

Quantum Optomechanical Heat Engine

~Ando-Lab Seminar on Sep. 2nd~

Kentaro Komori

Optomechanics \Leftrightarrow Thermodynamics

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Quantum Optomechanical Heat Engine

Keye Zhang,^{1,2} Francesco Bariani,² and Pierre Meystre²

¹*Quantum Institute for Light and Atoms, Department of Physics,
East China Normal University, Shanghai 200241, People's Republic of China*

²*B2 Institute, Department of Physics and College of Optical Sciences,
University of Arizona, Tucson, Arizona 85721, USA*

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We investigate theoretically a quantum optomechanical realization of a heat engine. In a generic optomechanical arrangement the optomechanical coupling between the cavity field and the oscillating end mirror results in polariton normal mode excitations whose character depends on the pump detuning and the coupling strength. By varying that detuning it is possible to transform their character from phononlike to photonlike, so that they are predominantly coupled to the thermal reservoir of phonons or photons, respectively. We exploit the fact that the effective temperatures of these two reservoirs are different to produce an Otto cycle along one of the polariton branches. We discuss the basic properties of the system in two different regimes: in the optical domain it is possible to extract work from the thermal energy of a mechanical resonator at finite temperature, while in the microwave range one can in principle exploit the cycle to extract work from the blackbody radiation background coupled to an ultracold atomic ensemble.

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Contents

- What is the heat engine?
- Hamiltonian and polaritons
- Otto cycle
- Work and efficiency
- Cosmic blackbody radiation

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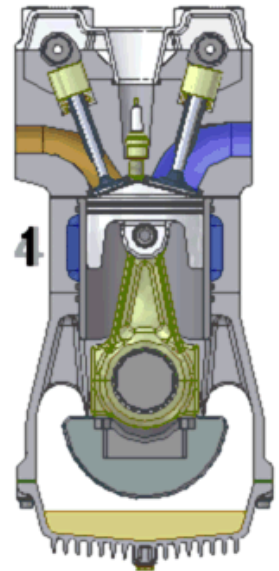
What is the heat engine?

✓ The machine producing **work** from **heat**

For example ...

✧ Car engine (a kind of otto cycle)

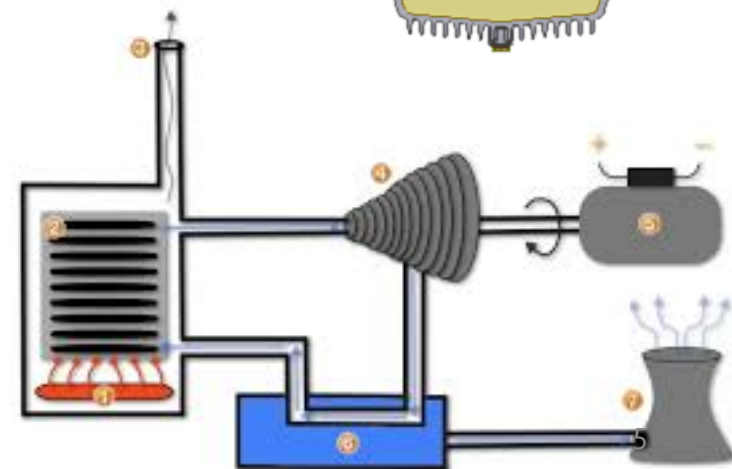
1. Intake, 2. Compression,
3. Power, 4. Exhaust



