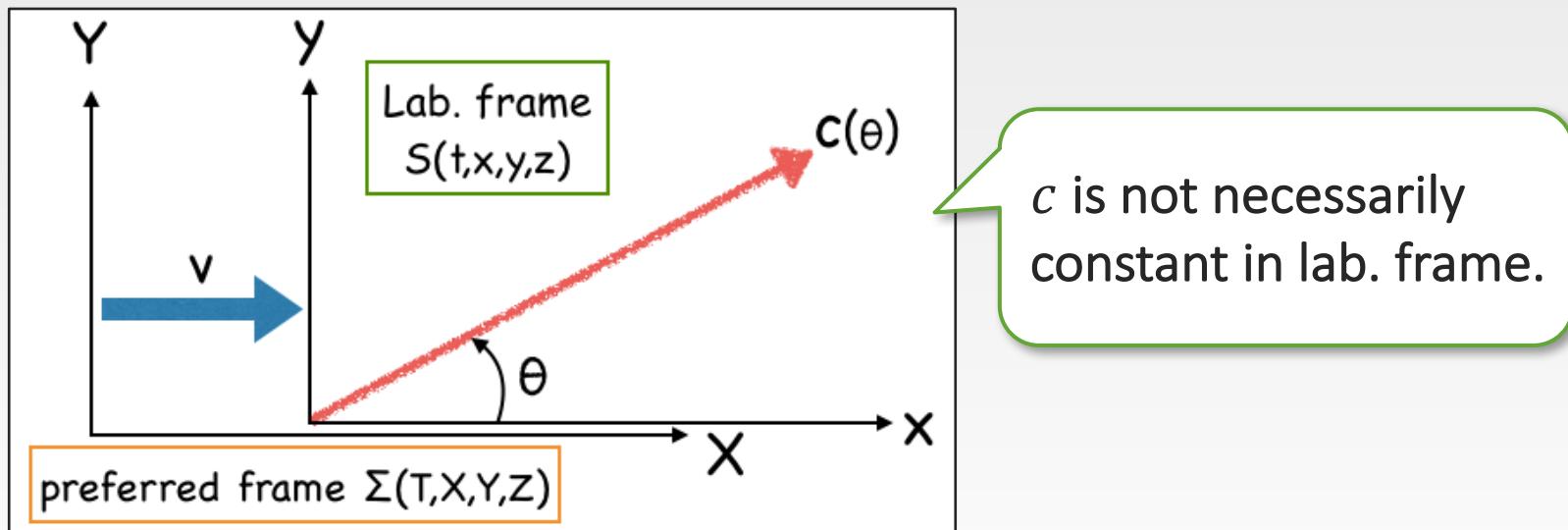
Test Theory

➤ 3 frameworks

- ✓ **Mansouri-Sexl theory** --- 道村さん修論
 - test theory of special relativity with 3 parameters (α, β, δ)
 - $\alpha + 1/2 = 0 \Leftrightarrow$ one-way speed of light is constant
- ✓ **Standard model extension** --- 道村さん(修論)・D論
 - more general theory with many parameters
 - can describe many particle (not only photon)
- ✓ **Spherical harmonic decomposition** --- 道村さんD論
 - no theoretical assumptions or background
 - $c(\theta, \phi) = 1 + \sum_{lm} \text{Re}[(\bar{y}_l^m)^* Y_l^m(\theta, \phi)]$

Mansouri-Sexl Theory

- Special relativity
 - ① Principle of relativity ② Invariance of c \Rightarrow Lorentz Invariance
- Mansouri-Sexl theory
 - ① There exists a preferred frame Σ , where c is constant
 - ② There is no preferred direction in Σ
 - + Clock synchronization by slow clock transport



↑Natural candidate: CMB rest frame

Speed of Light in MS Theory

① preferred frame Σ ($c = \text{const.}$) + ② no preferred direction in Σ

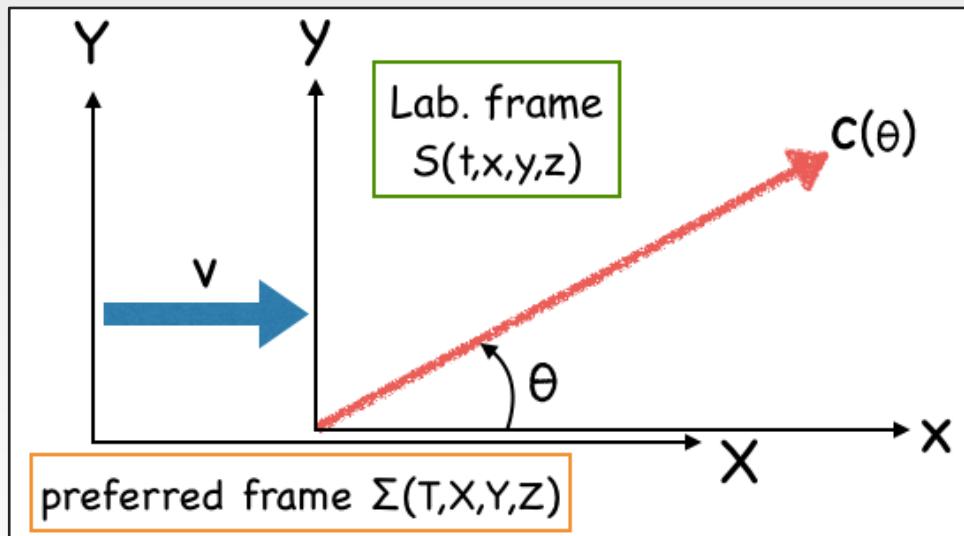


Dependence on v

$$c(\theta) = 1 - 2\left(\alpha + \frac{1}{2}\right)v \cos \theta - \left(\beta + \delta - \frac{1}{2}\right)v^2 \sin^2 \theta - (\alpha - \beta + 1)v^2 + \mathcal{O}(v^3)$$

One-way speed of light

Round-trip speed of light



In Special relativity,
 $\alpha = -\frac{1}{2}, \beta = \frac{1}{2}, \delta = 0$

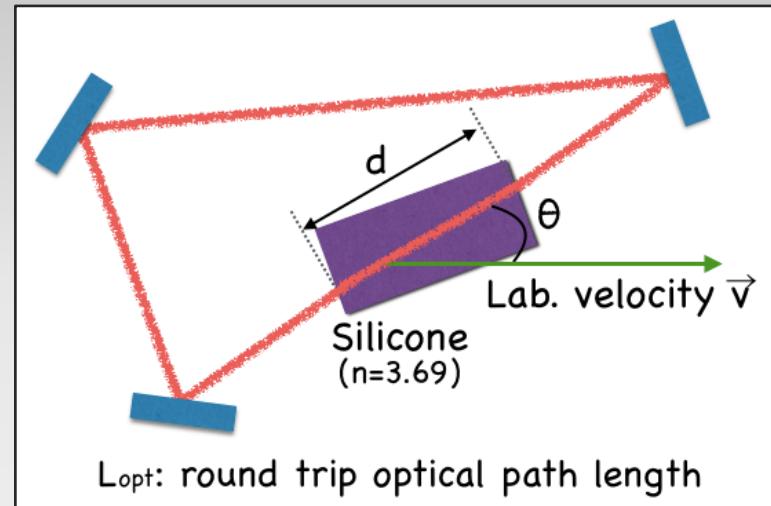
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Optical Ring Cavity (1/2)

➤ Speed of light

$$c(\theta) = 1 - \underbrace{2\left(\alpha + \frac{1}{2}\right)\nu \cos \theta}_{\text{One-way speed of light}}$$



➤ Resonant frequency

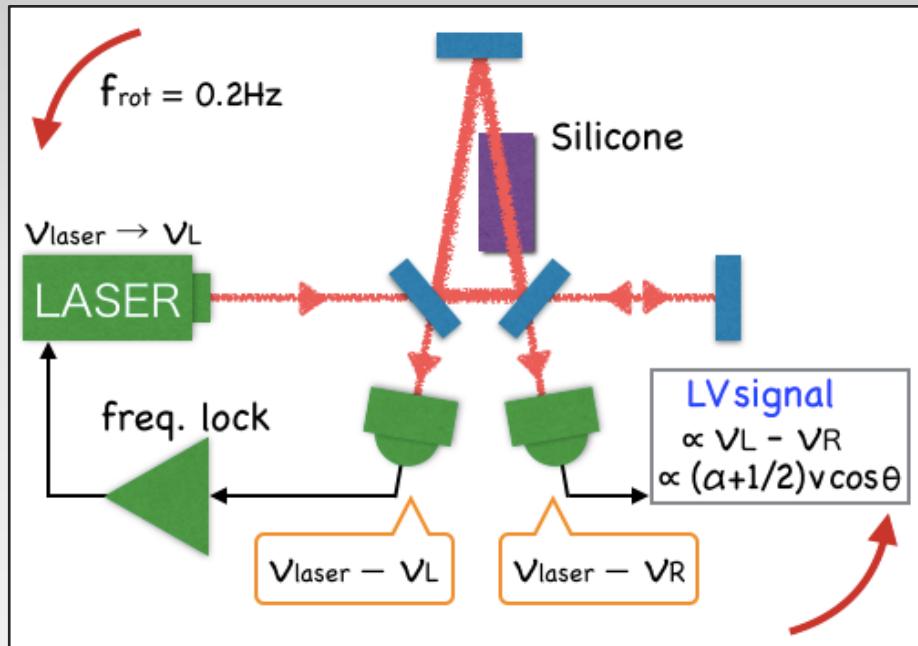
$$\begin{cases} \nu_L = \nu + \delta\nu/2 \text{ for clockwise} \\ \nu_R = \nu - \delta\nu/2 \text{ for counterclockwise} \end{cases}$$

where $\nu \equiv \frac{m}{L_{\text{opt}}} (m \in \mathbb{N})$, $\frac{\delta\nu}{2} \equiv -\nu \frac{2(n-1)d}{L_{\text{opt}}} \left(\alpha + \frac{1}{2}\right) \nu \cos \theta$

➤ Resonant frequency difference

$$\frac{\delta\nu}{\nu} = \frac{\nu_L - \nu_R}{\nu} = \frac{4(n-1)d}{L_{\text{opt}}} \left(\alpha + \frac{1}{2}\right) \nu \cos \theta$$

Optical Ring Cavity (2/2)



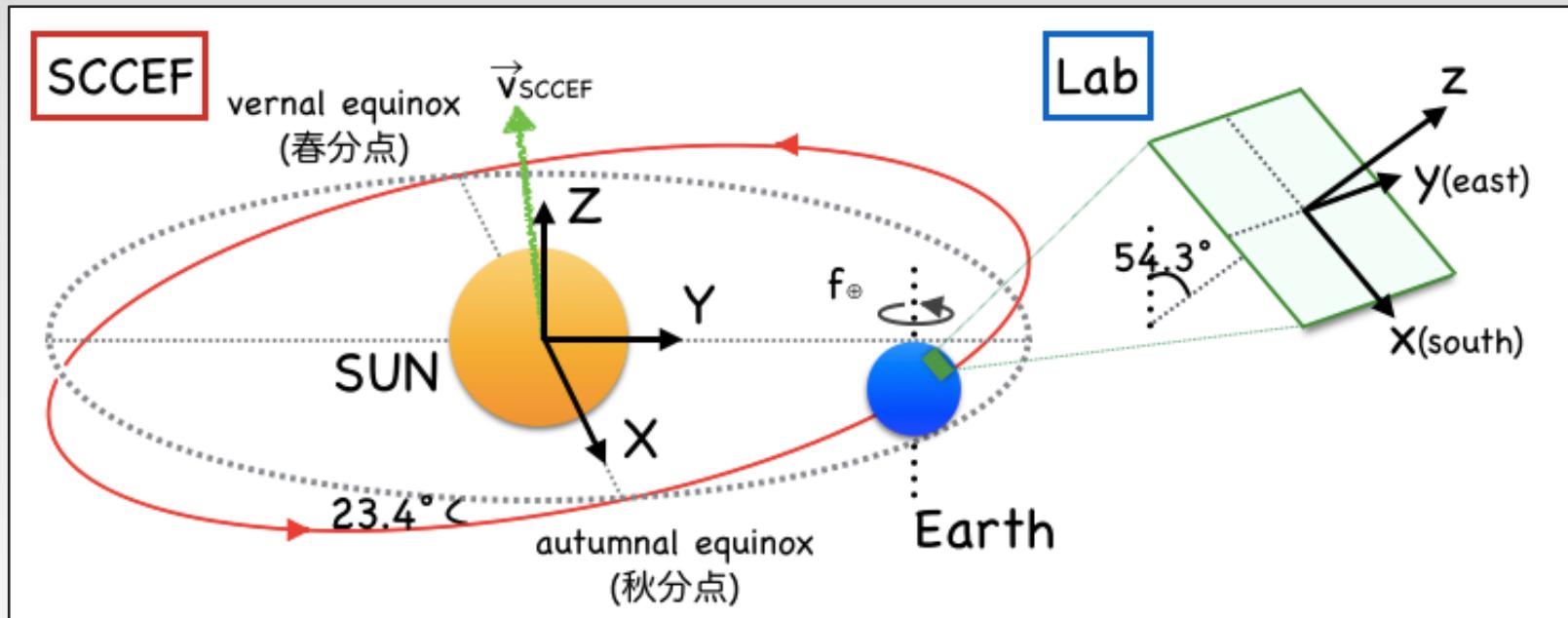
- Measuring $\nu_L - \nu_R$ using double-pass configuration.
- Getting the error signal ($\nu_{\text{laser}} - \nu_{L/R}$) by the Hänsch-Couillaud method.
= 偏光解析法
- Lorentz Violation signal → modulation by rotating the cavity
 - ↳ $\theta = \omega_{\text{rot}} t$? --- This is not true because of the Earth's rotation.
- We have to define a frame moving at a constant velocity with respect to Σ . → SCCEF

SCCEF

天の赤道

➤ SCCEF = Sun-Centered Celestial Equatorial Frame

⇒ Velocity with respect to Σ (CMB rest frame) is approximately constant.