✧ Steam turbine

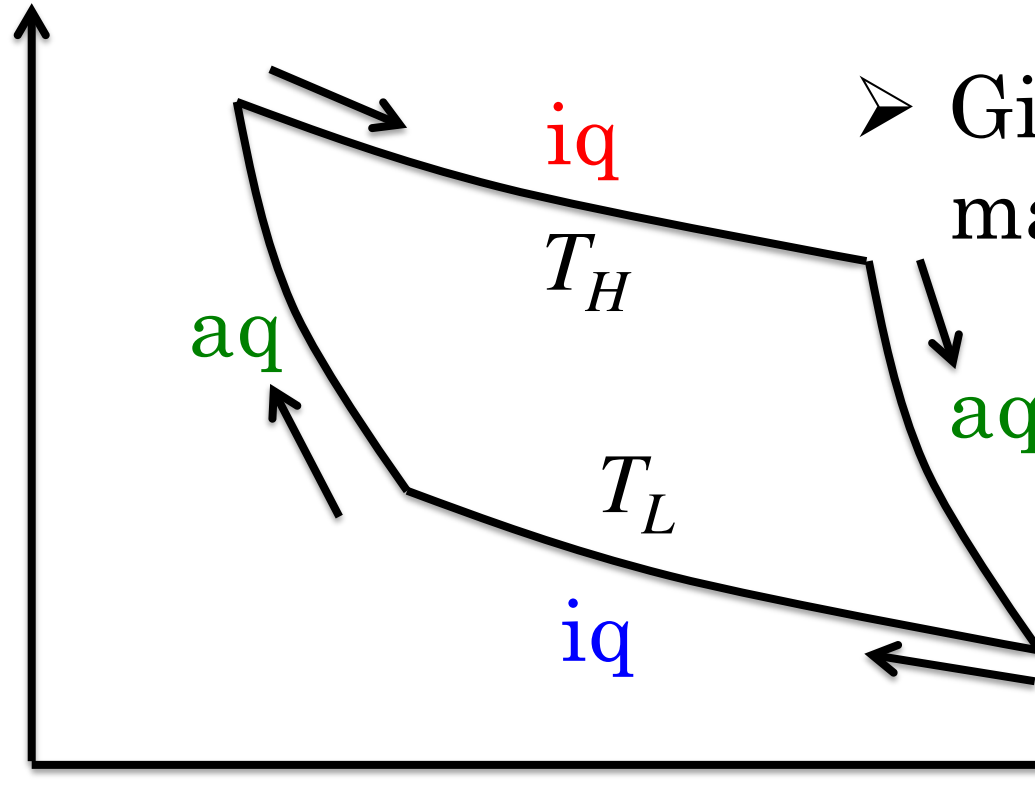
Fuel → Gas

Heat  Work 



Cycles at thermodynamics

- Carnot cycle iq = isothermal quasistatic
 aq = adiabatic quasistatic

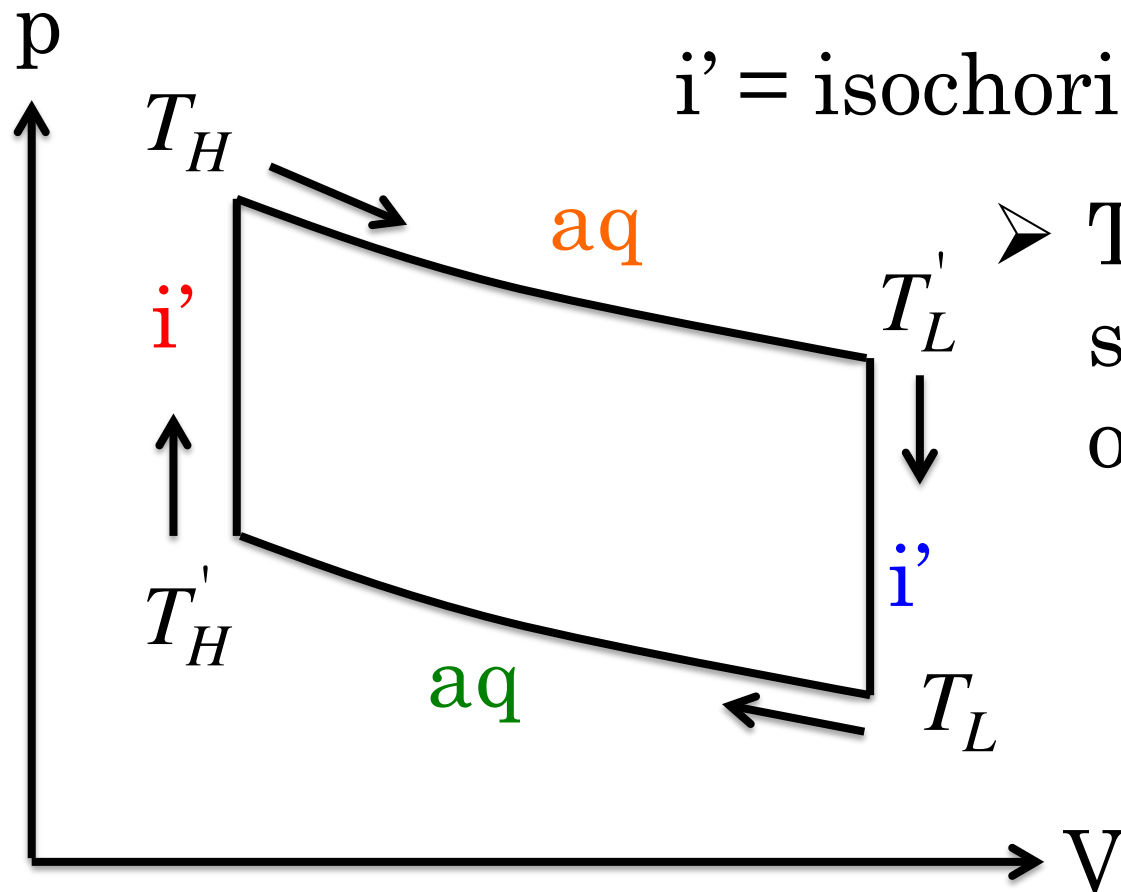


- Giving the maximum efficient

$$\varepsilon_0 = 1 - \frac{T_L}{T_H}$$

Cycles at thermodynamics

- Otto cycle (considered at this seminar)



- The efficient is smaller than one of carnot

$$\varepsilon = 1 - \frac{T_L}{T'_H} < \varepsilon_0$$

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Hamiltonian

➤ Linearized optomechanical coupling

$$H = \underbrace{-\hbar\Delta\hat{a}^\dagger\hat{a}}_{\text{optical}} + \underbrace{\hbar\omega_m\hat{b}^\dagger\hat{b}}_{\text{mechanical}} + \underbrace{\hbar G(\hat{b} + \hat{b}^\dagger)(\hat{a} + \hat{a}^\dagger)}_{\text{optomechanical}}$$

$\bar{\alpha} + \hat{a}$: average + photon annihilation

\hat{b} : phonon annihilation

Δ : cavity detuning

ω_m : mechanical resonant frequency

Hamiltonian

➤ Linearized optomechanical coupling

$$\begin{aligned} H &= -\hbar\Delta\hat{a}^\dagger\hat{a} + \hbar\omega_m\hat{b}^\dagger\hat{b} + \hbar G(\hat{b} + \hat{b}^\dagger)(\hat{a} + \hat{a}^\dagger) \\ &= -\hbar\Delta\frac{X_c^2 + P_c^2}{2} + \hbar\omega_m\frac{X_m^2 + P_m^2}{2} + 2\hbar GX_mX_c \end{aligned}$$

$$X_c = \frac{\hat{a} + \hat{a}^\dagger}{\sqrt{2}}, P_c = \frac{\hat{a} - \hat{a}^\dagger}{\sqrt{2}i} \quad : \text{optical quadrature}$$

$$X_m = \frac{\hat{b} + \hat{b}^\dagger}{\sqrt{2}}, P_m = \frac{\hat{b} - \hat{b}^\dagger}{\sqrt{2}i} \quad : \text{mechanical quadrature}$$

Hamiltonian

➤ Linearized optomechanical coupling

$$\begin{aligned} H &= -\hbar\Delta\hat{a}^\dagger\hat{a} + \hbar\omega_m\hat{b}^\dagger\hat{b} + \hbar G(\hat{b} + \hat{b}^\dagger)(\hat{a} + \hat{a}^\dagger) \\ &= -\hbar\Delta\frac{X_c^2 + P_c^2}{2} + \hbar\omega_m\frac{X_m^2 + P_m^2}{2} + 2\hbar GX_mX_c \\ &= \hbar\omega_+\frac{X_+^2 + P_+^2}{2} + \hbar\omega_-\frac{X_-^2 + P_-^2}{2} \\ &= \hbar\omega_+\underline{\hat{A}^\dagger\hat{A}} + \hbar\omega_-\underline{\hat{B}^\dagger\hat{B}} \end{aligned}$$

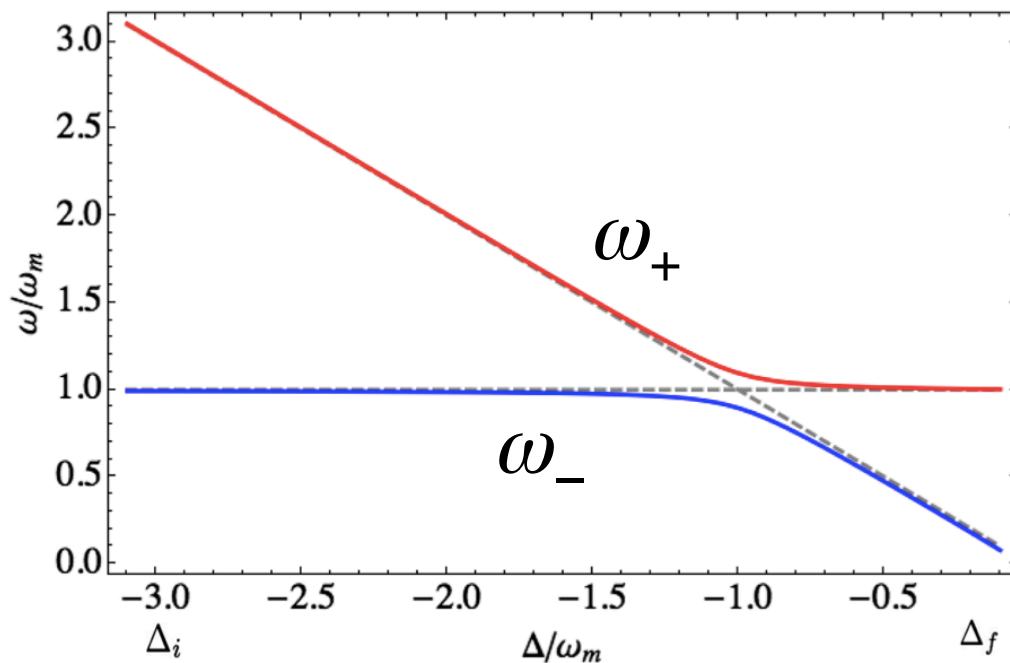

diagonalization

new bosonic annihilation operators
(linear combinations of $\hat{a}, \hat{a}^\dagger, \hat{b}, \hat{b}^\dagger$)

Hamiltonian

➤ the frequency of polaritons

$$\omega_{\pm} = \sqrt{\frac{\Delta^2 + \omega_m^2 \pm \sqrt{(\Delta^2 - \omega_m^2)^2 - 16G^2\Delta\omega_m}}{2}}$$