➤ Velocity of SCCEF

$$\vec{v}_{\text{SCCEF}} = 369 \text{ km/s} \times \begin{pmatrix} -0.970 \\ 0.206 \\ 0.126 \end{pmatrix}$$

We can neglect

- orbital speed $\sim 30 \text{ km/s}$
- rotational speed $\sim 0.5 \text{ km/s}$

Lorentz Violation Signal

- Coordinate transformation matrix (**SCCEF** → **Lab.**)

$$R = \begin{pmatrix} \cos \chi \cos \omega_{\oplus} t & \cos \chi \sin \omega_{\oplus} t & -\sin \chi \\ -\sin \omega_{\oplus} t & \cos \omega_{\oplus} t & 0 \\ \sin \chi \cos \omega_{\oplus} t & \sin \chi \sin \omega_{\oplus} t & \cos \chi \end{pmatrix}$$

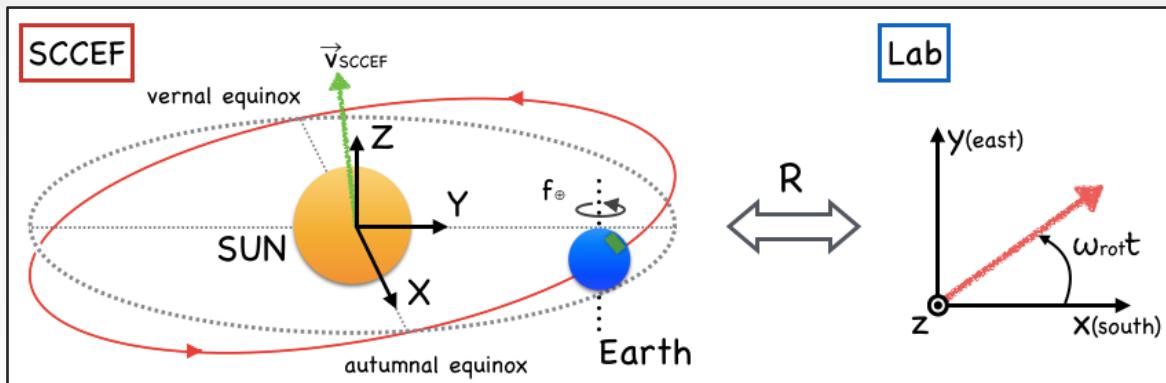


➤ LV signal

$$\frac{\delta\nu}{\nu} = \frac{4(n-1)d}{L_{\text{opt}}} \left(\alpha + \frac{1}{2} \right) v \cos \theta$$

$$= \frac{4(n-1)d}{L_{\text{opt}}} \left(\alpha + \frac{1}{2} \right) \vec{v} \cdot R^T \begin{pmatrix} \cos \omega_{\text{rot}} t \\ \sin \omega_{\text{rot}} t \\ 0 \end{pmatrix}$$

Unit vector along the light propagating in silicone
(in Lab. frame)



$\chi = 54.3^\circ$ (Colatitude of Lab.)

$$\frac{\omega_{\oplus}}{2\pi} = \frac{1}{23h\ 56m\ 4s}$$

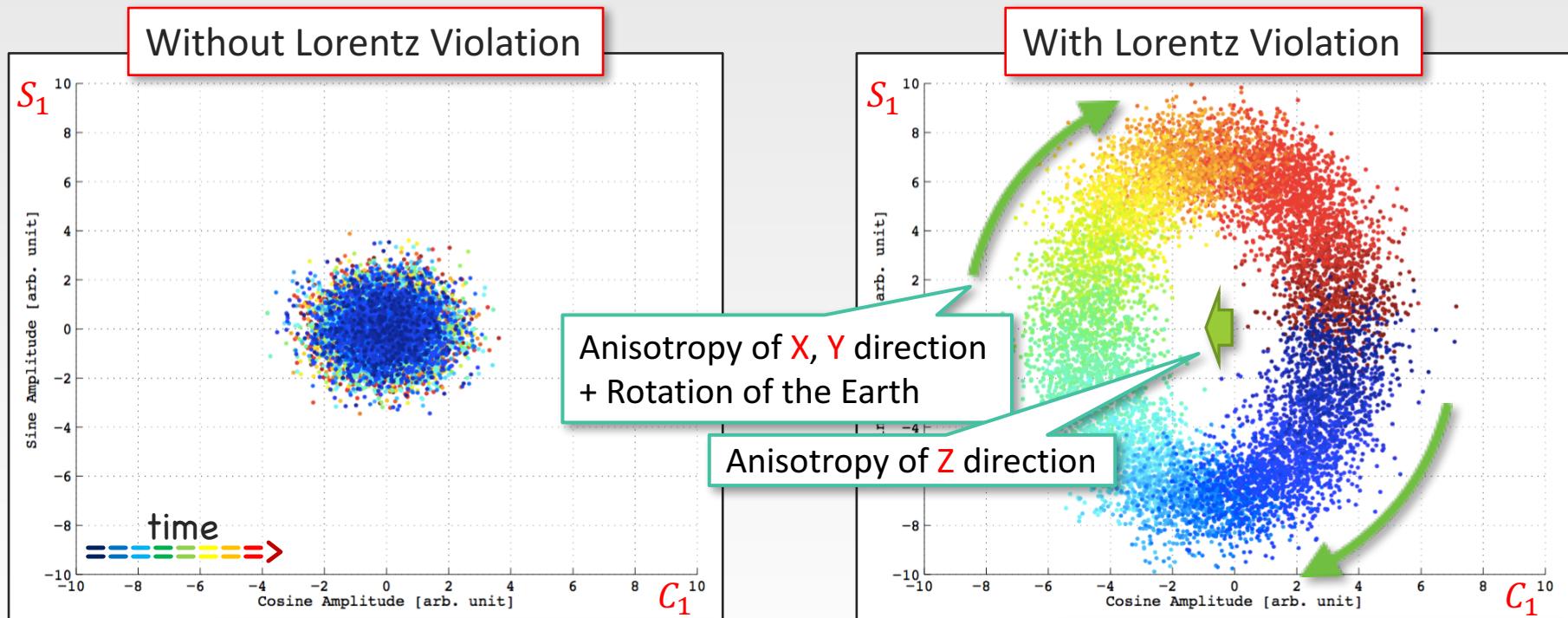
(Rotational frequency of the Earth)

Cos/Sin Amplitude of LV Signal

➤ LV signal

$$\frac{\delta\nu}{\nu} = \frac{4(n-1)d}{L_{\text{opt}}} \left(\alpha + \frac{1}{2} \right) \vec{v} \cdot R^T \begin{pmatrix} \cos \omega_{\text{rot}} t \\ \sin \omega_{\text{rot}} t \\ 0 \end{pmatrix}$$
$$= C_1 \cos \omega_{\text{rot}} t + S_1 \sin \omega_{\text{rot}} t$$

➤ Simulation (made by random number)

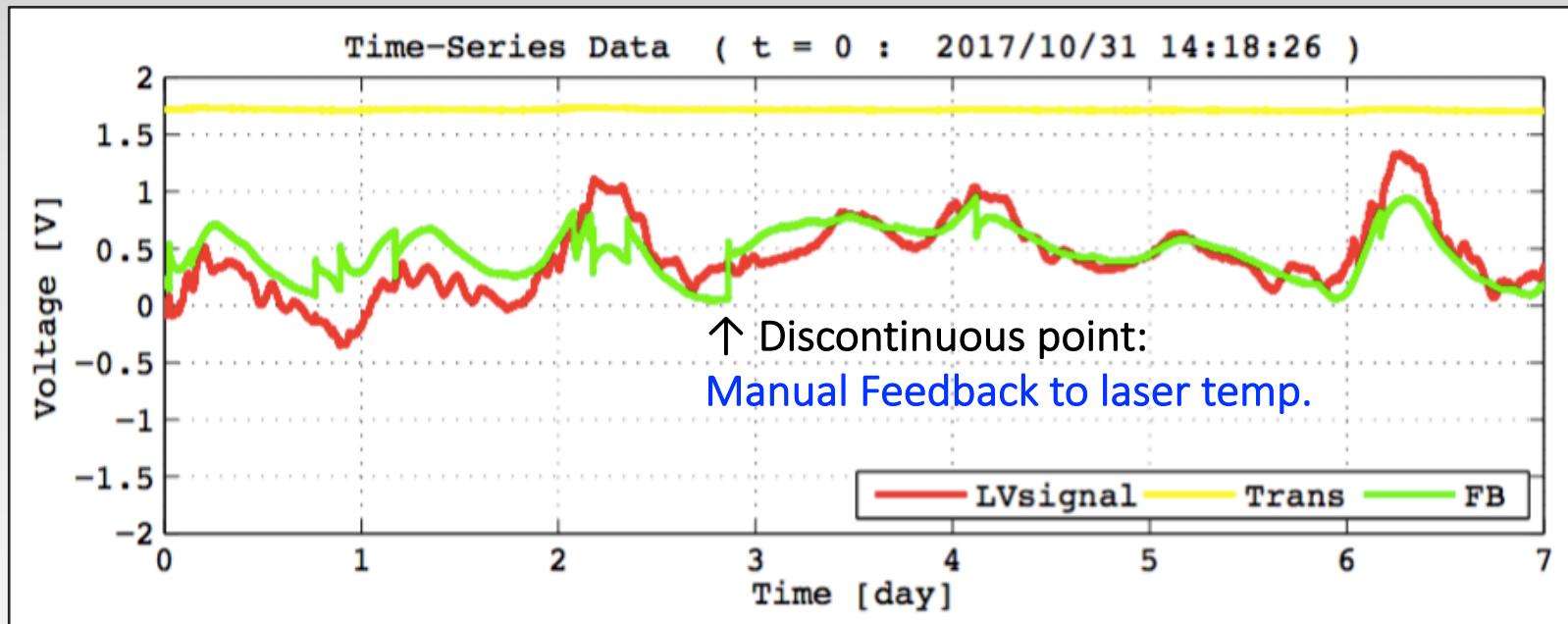


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Time-Series Data (1/2)

- Freq. lock had continued for 7 days.

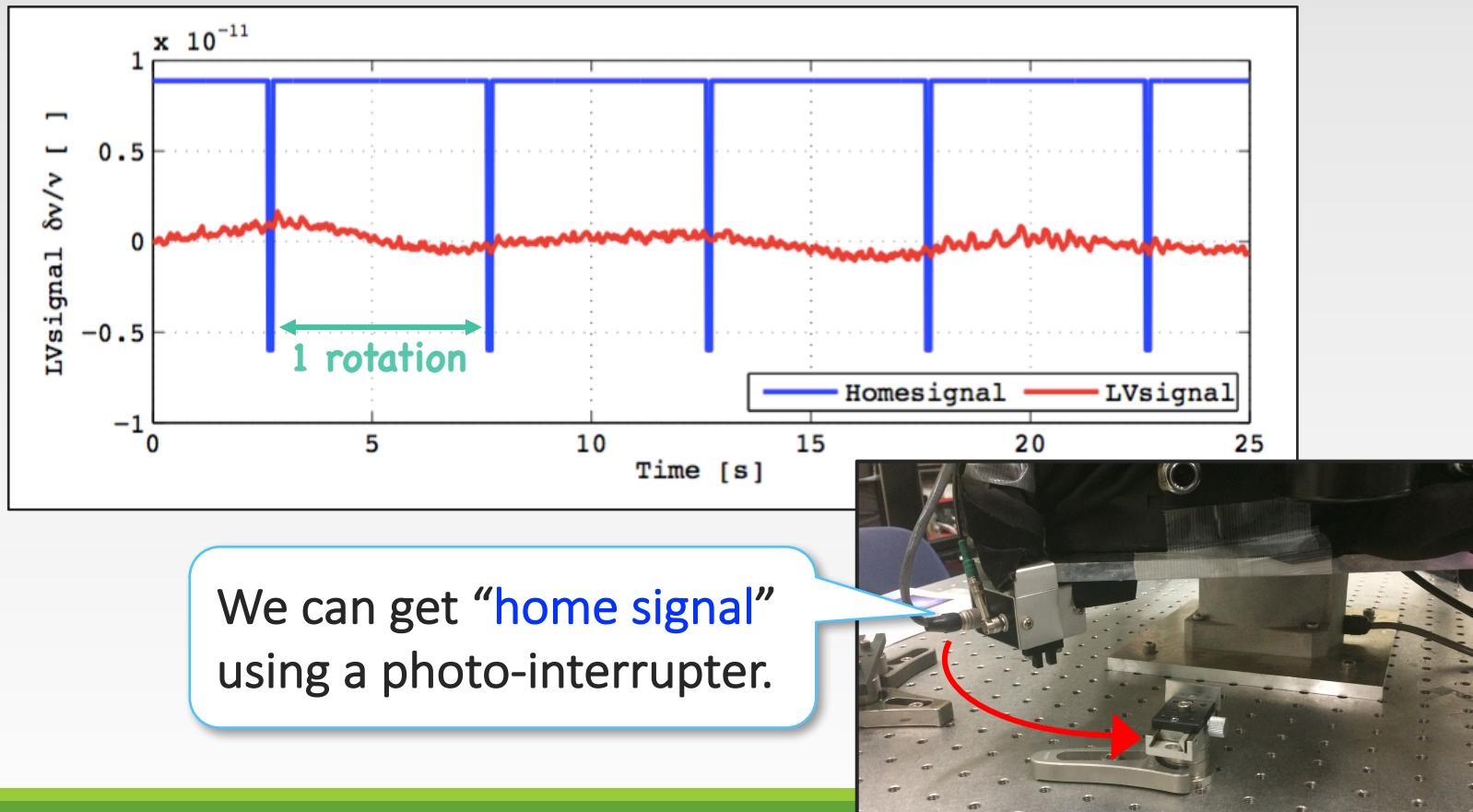


- Some correlation between **LV** and **FB**
⇒ If caused by cavity length change, $\text{CMRR} \lesssim 10^{-3}$