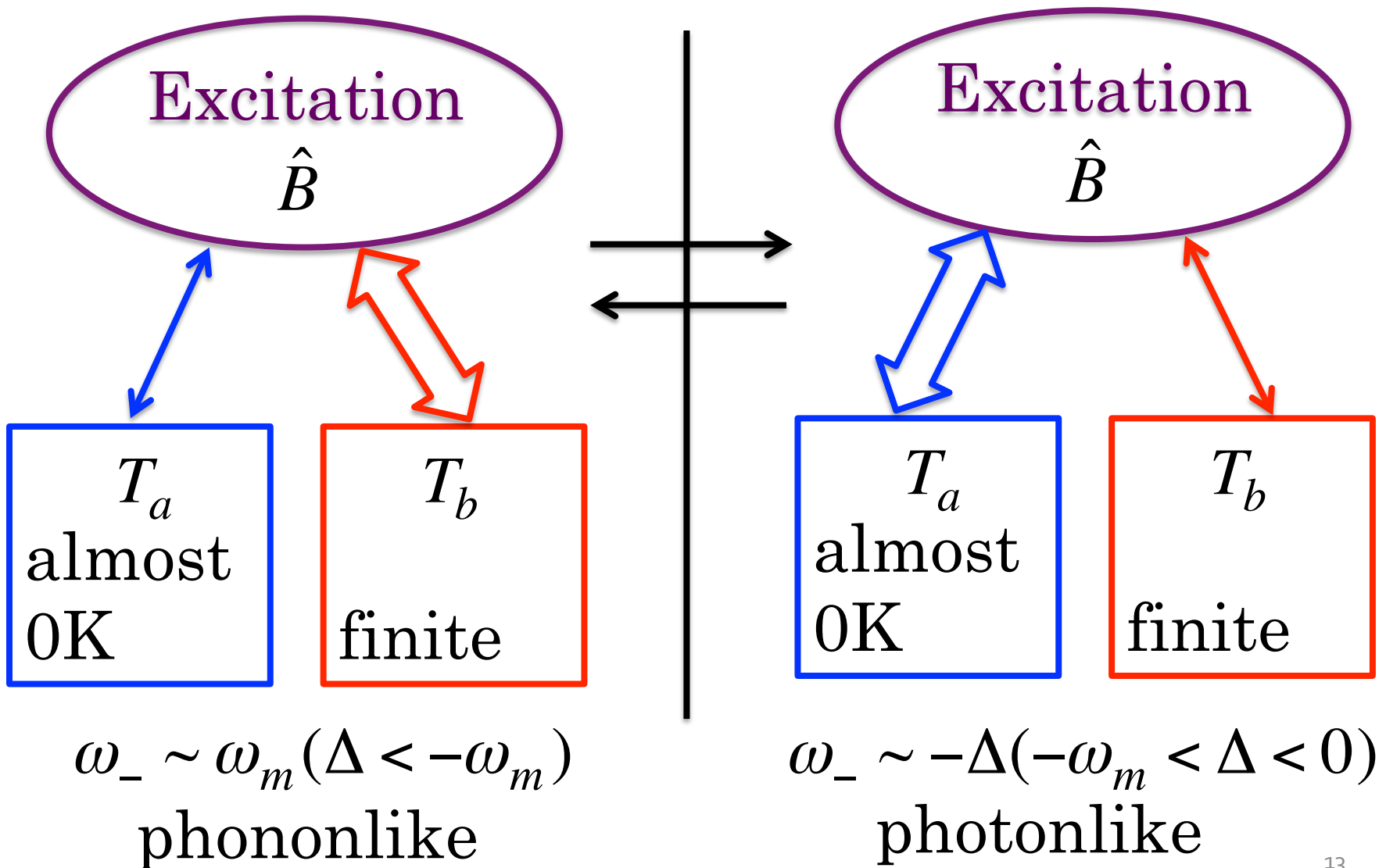
Focusing on the
normal-mode \hat{B}

$\omega_- \sim \omega_m (\Delta < -\omega_m)$
phononlike



$\omega_- \sim -\Delta (-\omega_m < \Delta < 0)$
photonlike

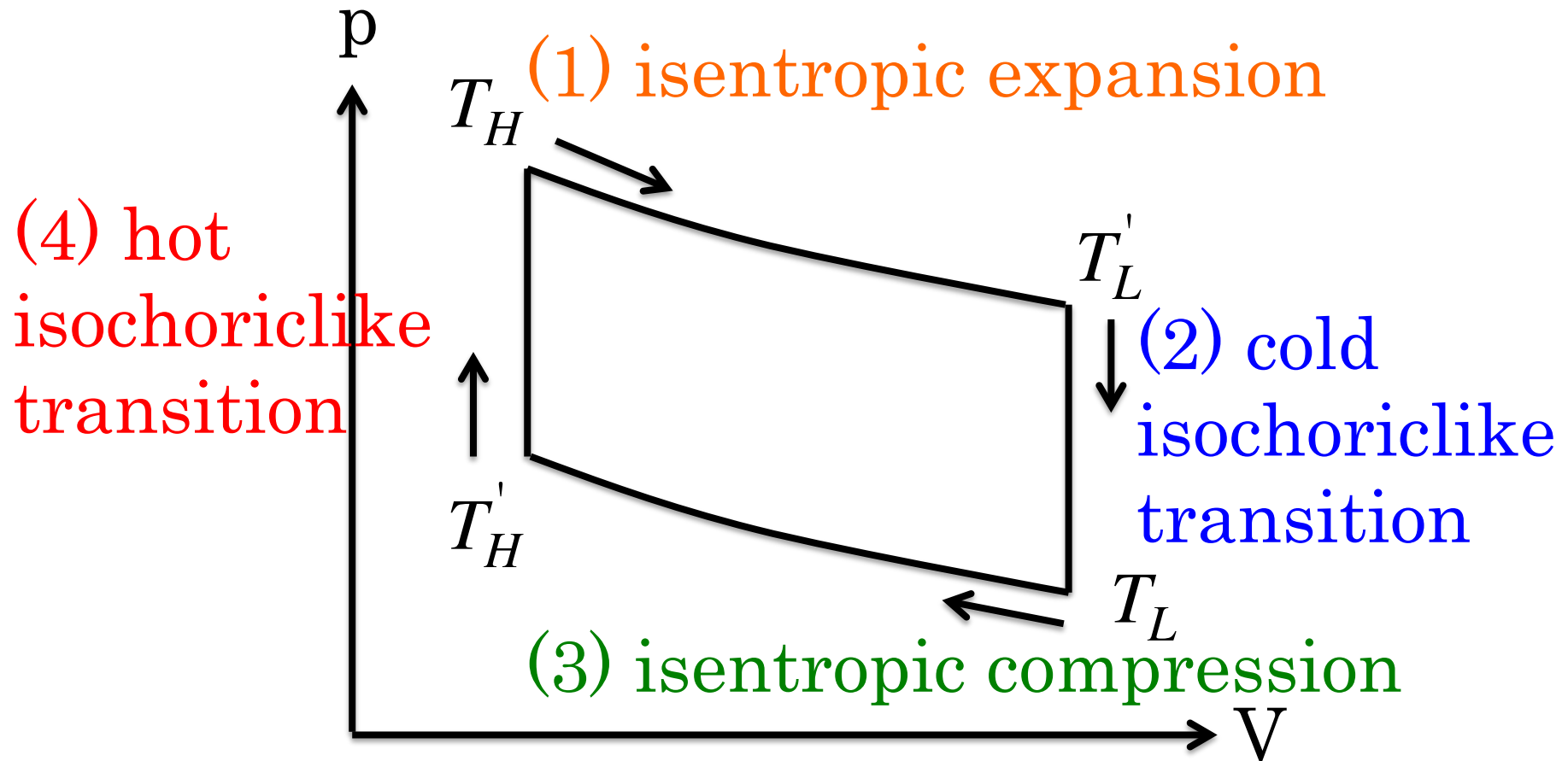
Situation



Contents

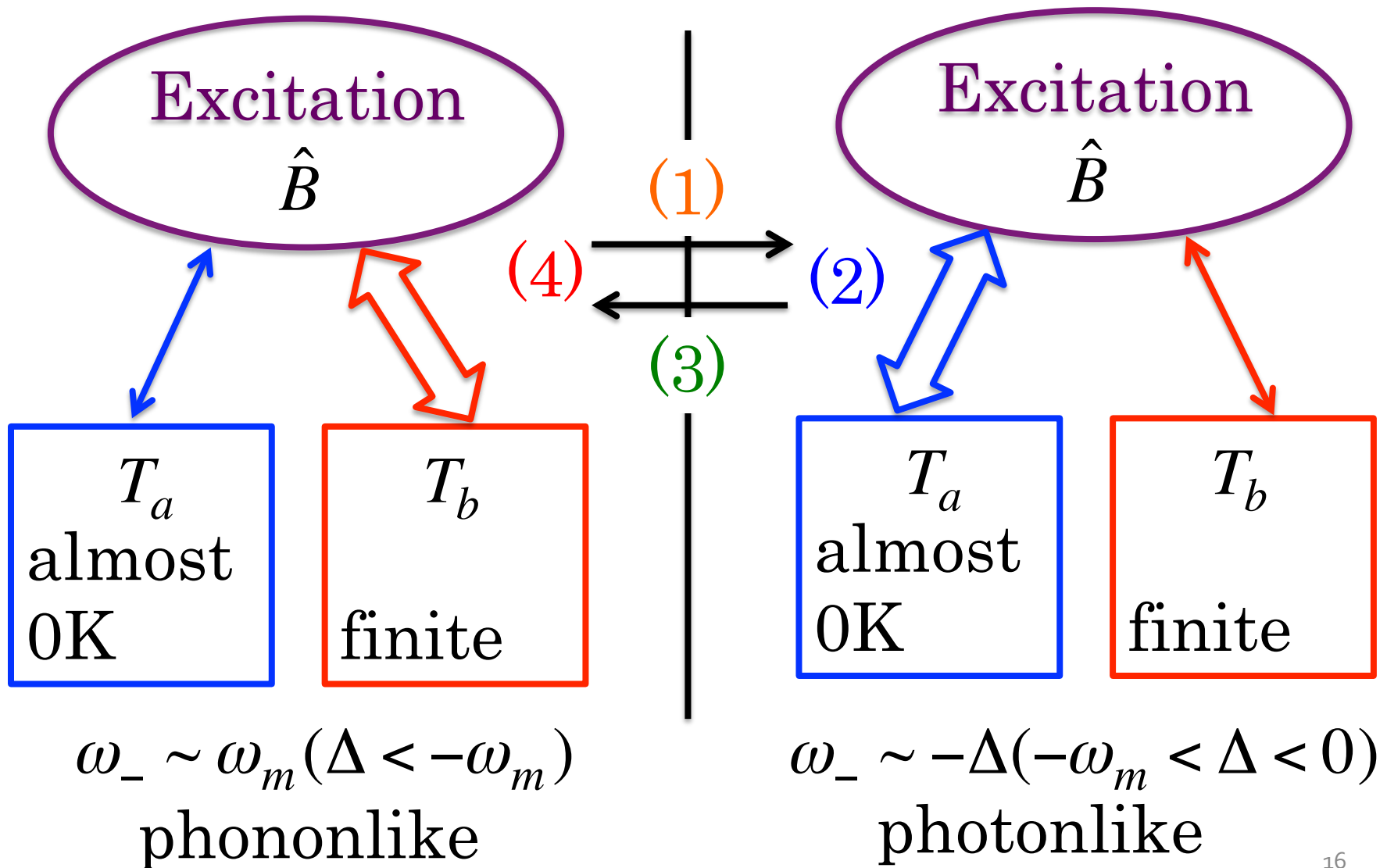
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Otto cycle