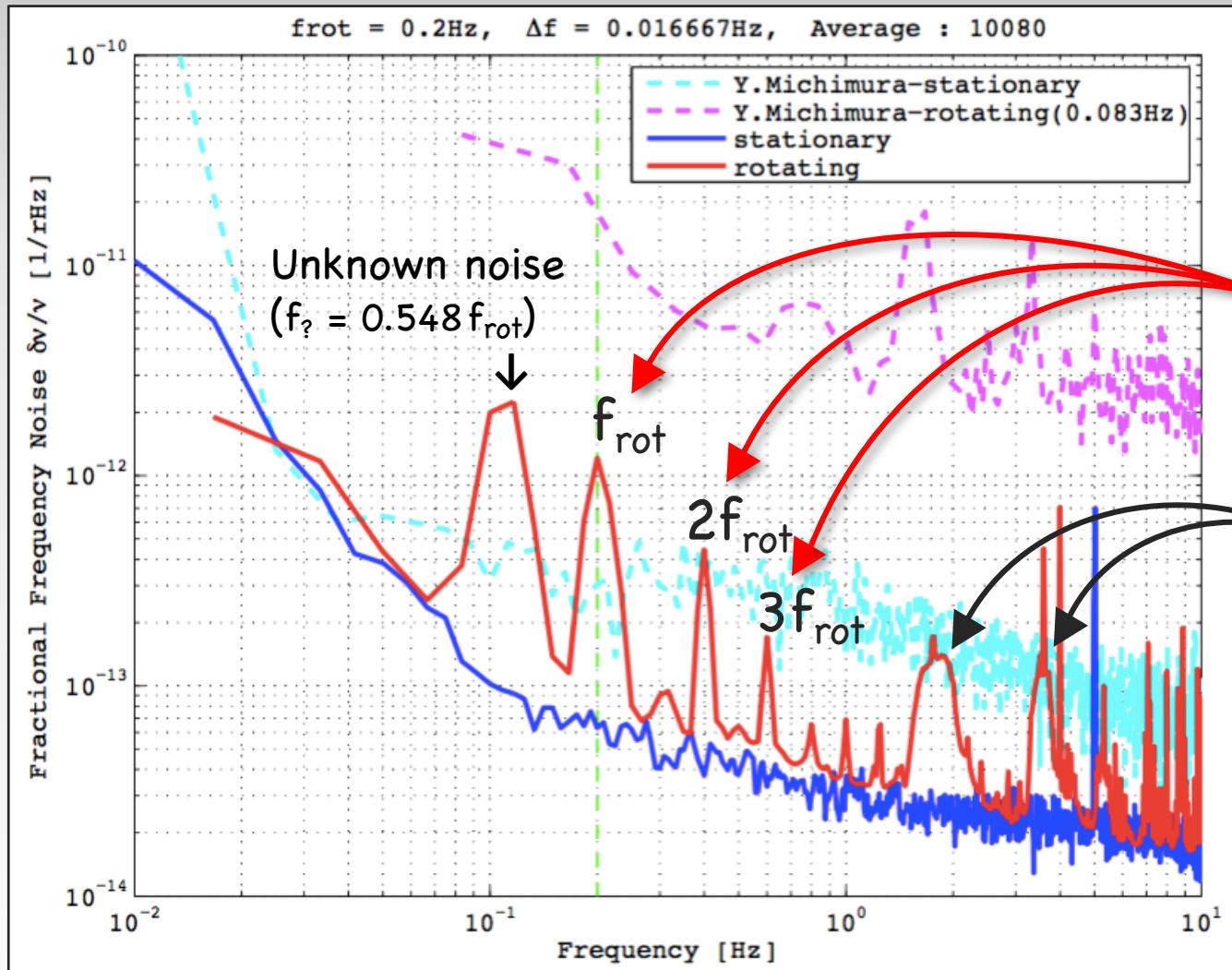
$$\nu_L - \nu_R$$

Time-Series Data (2/2)

- Enlarged view of LV signal
 - Cannot find rotational frequency noise visually
 - Unknown noise with ~2-rotation period



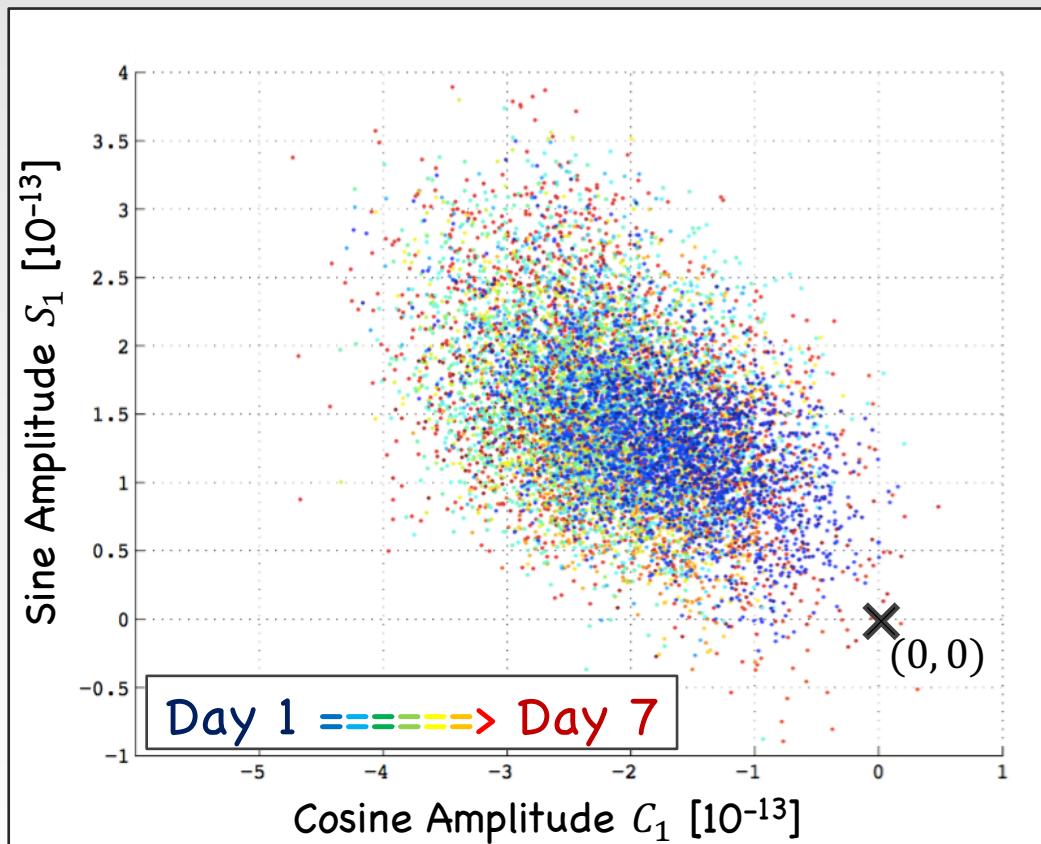
Power Spectrum of LV signal



Cos/Sin Amplitudes (1/2)

- Cos/Sin amplitudes (calculated every 12 rotations = 1min.)

$$\delta\nu/\nu = C_1 \cos \omega_{\text{rot}} t + S_1 \sin \omega_{\text{rot}} t$$