(1),(3) changing detuning (laser frequency)

Situation

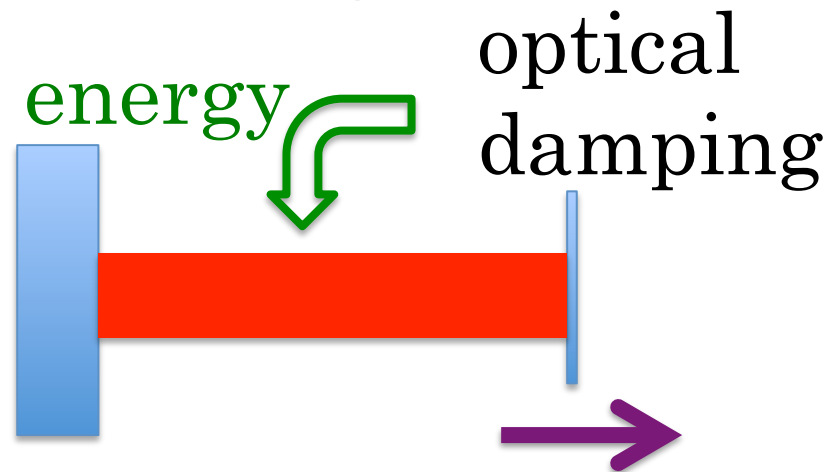


(1) Isentropic expansion

$\omega_i \sim \omega_m (\Delta < -\omega_m)$
phononlike

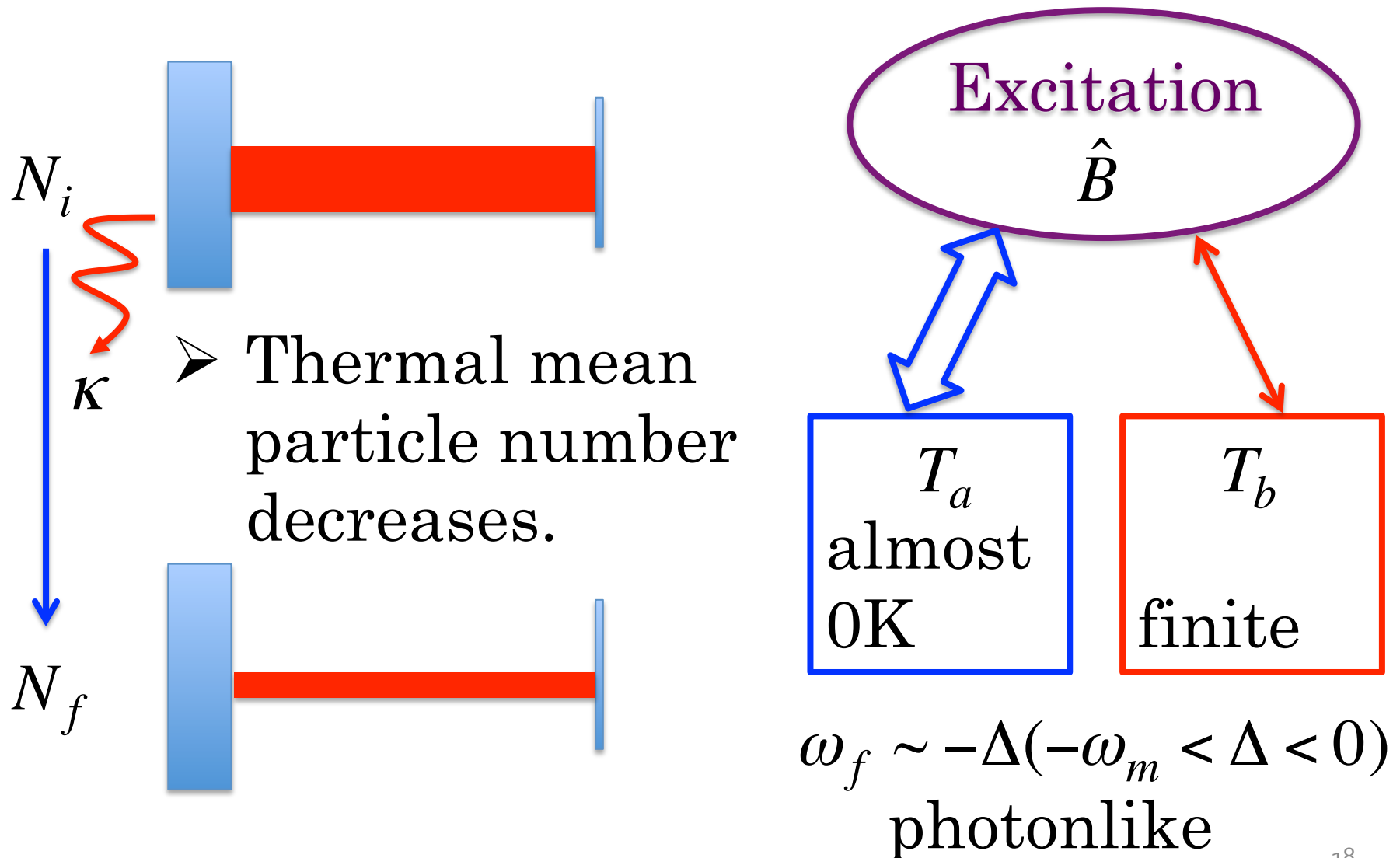
$N_i = \langle \hat{B}^\dagger \hat{B} \rangle_{\omega_i, T_i}$ remains unchanged