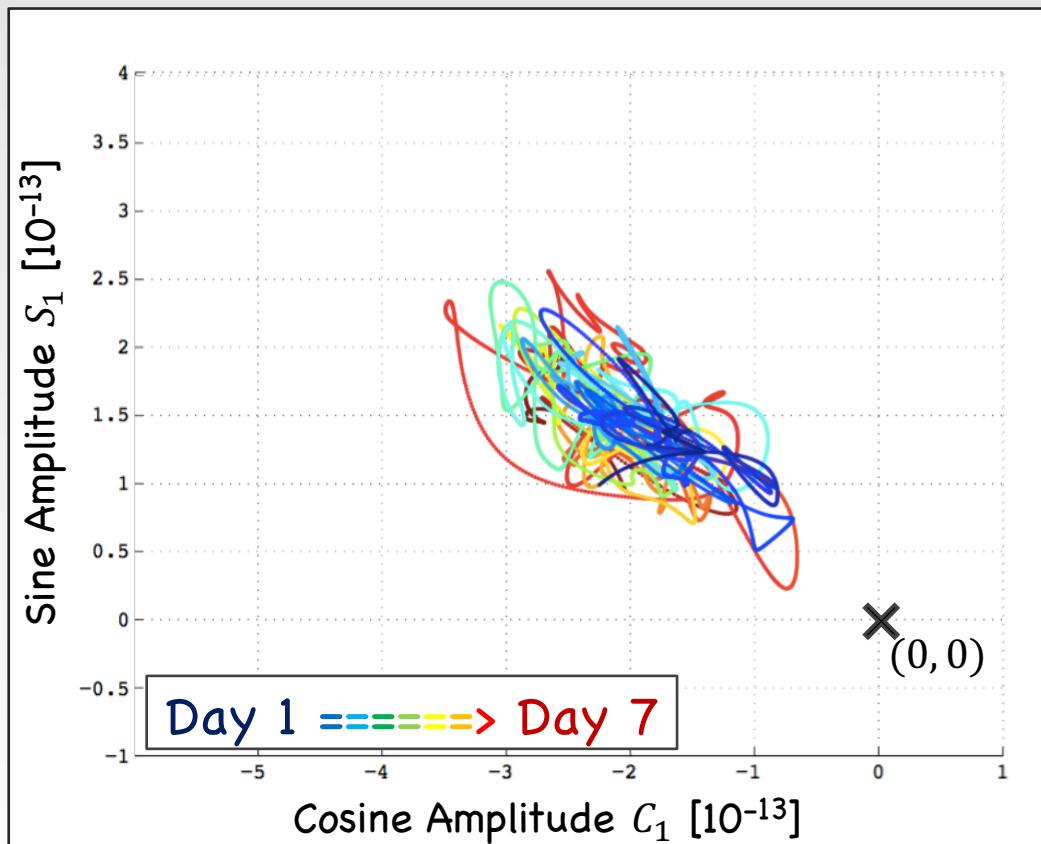
$$\overline{C}_1, \overline{S}_1 \neq 0$$

··· Rot. freq. noise

(If this gap is caused by anisotropy, $\delta c/c \sim 10^{-12}$)

Cos/Sin Amplitudes (2/2)

- Taking moving average (time constant ~ 1 hour) 前後1時間のデータと移動平均をとってみる
- Local mean value changes with time --- *Big problem!!!*



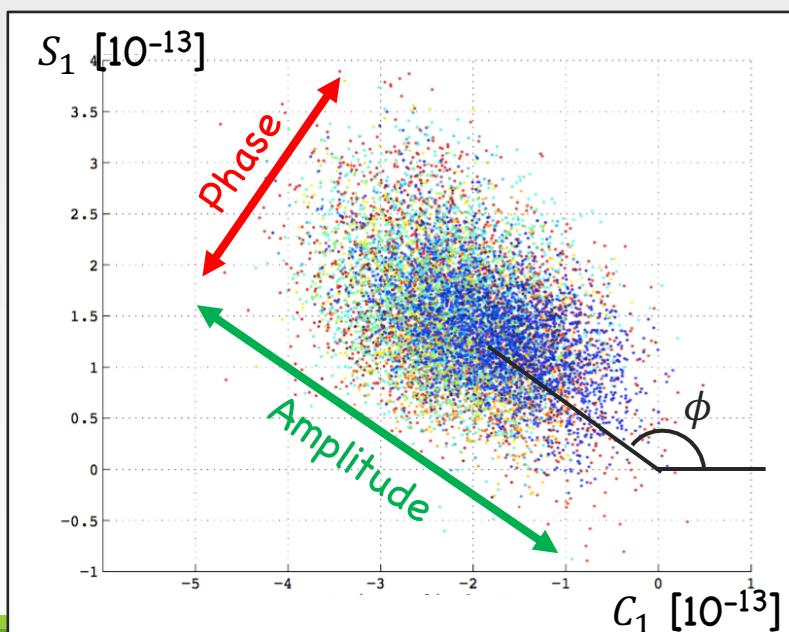
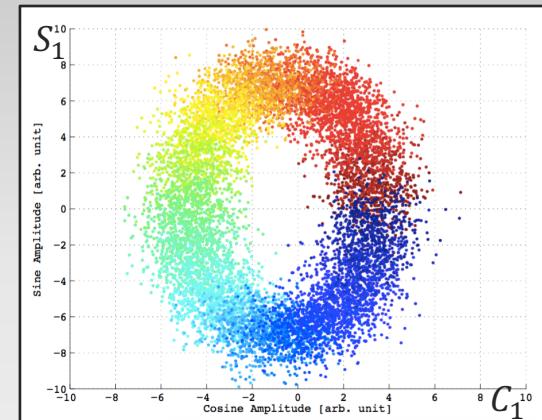
because we cannot
distinguish mean value
change **by anisotropy**
with **by noise**.

Current sensitivity is
limited by this effect.

(noise at $f_{\text{rot}} \pm f_{\oplus}$)

How to calculate $\alpha + 1/2$

- If anisotropy exists,
 - C_1 is drifted Anisotropy of Z direction
 - C_1 changes @ f_\oplus } Anisotropy of X, Y direction
 - S_1 changes @ f_\oplus } + Rotation of the Earth
- ⇒ These 3 components has the same information about $\alpha + 1/2$.



- I used “phase component”
 $p_1 \equiv -\sin \phi (C_1 - \bar{C}_1) + \cos \phi (S_1 - \bar{S}_1)$
(ϕ : phase of rot. Freq. noise)

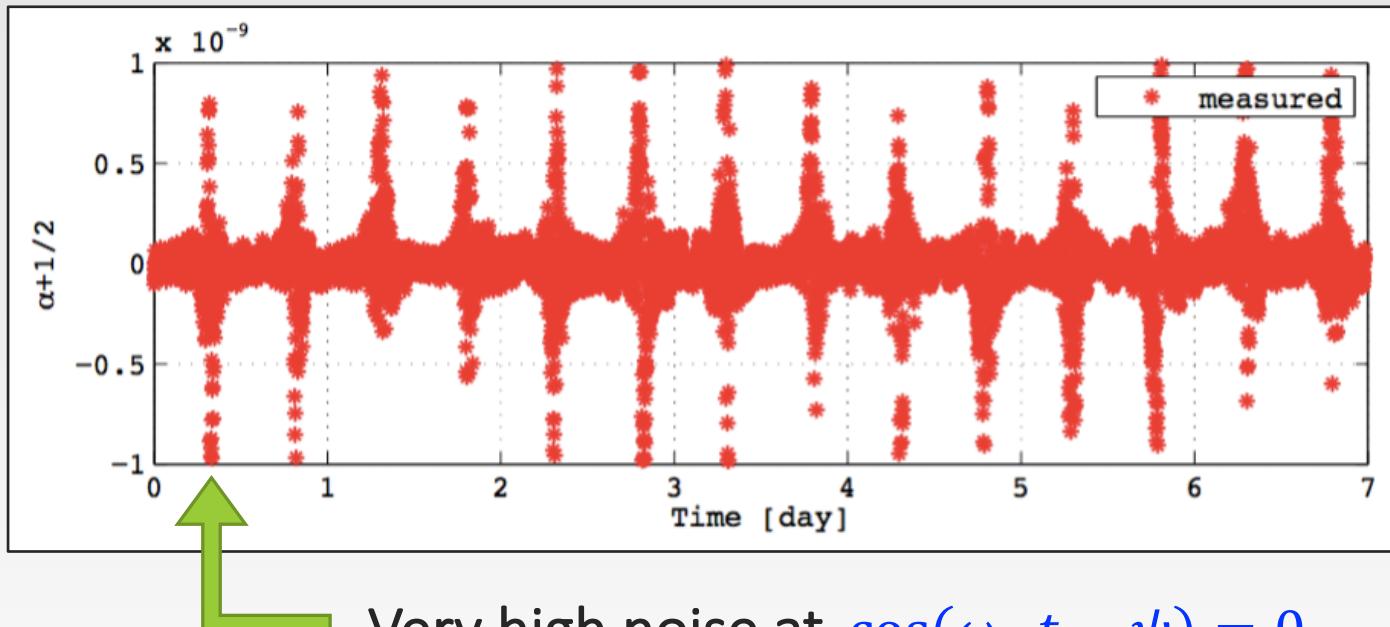
↓

$$p_1 = A \left(\alpha + \frac{1}{2} \right) \cos(\omega_\oplus t - \psi)$$

$$\alpha + 1/2$$

- I calculated $\alpha + 1/2$ every 12 rotations

$$\alpha + \frac{1}{2} = \frac{p_1}{A \cos(\omega_{\oplus} t - \psi)}$$



Very high noise at $\cos(\omega_{\oplus} t - \psi) = 0$

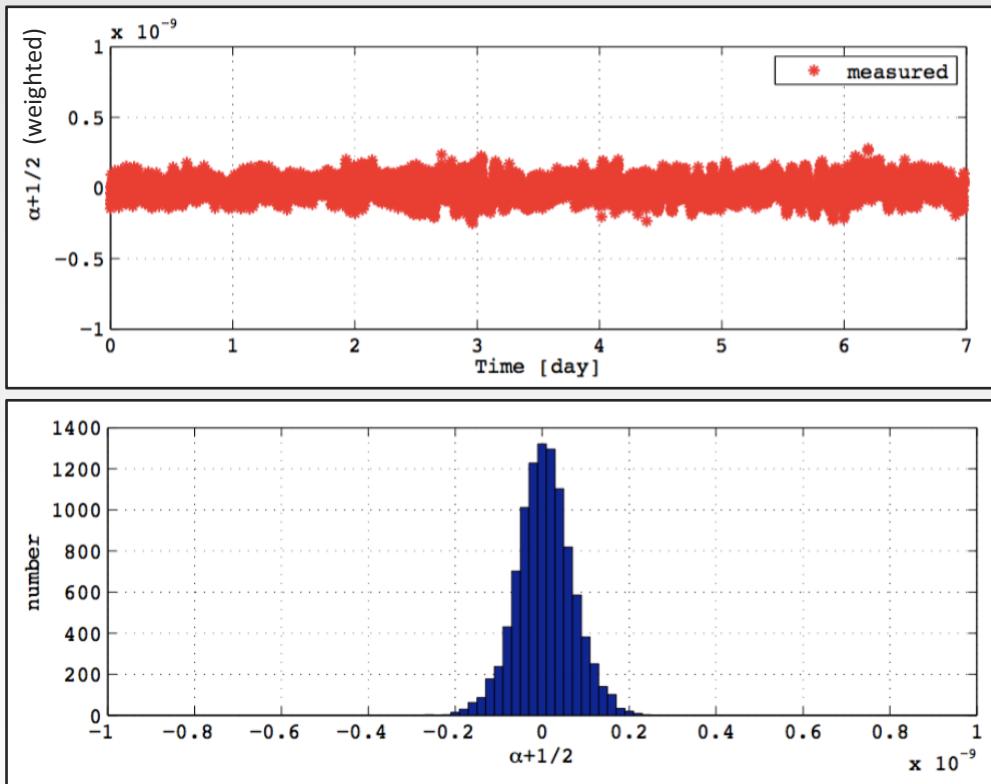
- Need to take weighted average (Normal average \rightarrow Sensitivity \downarrow)

Weighted $\alpha + 1/2$

- Taking weighted average

$$\alpha + \frac{1}{2} = \frac{p_1}{A \cos(\omega_{\oplus} t - \psi)}$$

Weight $\propto |\cos(\omega_{\oplus} t - \psi)|$



- Weighted average:

$$\overline{\left(\alpha + \frac{1}{2}\right)_W} = 5 \times 10^{-13}$$

- Standard deviation:

$$\delta \left(\alpha + \frac{1}{2} \right)_W \sim 10^{-10}$$



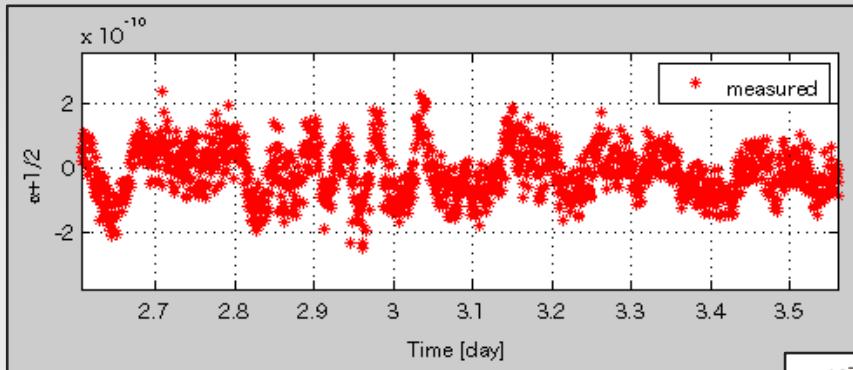
Number of data: $N \sim 10^5$

- Sensitivity:

$$\delta \overline{\left(\alpha + \frac{1}{2} \right)_W} \cancel{\propto} \frac{\delta \left(\alpha + \frac{1}{2} \right)_W}{\sqrt{N}} \sim 3 \times 10^{-13} \quad (?)$$

Really Independent?

- Enlarged view of weighted $\alpha + 1/2$



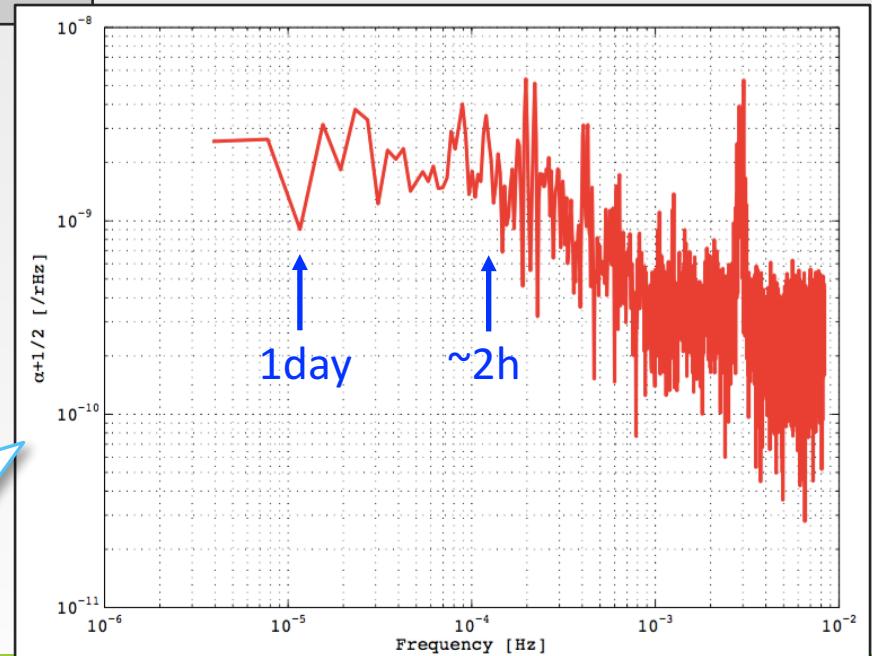
There is some trend!