$\omega_f \sim -\Delta (-\omega_m < \Delta < 0)$
photonlike



increased radiation pressure = work

(2) Cold Isochoriclike transition



(3) Isentropic compression



$$\omega_f \sim -\Delta (-\omega_m < \Delta < 0)$$

photonlike

$$N_f = \langle \hat{B}^\dagger \hat{B} \rangle_{\omega_f, T_f}$$

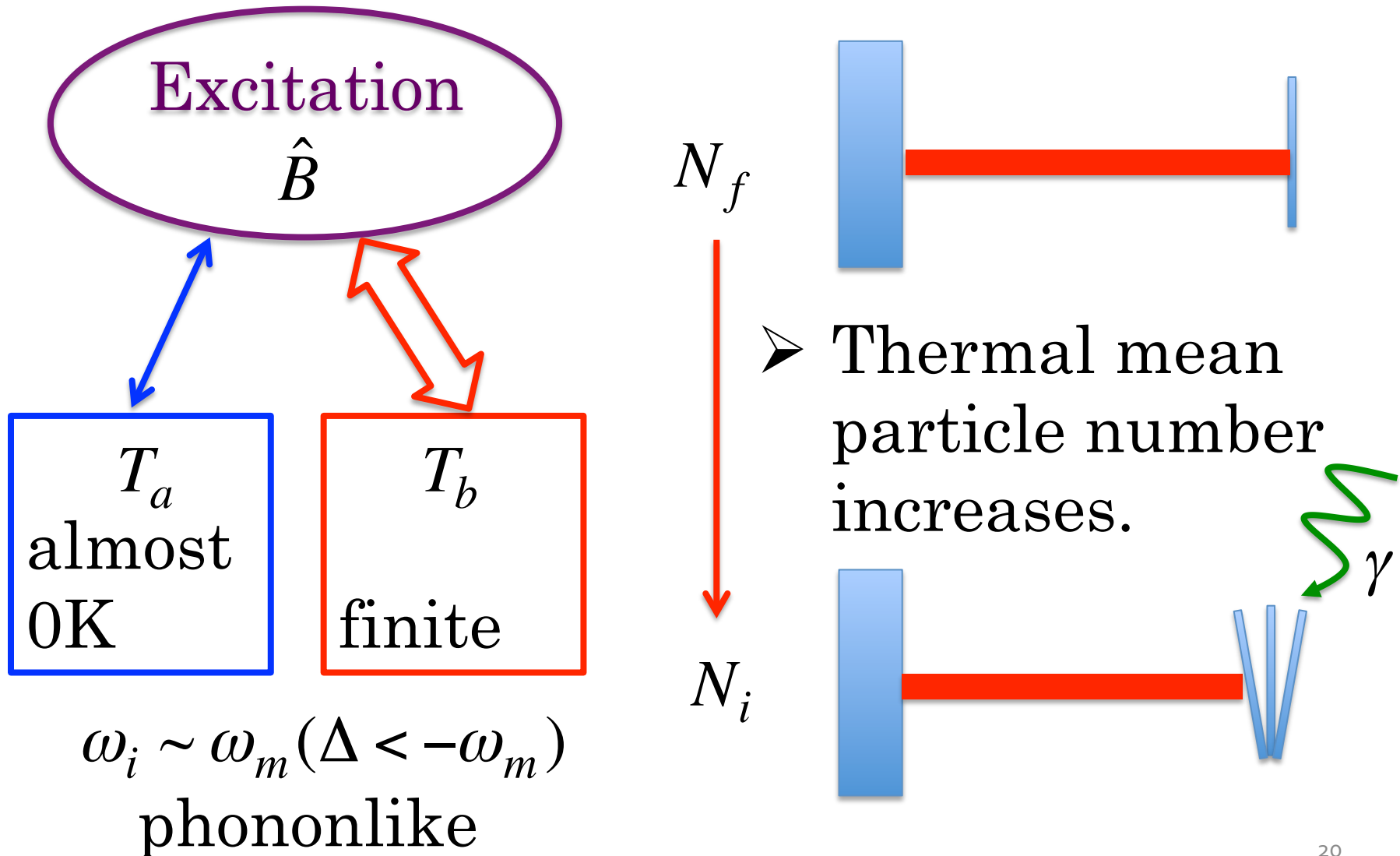
remains unchanged



$$\omega_i \sim \omega_m (\Delta < -\omega_m)$$

phononlike

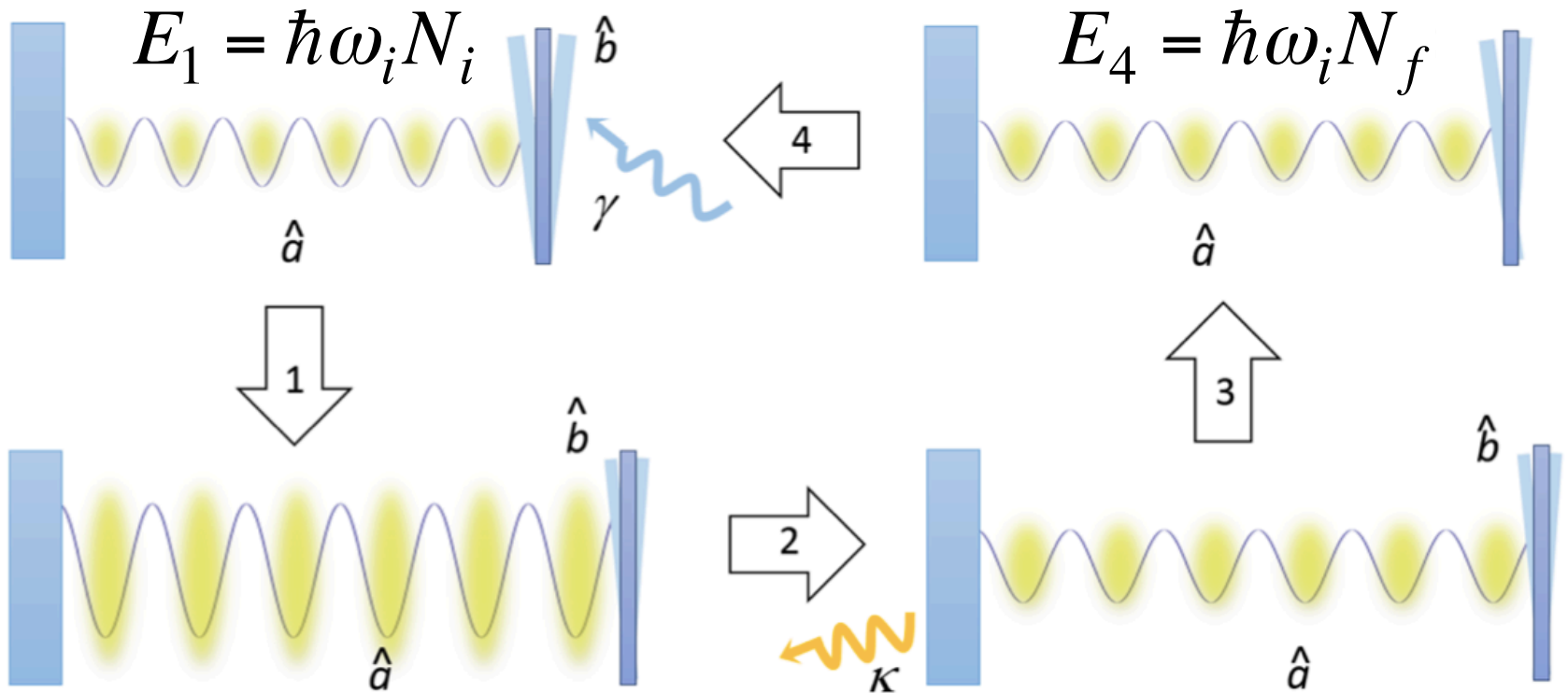
(4) Hot Isochoriclike transition



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Work and efficiency



$$E_2 = \hbar\omega_f N_i$$

➤ total work

$$W = E_1 - E_2 + E_3 - E_4$$

$$E_3 = \hbar\omega_f N_f$$

➤ received heat

$$Q = E_1 - E_4$$

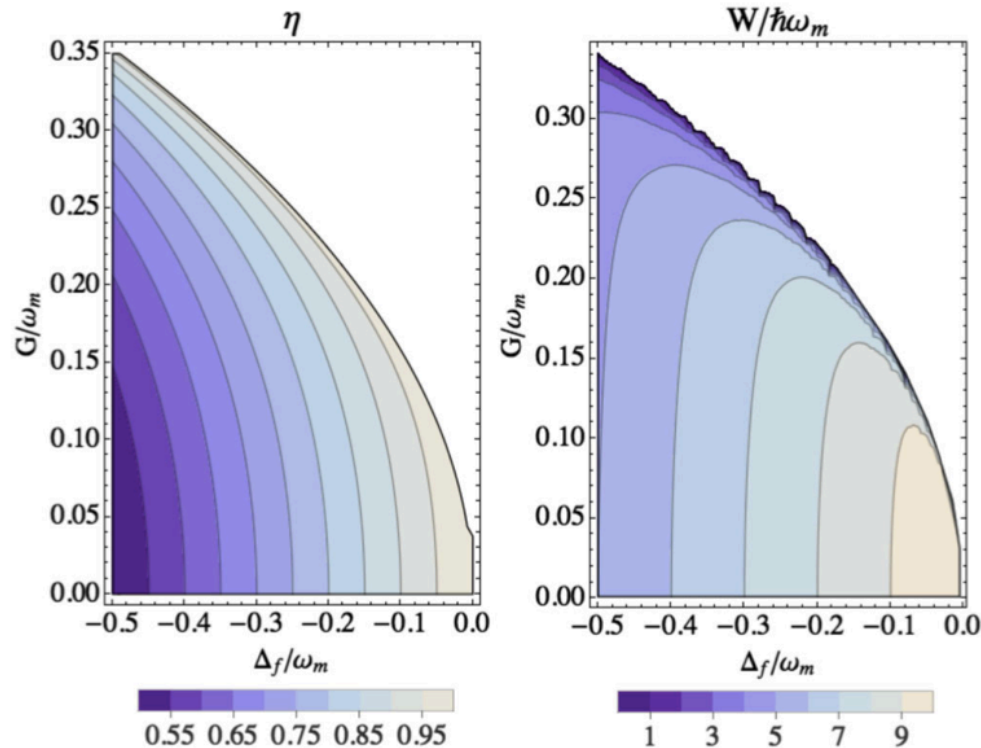
Work and efficiency

➤ efficiency

$$\eta = \frac{W}{Q} = 1 - \frac{\omega_f}{\omega_i}$$

$$W = \hbar(\omega_i - \omega_f)(N_i - N_f)$$

$$Q = \hbar\omega_i(N_i - N_f)$$



$$\Delta_i = -3\omega_m$$

$$T_a = 0, T_b = 0.1$$

Work and efficiency

When $\frac{G}{\omega_m} \ll 1$ (weak coupling)

and $-\frac{\Delta_f}{\omega_m} \ll 1$, the approximate values are

$$\omega_i = \omega_m$$

$$N_i = n_b$$

$$\omega_f = -\Delta_f - \frac{2G^2}{\omega_m}$$

$$N_f = \left(1 + \frac{4\Delta_f G^2}{\omega_m^3}\right) n_a + \frac{2G^2}{\omega_m^2} n_b$$

$$\frac{W}{\hbar\omega_m} = \left(\frac{\Delta_f}{\omega_m} + \frac{2G^2}{\omega_m^2} + 1\right) \left[\left(1 - \frac{2G^2}{\omega_m^2}\right) n_b - \frac{G^2}{\omega_m^2}\right]$$

Work and efficiency

In the high temperature limit of phonon bath,

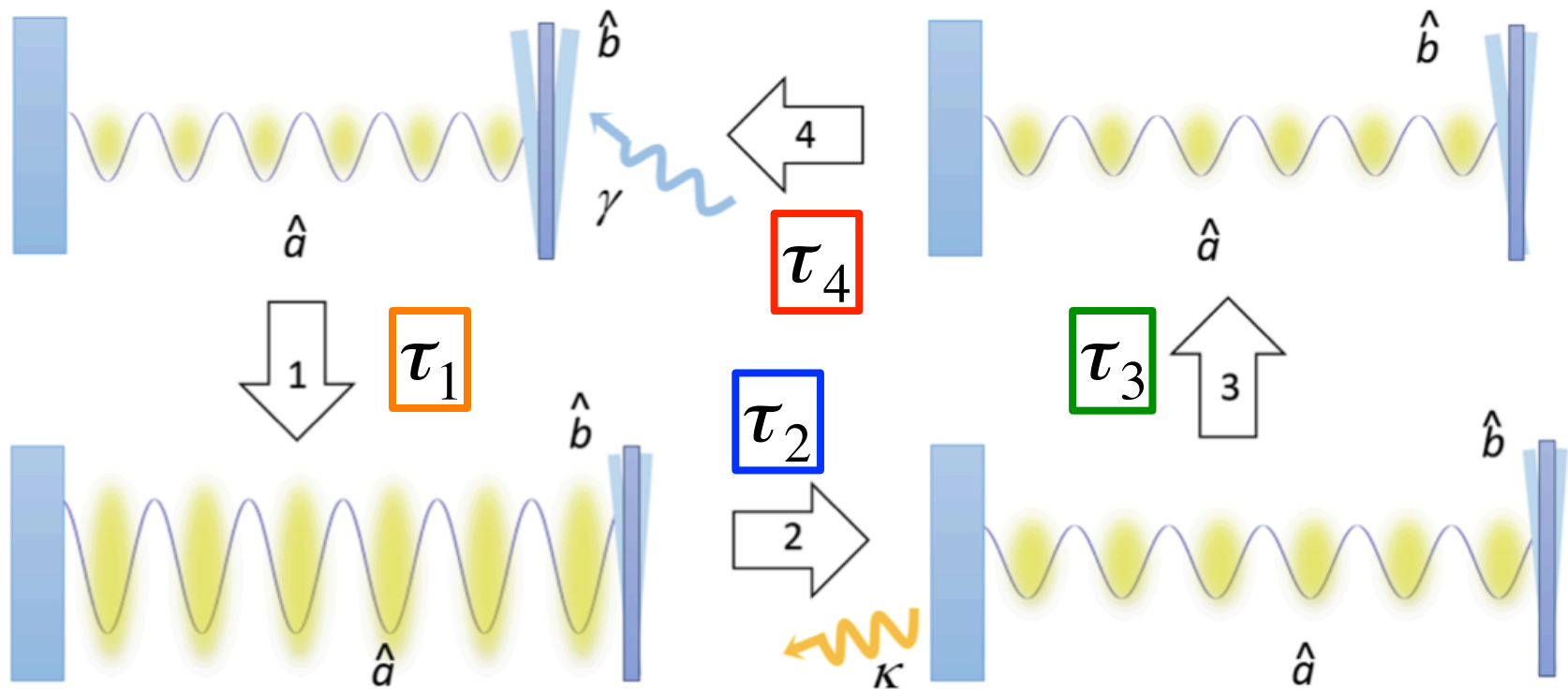
$\frac{k_B T_b}{\hbar \omega_m} \gg 1$, the efficiency at maximum work is

$$\eta_W = 1 - \left(\frac{-\Delta_f}{\omega_m} + \frac{\hbar \omega_m}{4k_B T_b} \right)$$
$$< 1 - \sqrt{\frac{-\hbar \Delta_f}{2k_B T_b}}$$