Nearby data points are
NOT independent.

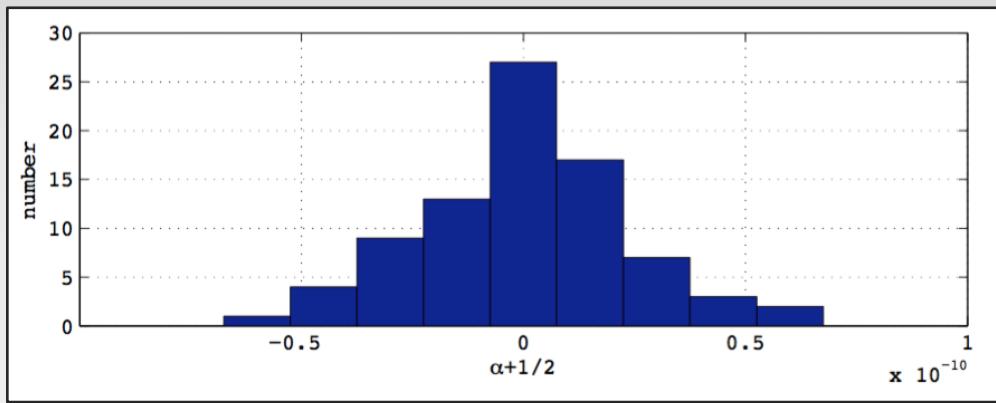
- Power spectrum of weighted $\alpha + 1/2$

2 points are independent if **more than 2 hours away** from each other.



Result

- Taking average **every 2 hours**



- Standard deviation:

$$\delta \left(\alpha + \frac{1}{2} \right)_w = 2.3 \times 10^{-11}$$

Number of data: $N \sim 84$

- Sensitivity:

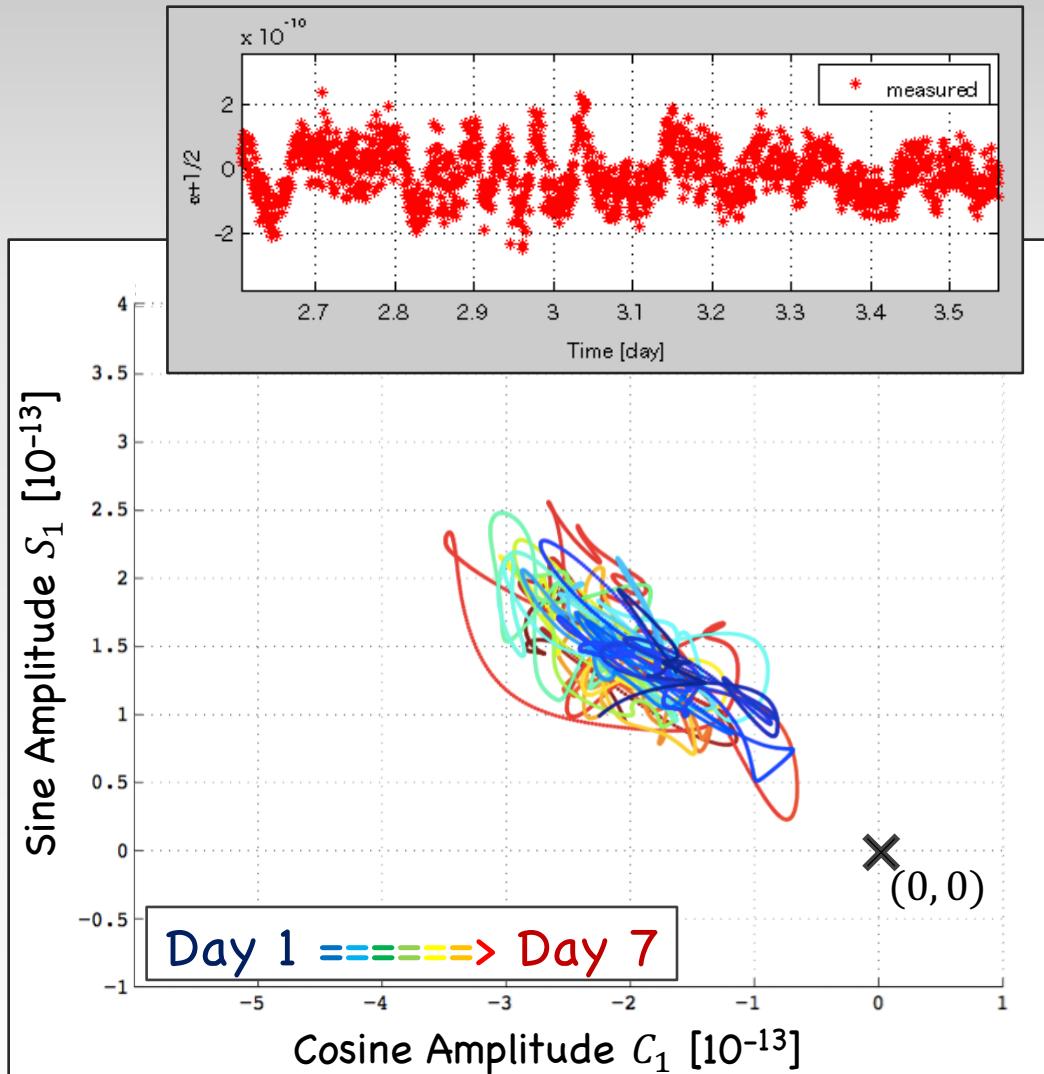
$$\delta \frac{\alpha + \frac{1}{2}}{w} = \frac{\delta \left(\alpha + \frac{1}{2} \right)_w}{\sqrt{N}} \sim 2.5 \times 10^{-12}$$

Result

$$\alpha + \frac{1}{2} = (-0.5 \pm 2.5) \times 10^{-12}$$

$$\frac{\delta c}{c} \sim 2 \left(\alpha + \frac{1}{2} \right) v_{SCCEF} \sim 6 \times 10^{-15}$$

Discussion



- Noise level is reduced, but now **noise at $f_{\text{rot}} \pm f_{\oplus}$** limits the sensitivity.
- In the framework of spherical harmonic decomposition, **anisotropy of Z direction cannot be measured unless rot. freq. noise is reduced by 3 order of magnitude.**

今の装置では限界？吊るしかない？

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Conclusion

- Data analysis in Mansouri-Sexl framework
(using a week-long data: Oct. 31st - Nov. 7th)

$$\alpha + 1/2 = (-0.5 \pm 2.5) \times 10^{-12} \quad \left(\frac{\delta c}{c} \sim 6 \times 10^{-15} \right)$$

- ✓ The same order of magnitude as world record
- ✓ Noise at $f_{\text{rot}} \pm f_{\oplus}$ is big problem
- ✓ 来年の異方性実験のテーマ: 光学系を吊った状態で
回せる連続回転機構(傾き制御付き)の開発

ほそく：回転周波数依存性

- 回転周波数ノイズ・謎のノイズとも振幅は変わらない