c.f. Carnot limit

$$\eta < 1 + \frac{\hbar \Delta_f}{2k_B T_b}$$

Timescale

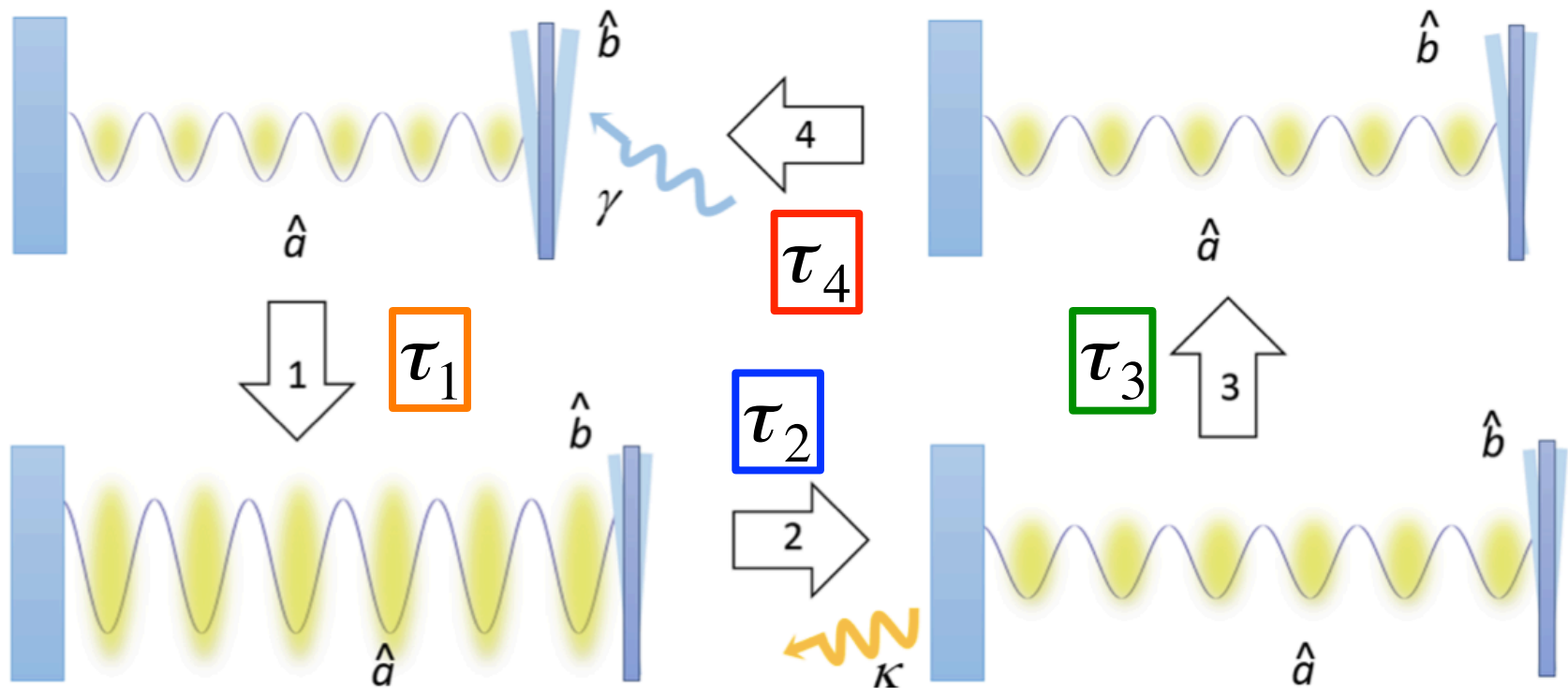


$$1/G \ll \tau_1, \tau_3 < 1/\kappa$$

preventing transition
between \hat{A} and \hat{B}

keeping N_i, N_f
unchanged

Timescale

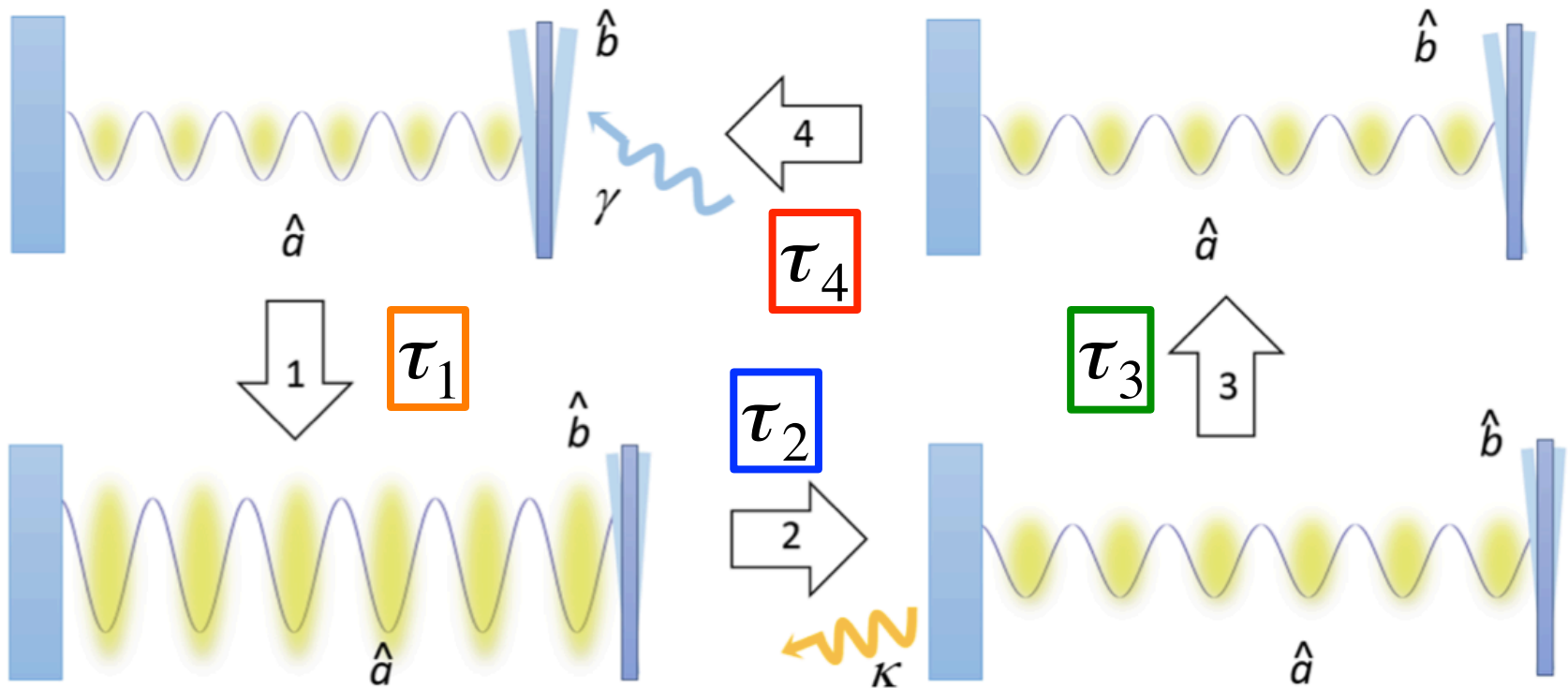


$$1/\kappa < \tau_2 \ll 1/\gamma$$

waiting
thermalization

preventing
thermalization of \hat{A}

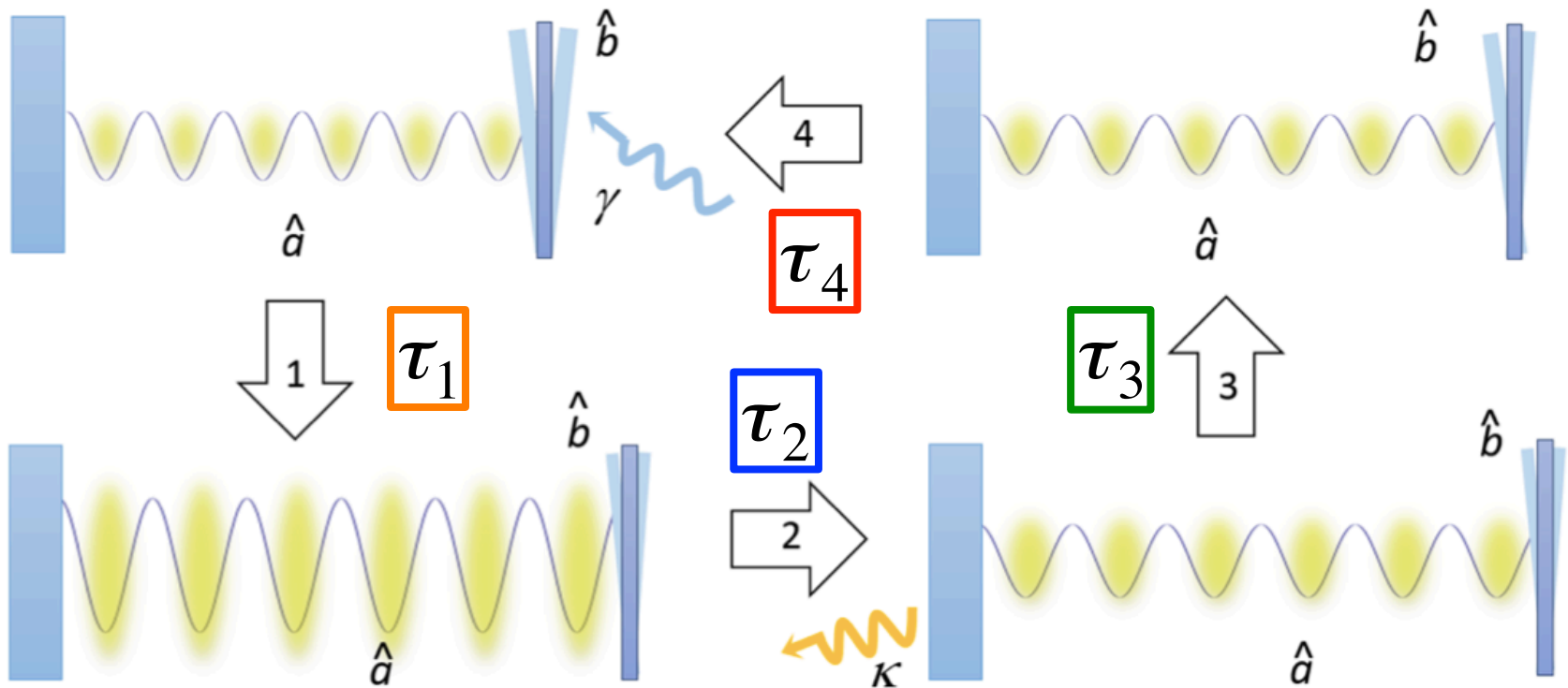
Timescale



$$1/\gamma < \tau_4$$

waiting
thermalization

Timescale



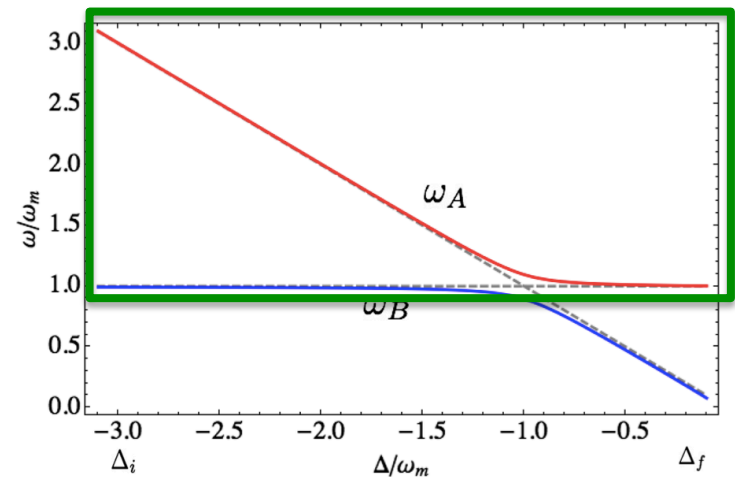
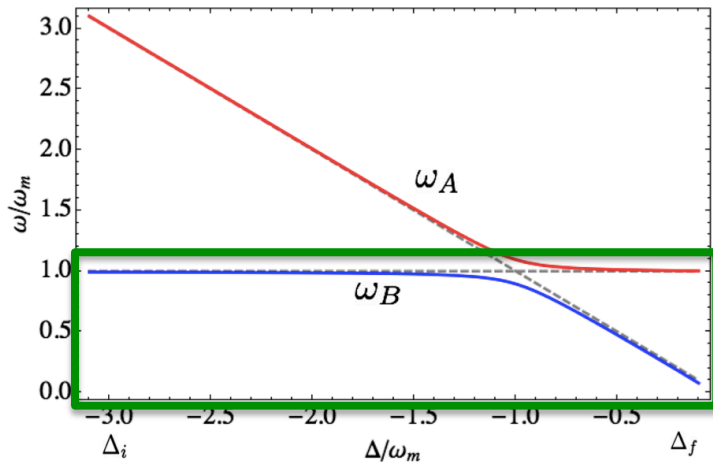
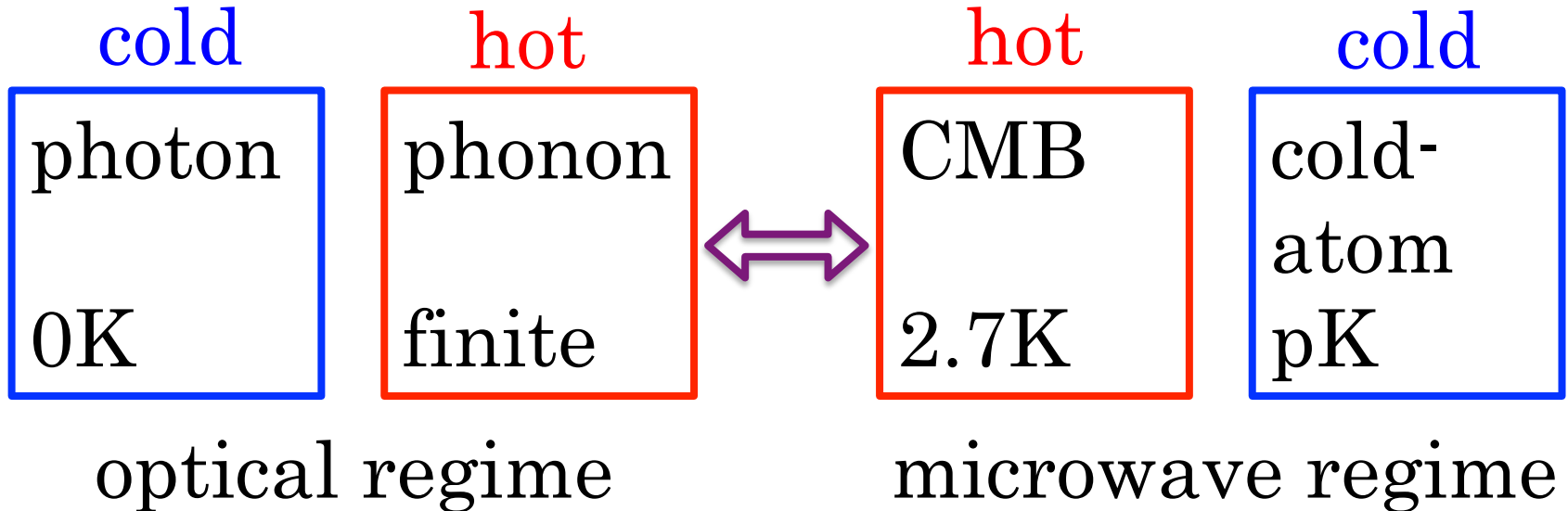
requirement

$$1 / \tau_4 < \gamma \ll 1 / \tau_2 < \kappa < 1 / \tau_{1,3} \ll G \ll \omega_m$$

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Cosmic blackbody radiation



Cosmic blackbody radiation

- Can extract energy from CMB !

requirement

thermal motion < coherent momentum recoil $2\hbar k$

In a case of lithium atoms, not to exceed a pK for 300 GHz photons

hot

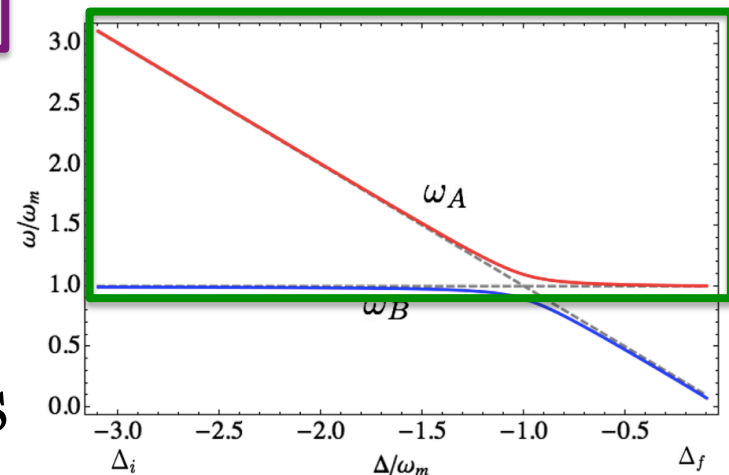
CMB

2.7K

cold

cold-atom
pK

microwave regime



Future work

- Evaluating the role of imperfections due to the coupling to the thermal reservoirs.
- Considering the effects of non-adiabatic transitions between polaritons and non-ideal control of maintaining the intra-cavity power.
- Method to extract work from the cycle actually

Summary

- Heat engine can be realized with opto-mechanical system.
- It is possible to produce an Otto cycle by varying the cavity detuning in terms of two effective reservoirs, which consist phonon and photon.
- In principle, this system can extract work from cosmic microwave background.